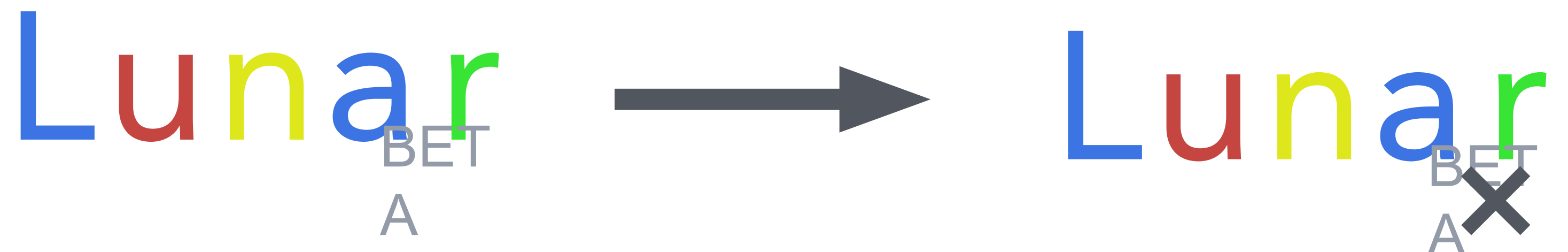


# **AERONET Nighttime AOD Product**

**A summary of method and the end of ‘provisional’ status**

**AERONET has been acquiring lunar observations from the majority of model T cimels for many years and producing a night-time AOD data set which currently includes observations at 492 sites dating back as far as 2014 for some locations.**

- 1) How has lunar evolved from original provisional version?
- 2) Summary of lunar AOD product evaluation?



# Updated Empirical Bias Correction

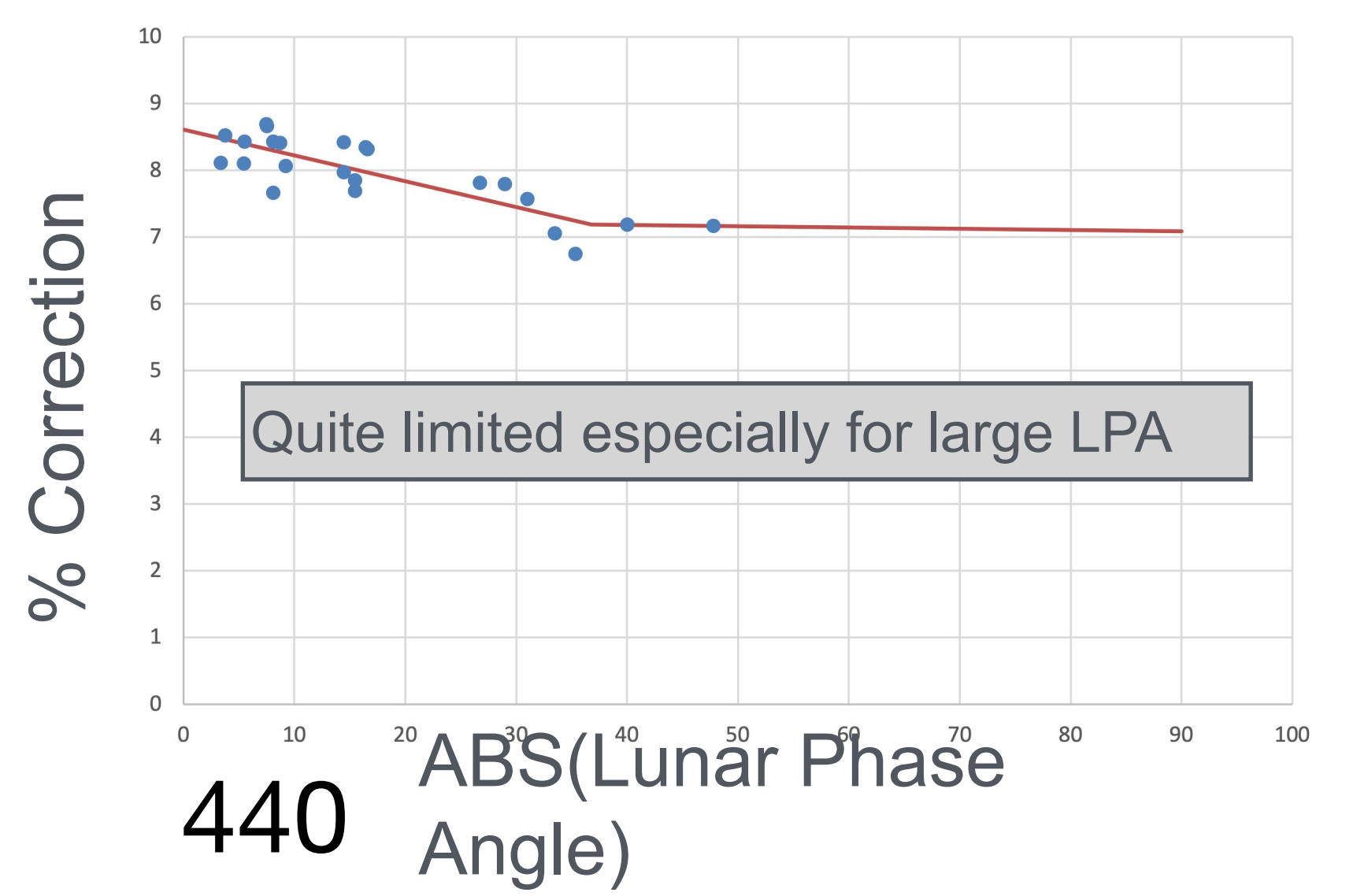
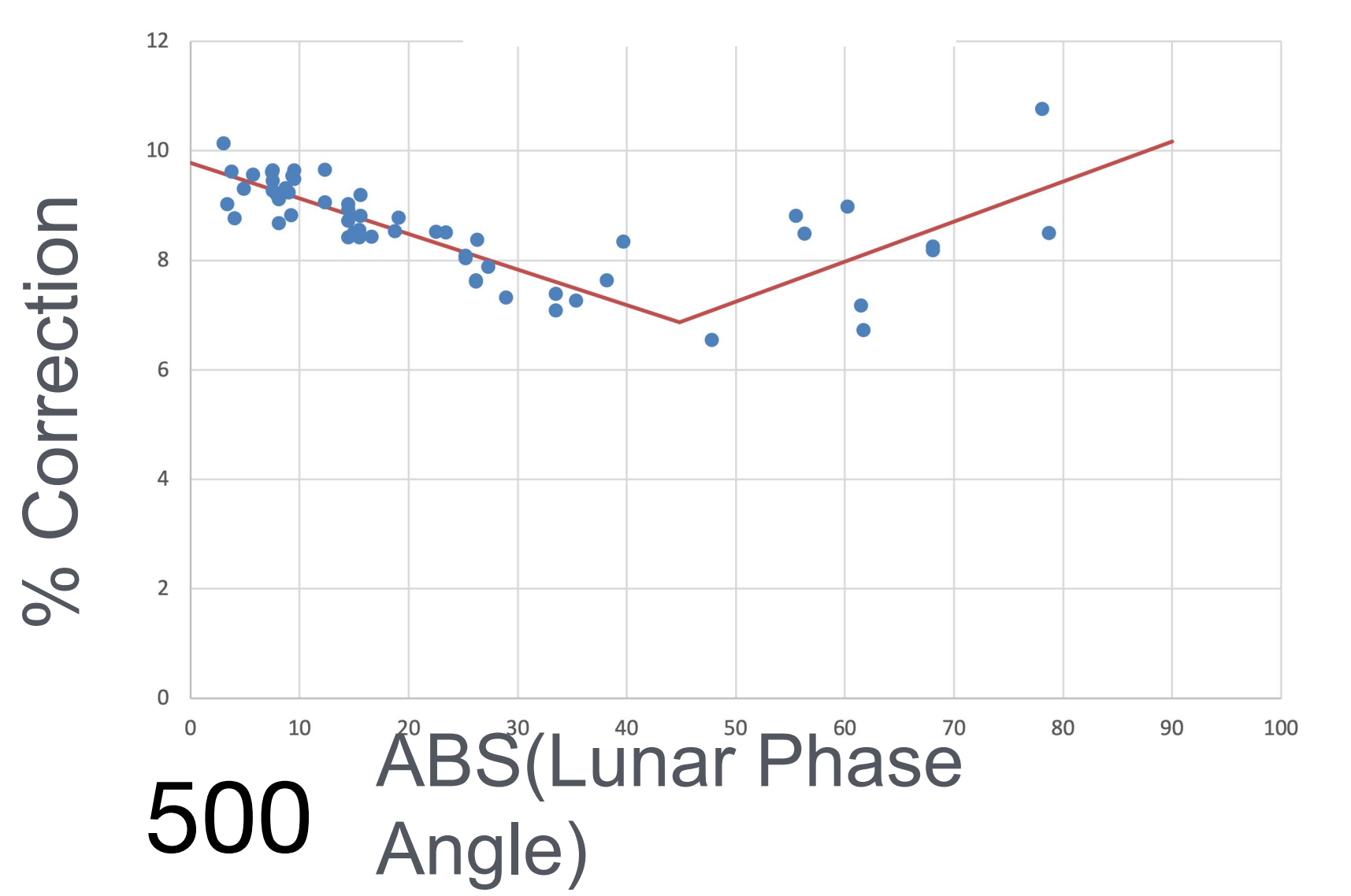
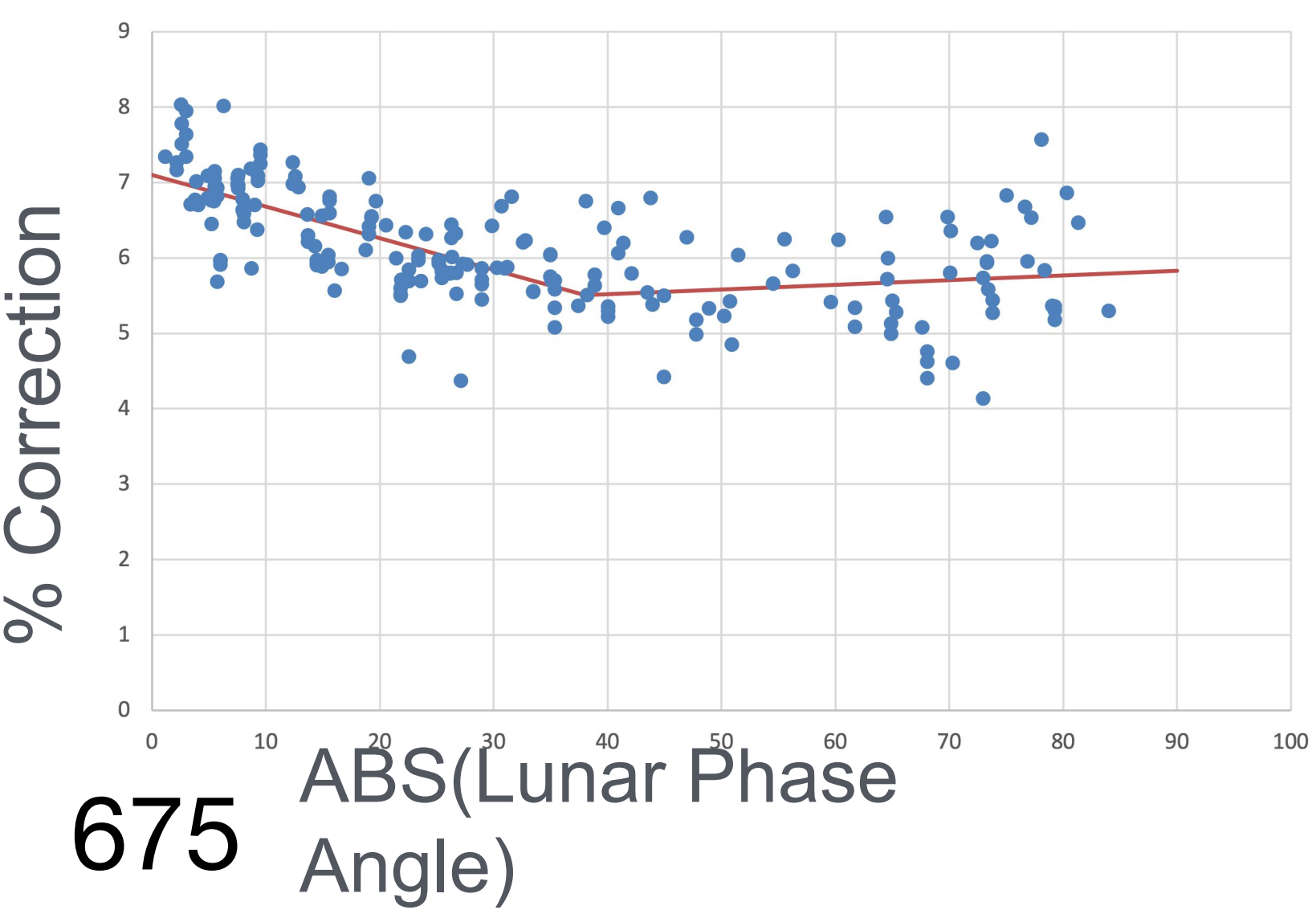
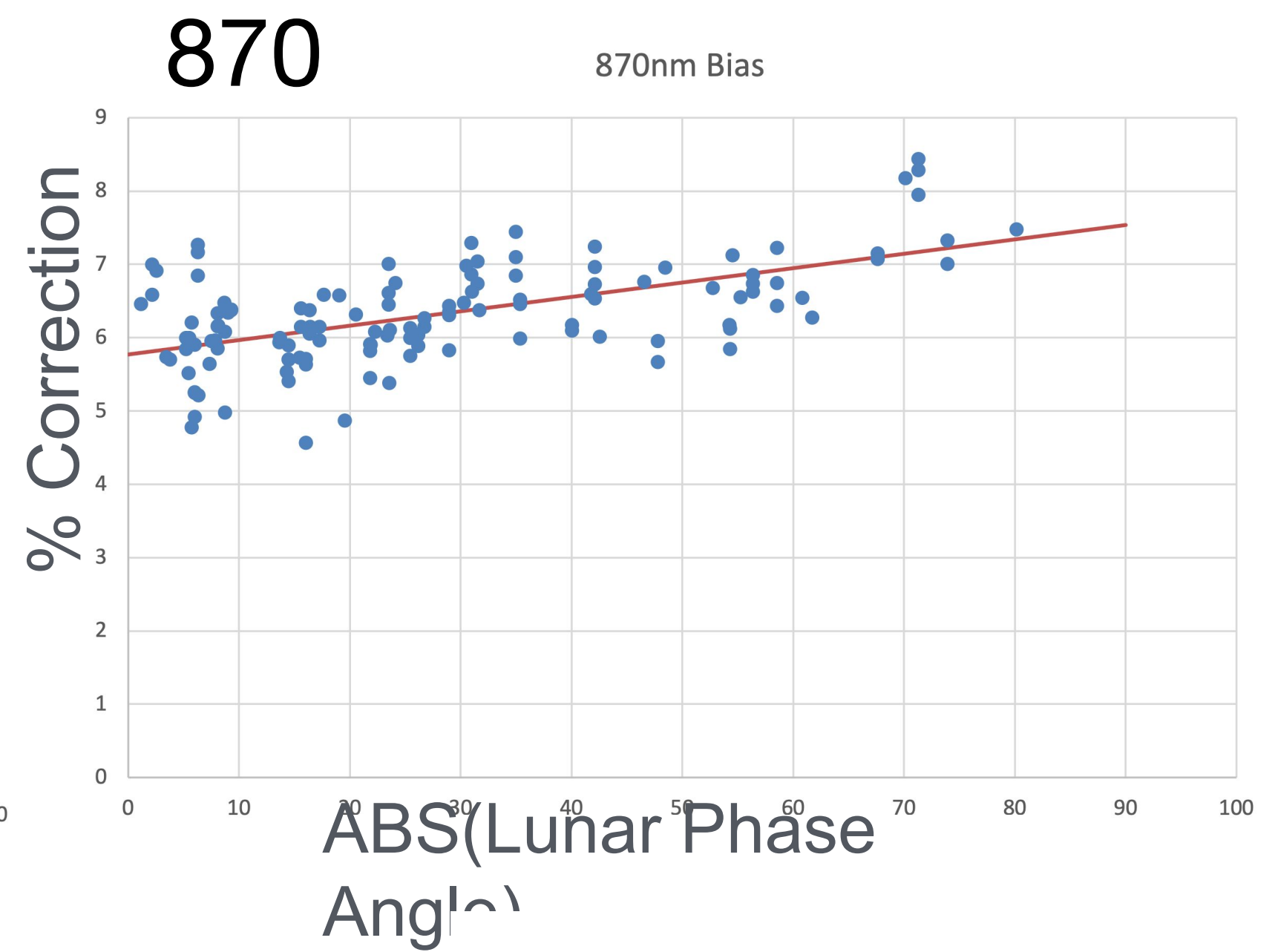
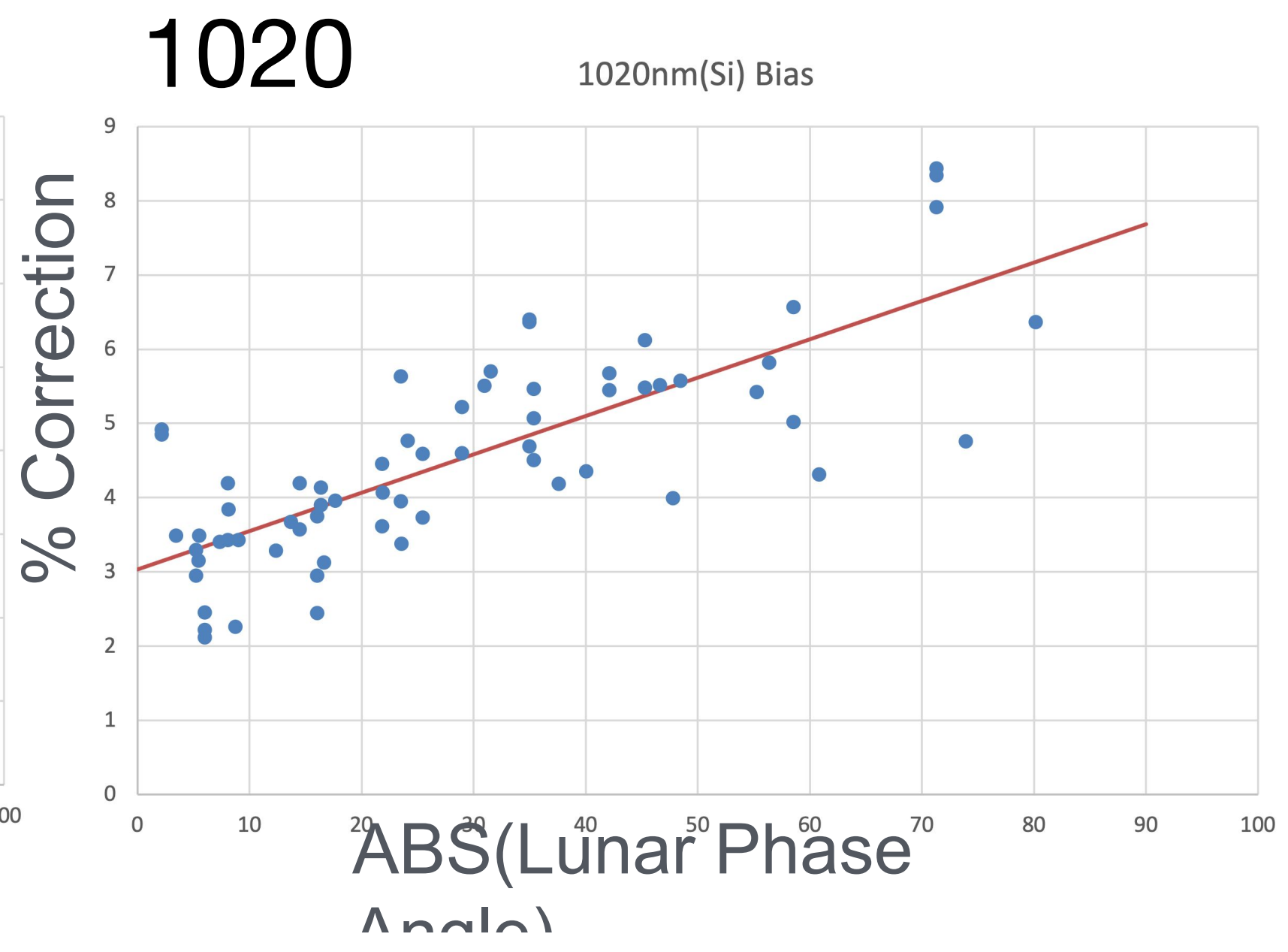
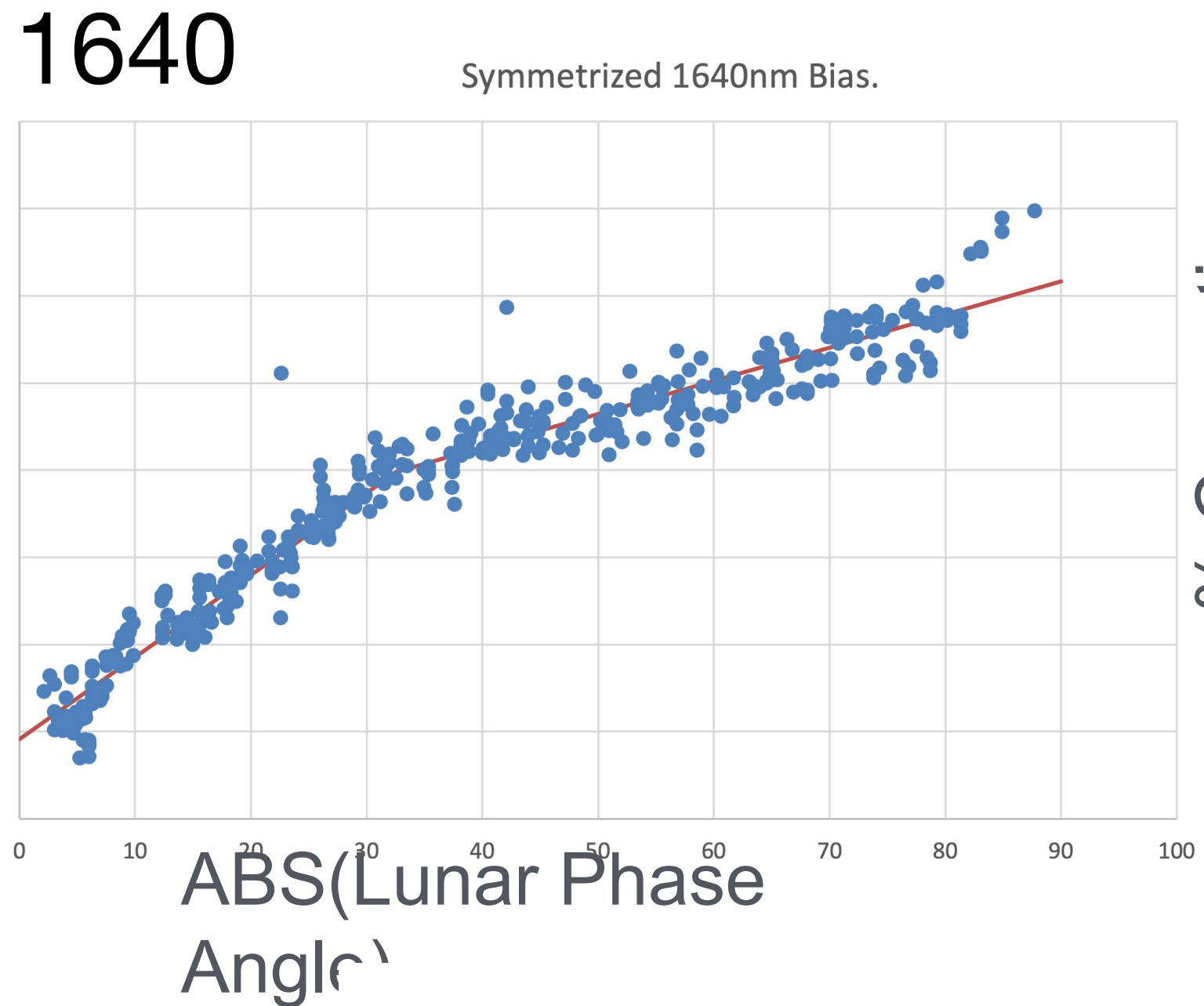
**Based on lunar Langleys from MLO and Izaña from 2015-2024**

The AERONET approach to address errors in ROLO irradiance has been via comparison of  $V_o$ 's of high elevation lunar Langleys with solar Langleys at Mauna Loa Observatory (MLO) and Izaña calibration facilities.

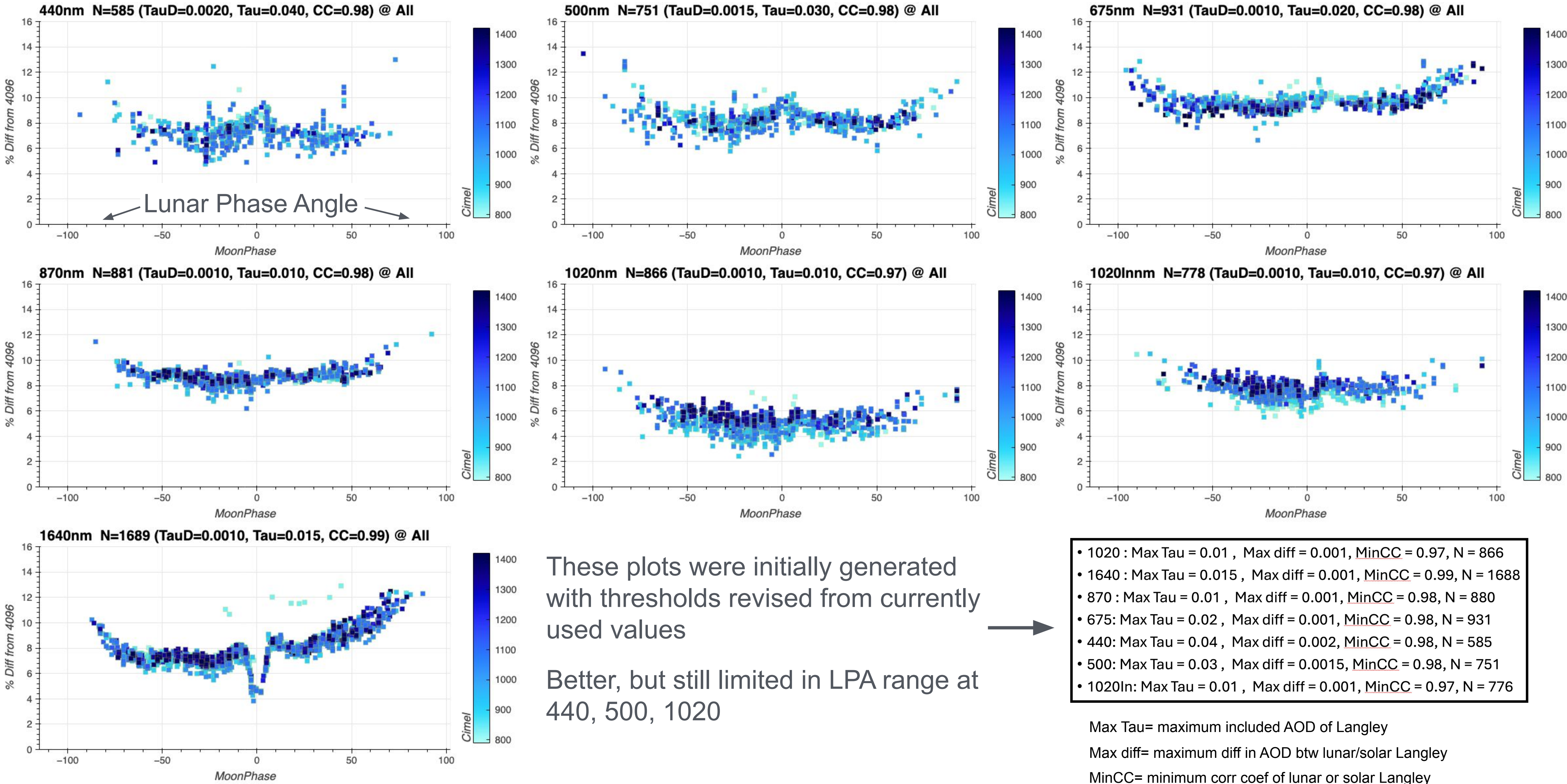
For a given instrument, solar and lunar measurements are made with the same optical/electronic path so in principal the ratios of solar to lunar  $V_o$  should match the sun/sky gain ratio of 4096. In practice, this is not an option due to errors in the lunar irradiance model.

Thus, observed deviations from this nominal value could be determined as a function of lunar phase angle (LPA) and used as an empirical correction factor by which the base ROLO irradiances are modified.

**Original analyses based on dozens of solar/lunar Langleys from MLO and Izaña from 20 cimels over 5 years**



# MLO + Izaña Observed Deviation from Nominal sun/sky gain= 4096



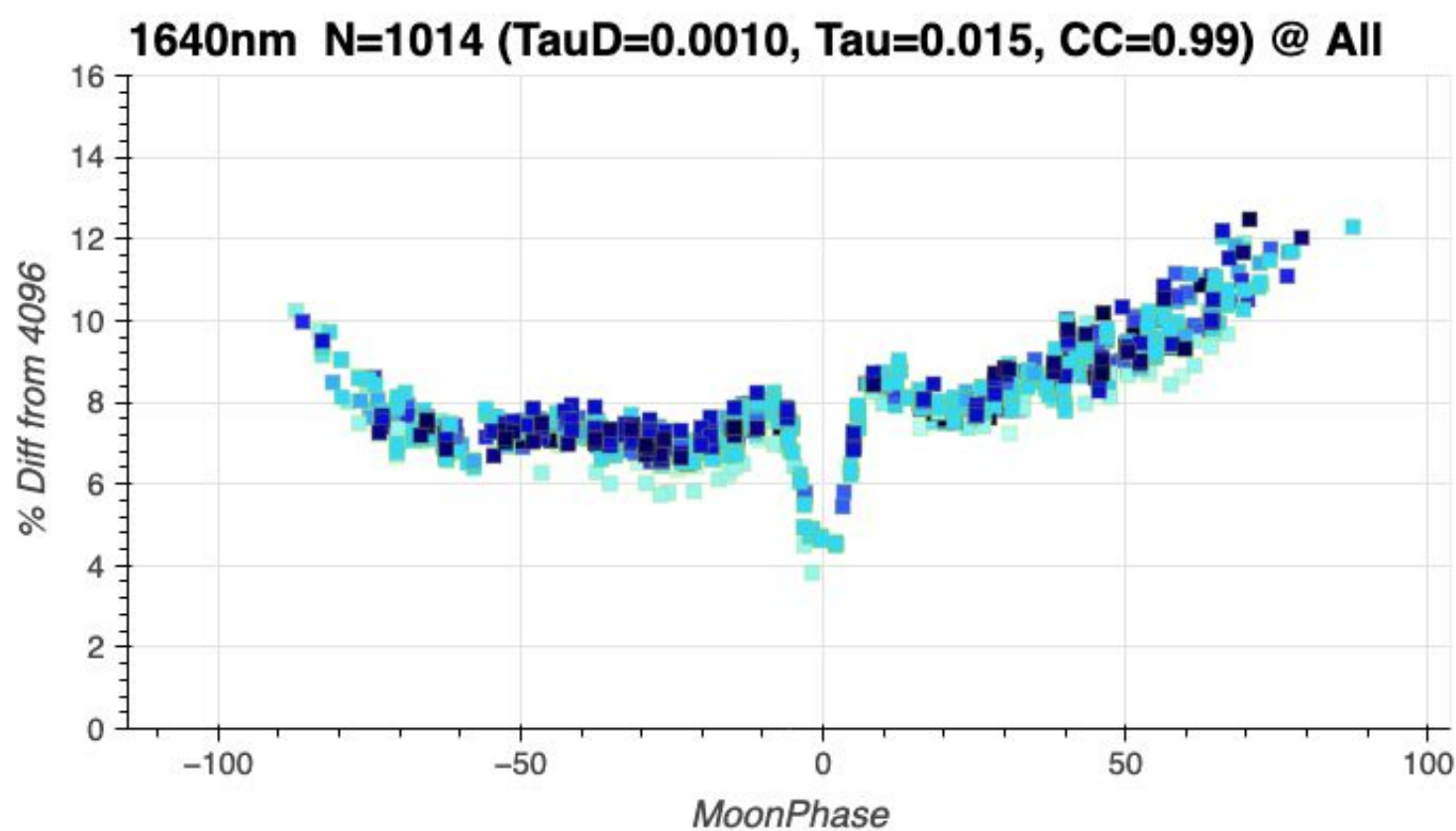
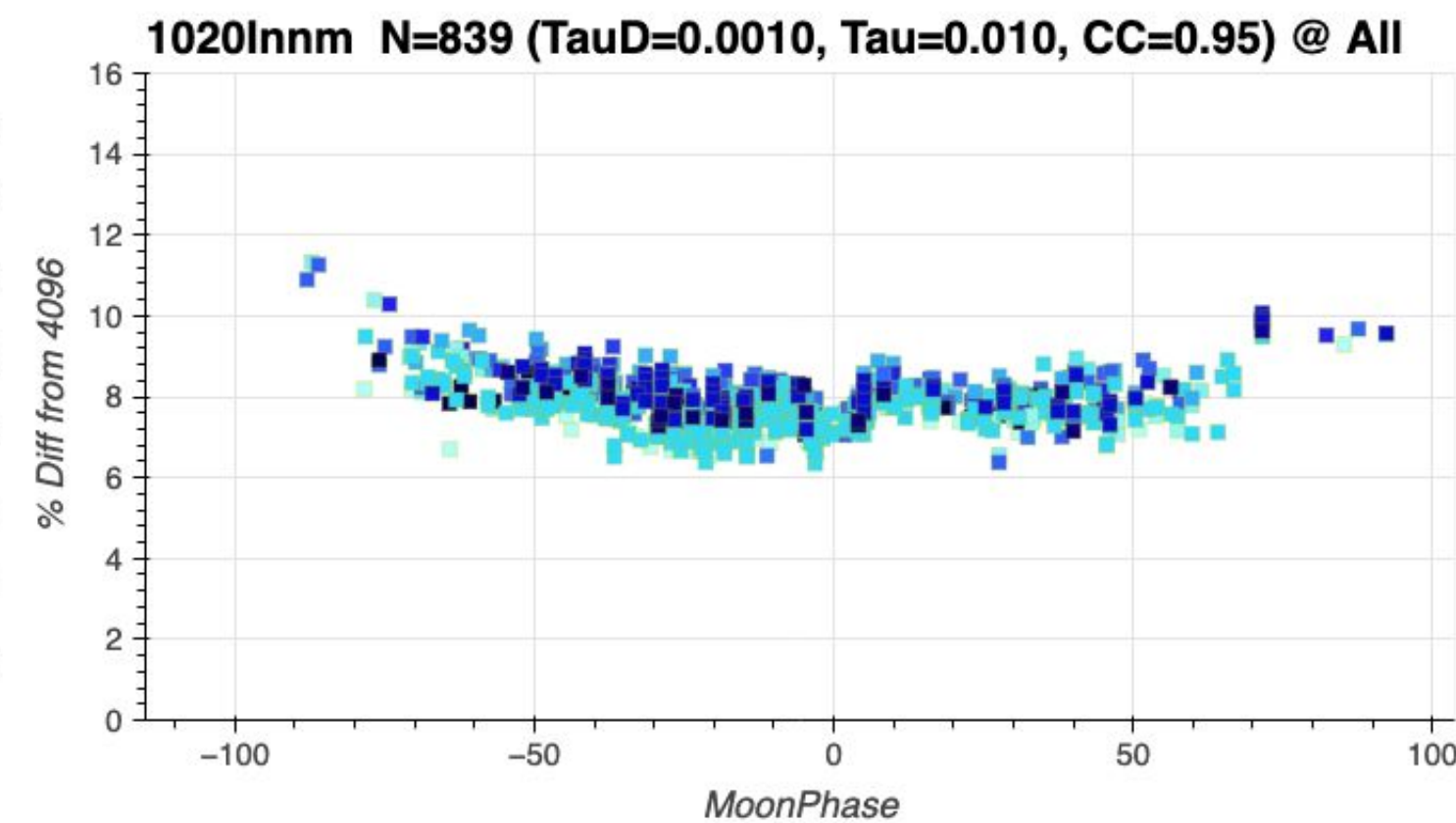
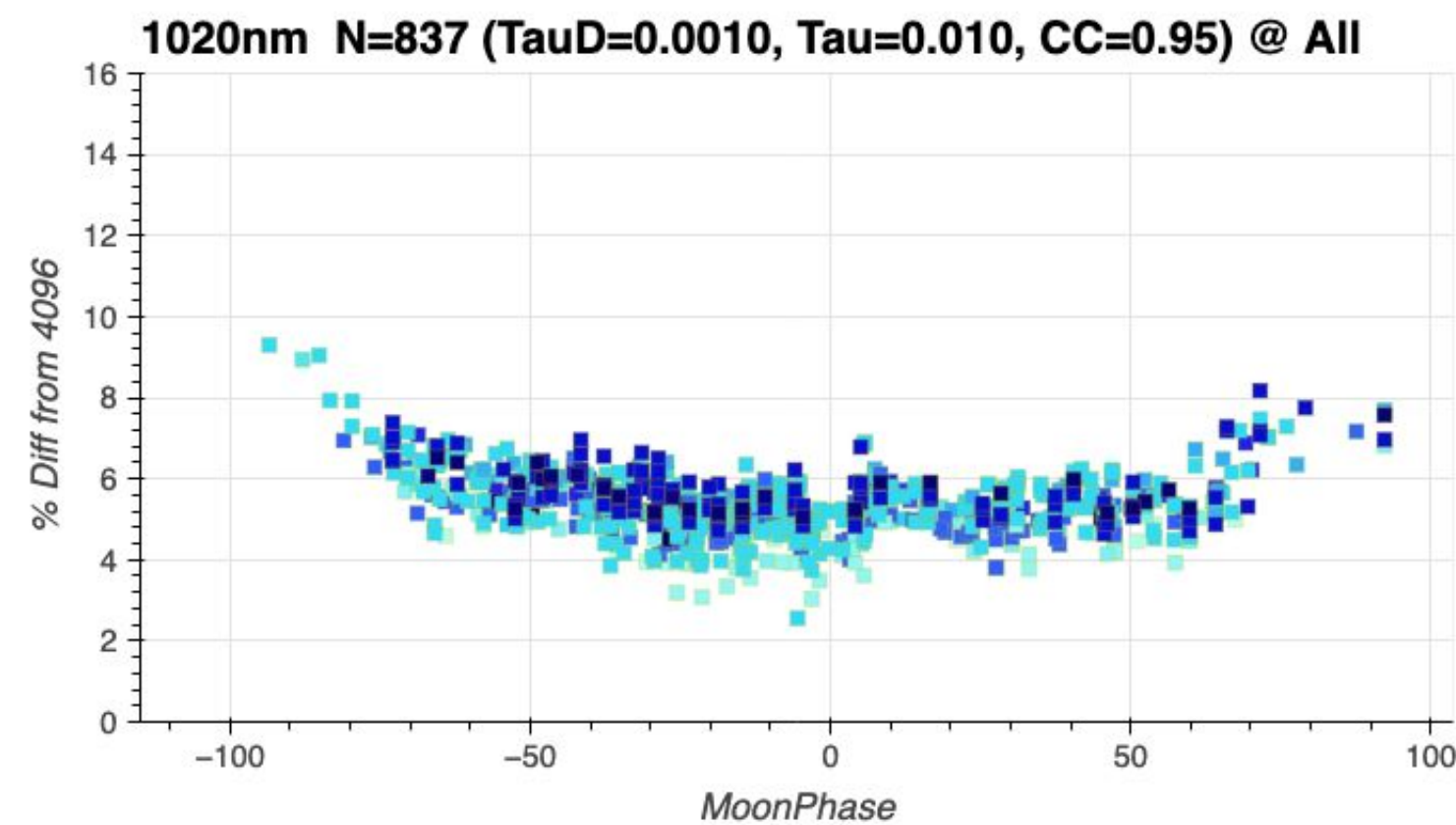
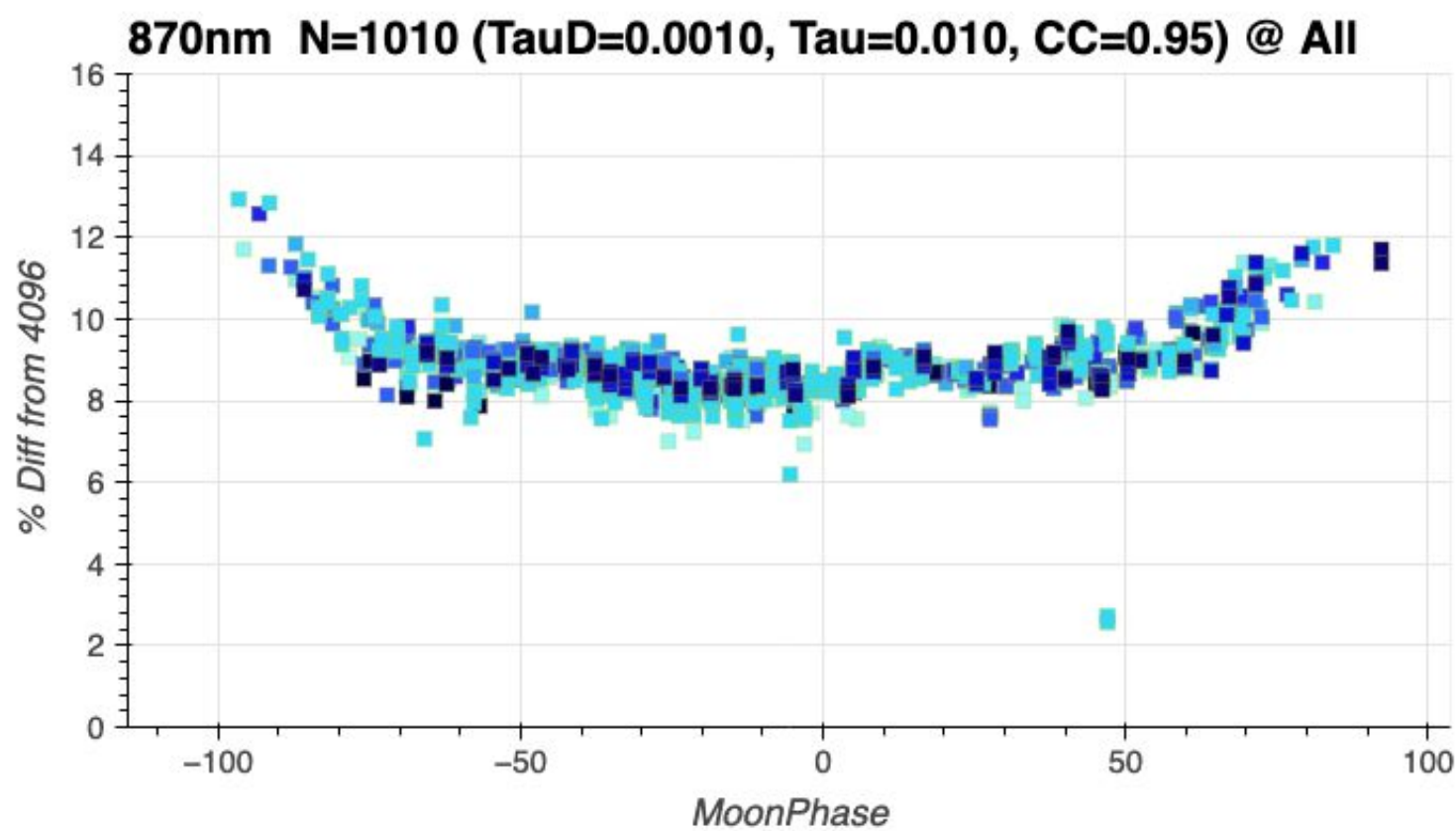
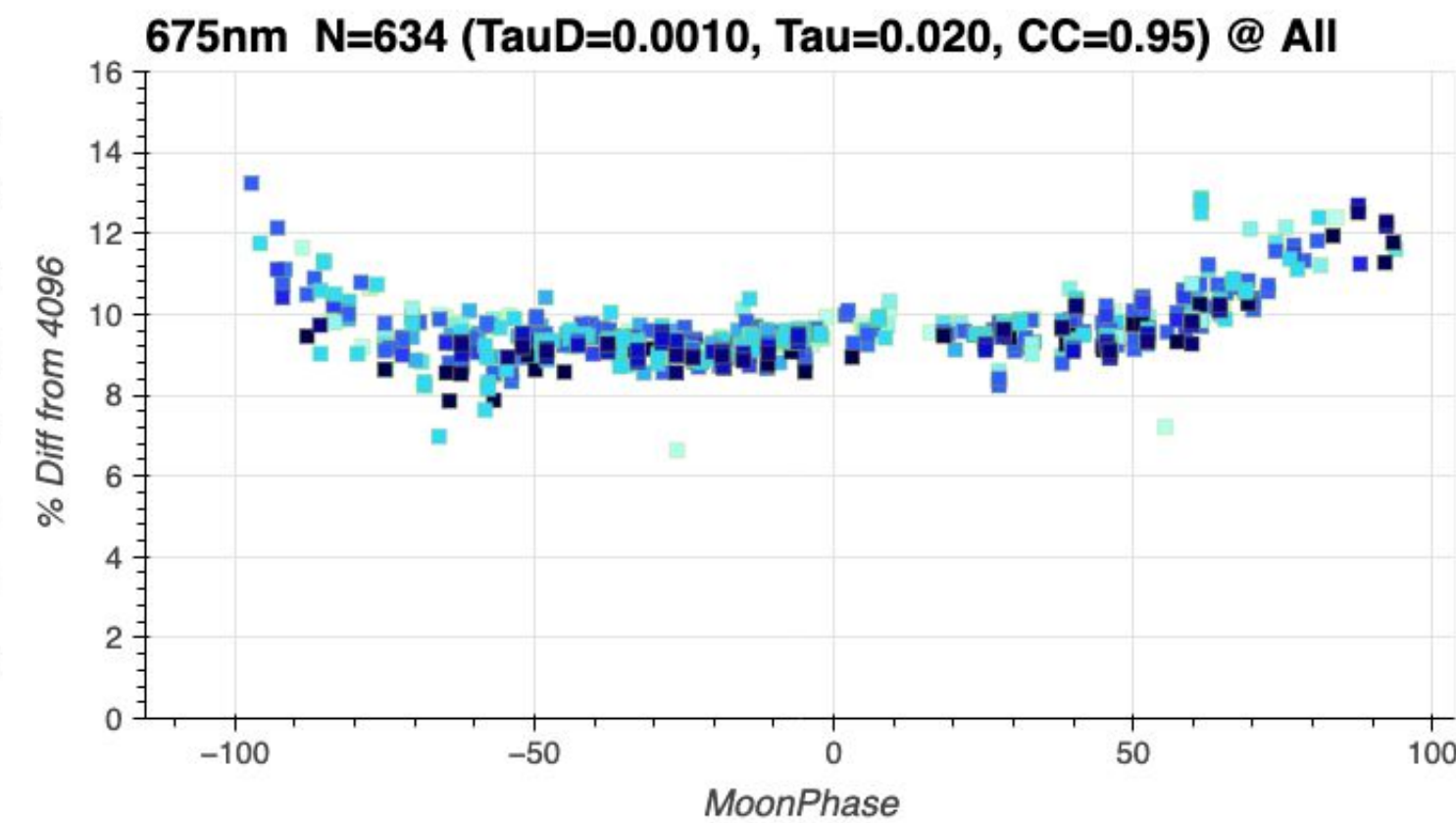
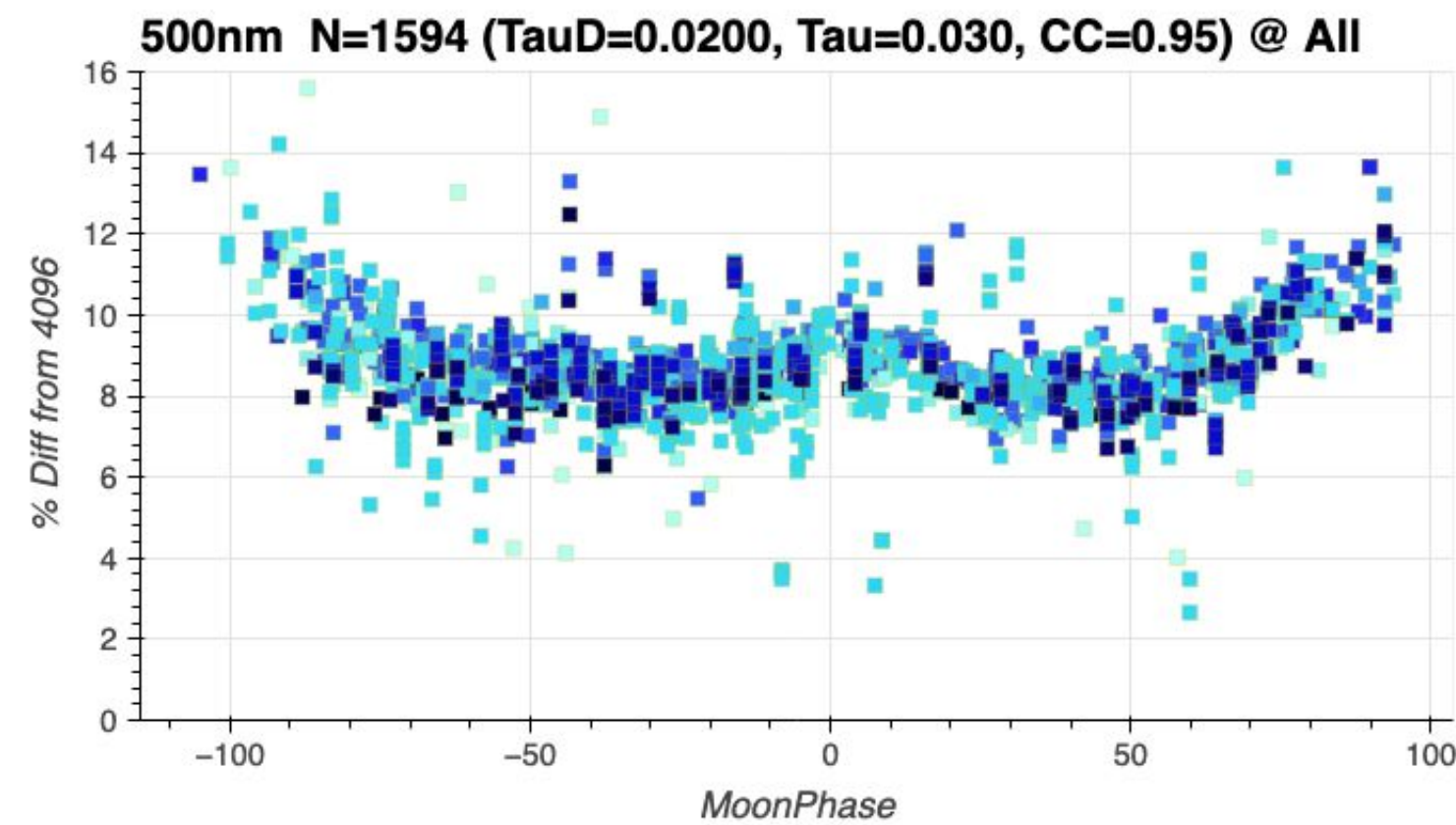
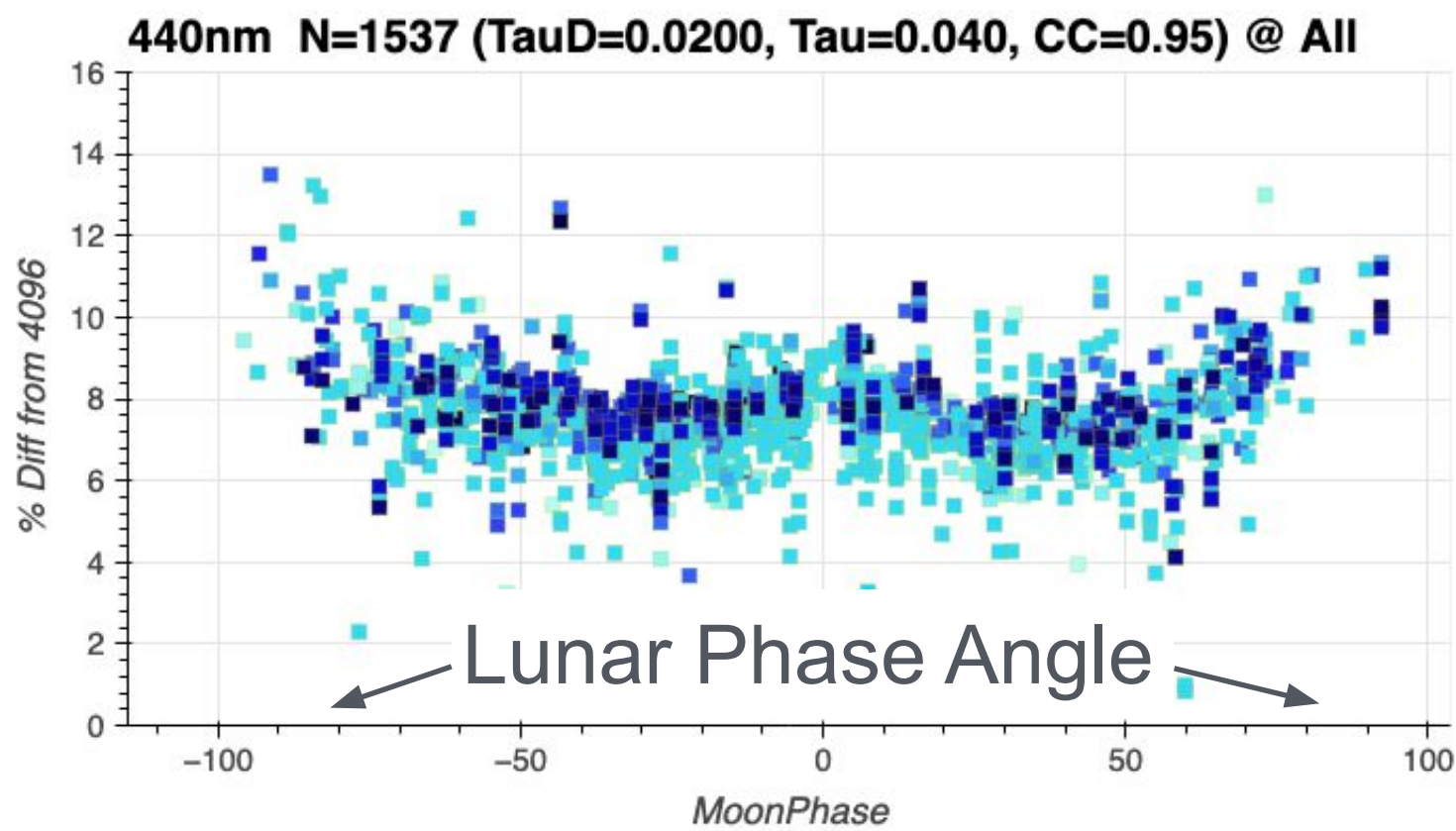
These plots were initially generated with thresholds revised from currently used values

Better, but still limited in LPA range at 440, 500, 1020

- 1020 : Max Tau = 0.01 , Max diff = 0.001, MinCC = 0.97, N = 866
- 1640 : Max Tau = 0.015 , Max diff = 0.001, MinCC = 0.99, N = 1688
- 870 : Max Tau = 0.01 , Max diff = 0.001, MinCC = 0.98, N = 880
- 675: Max Tau = 0.02 , Max diff = 0.001, MinCC = 0.98, N = 931
- 440: Max Tau = 0.04 , Max diff = 0.002, MinCC = 0.98, N = 585
- 500: Max Tau = 0.03 , Max diff = 0.0015, MinCC = 0.98, N = 751
- 1020ln: Max Tau = 0.01 , Max diff = 0.001, MinCC = 0.97, N = 776

Max Tau= maximum included AOD of Langley  
 Max diff= maximum diff in AOD btw lunar/solar Langley  
 MinCC= minimum corr coef of lunar or solar Langley

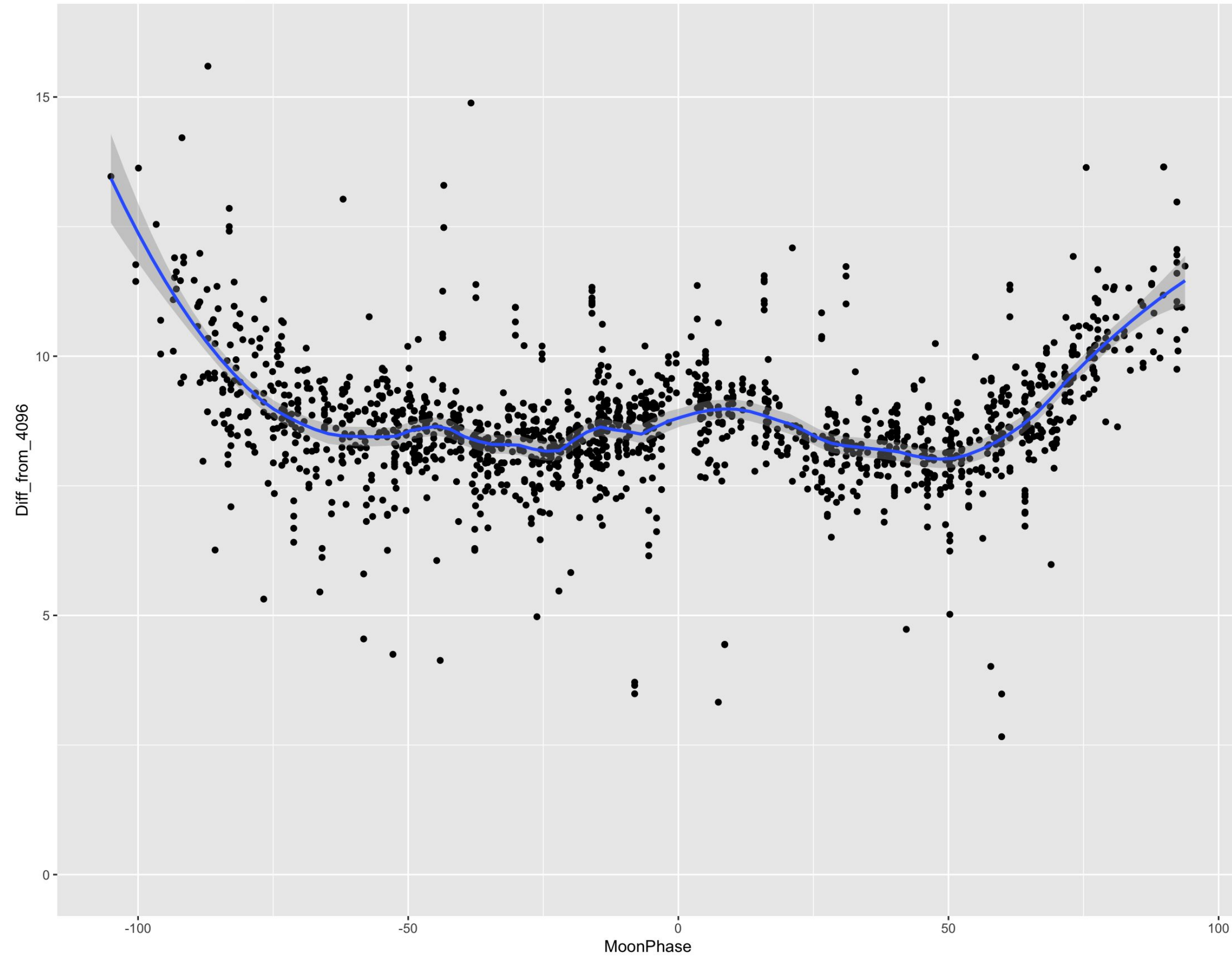
# Updated empirical bias data with additional 5 years of Langley observations



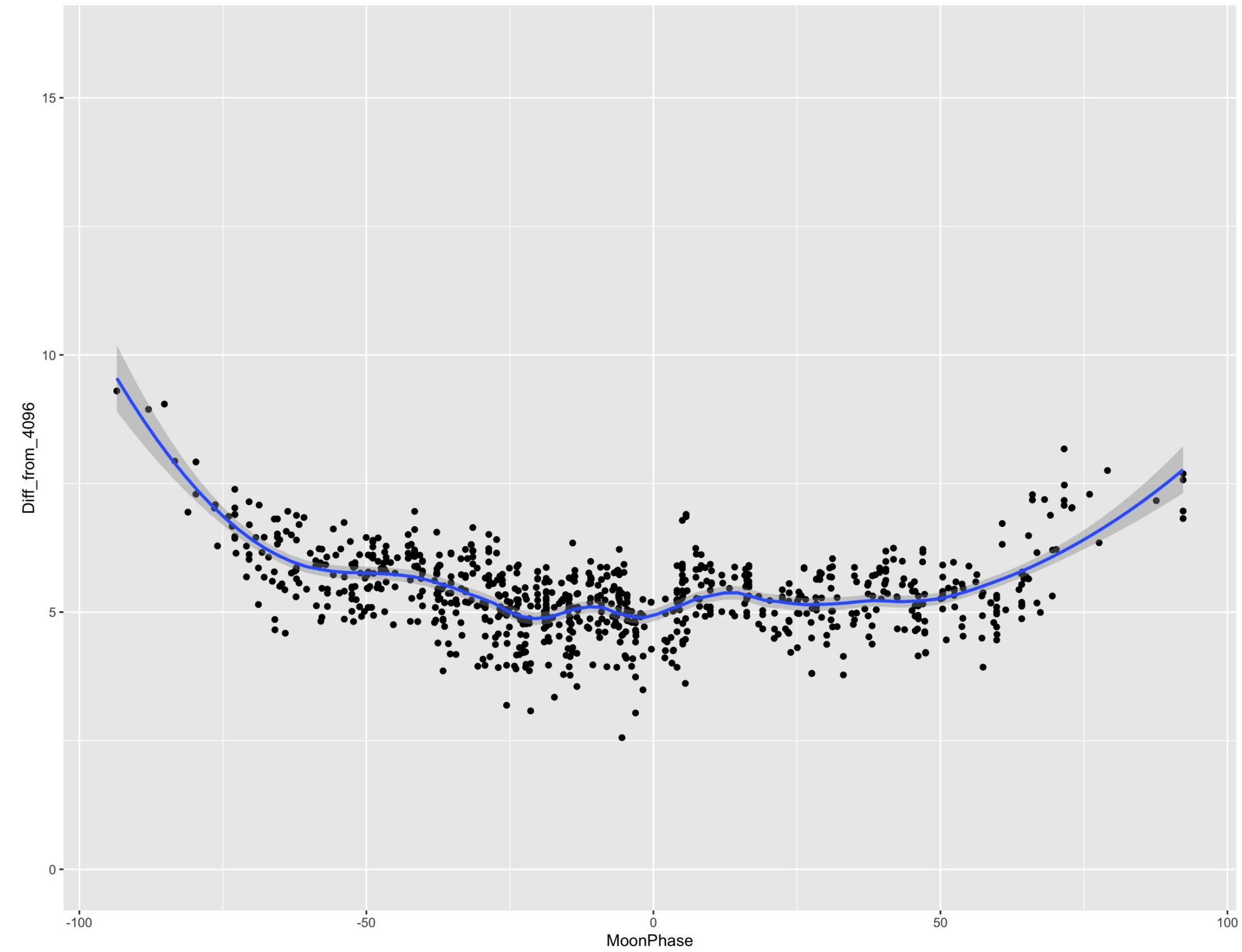
Original comparison thresholds modified to increase data set  
Sufficient data density for robust trends for full MPA range

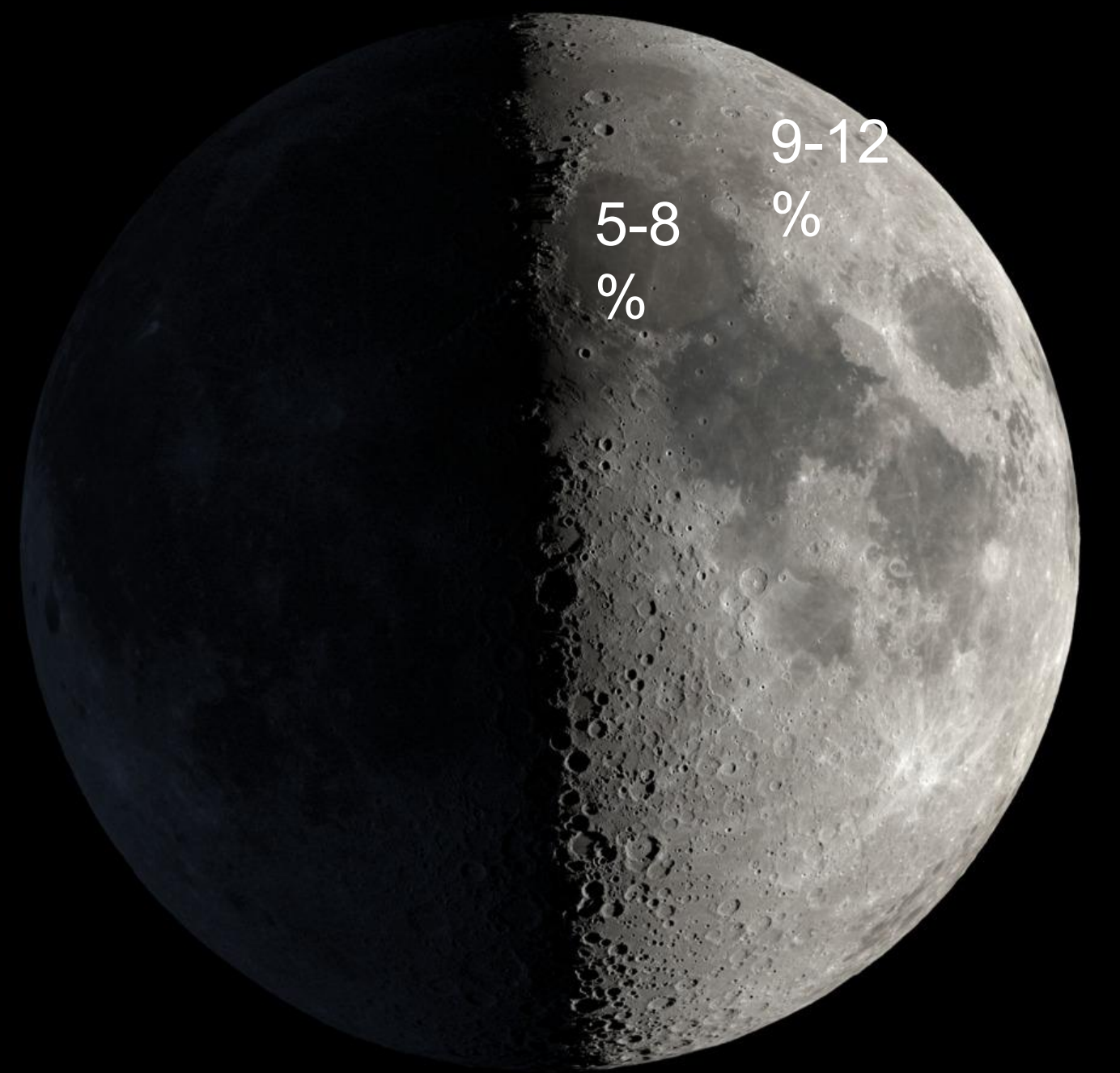
# MLO + Izaña Combined Langley Dataset (LOESS fit)

500nm



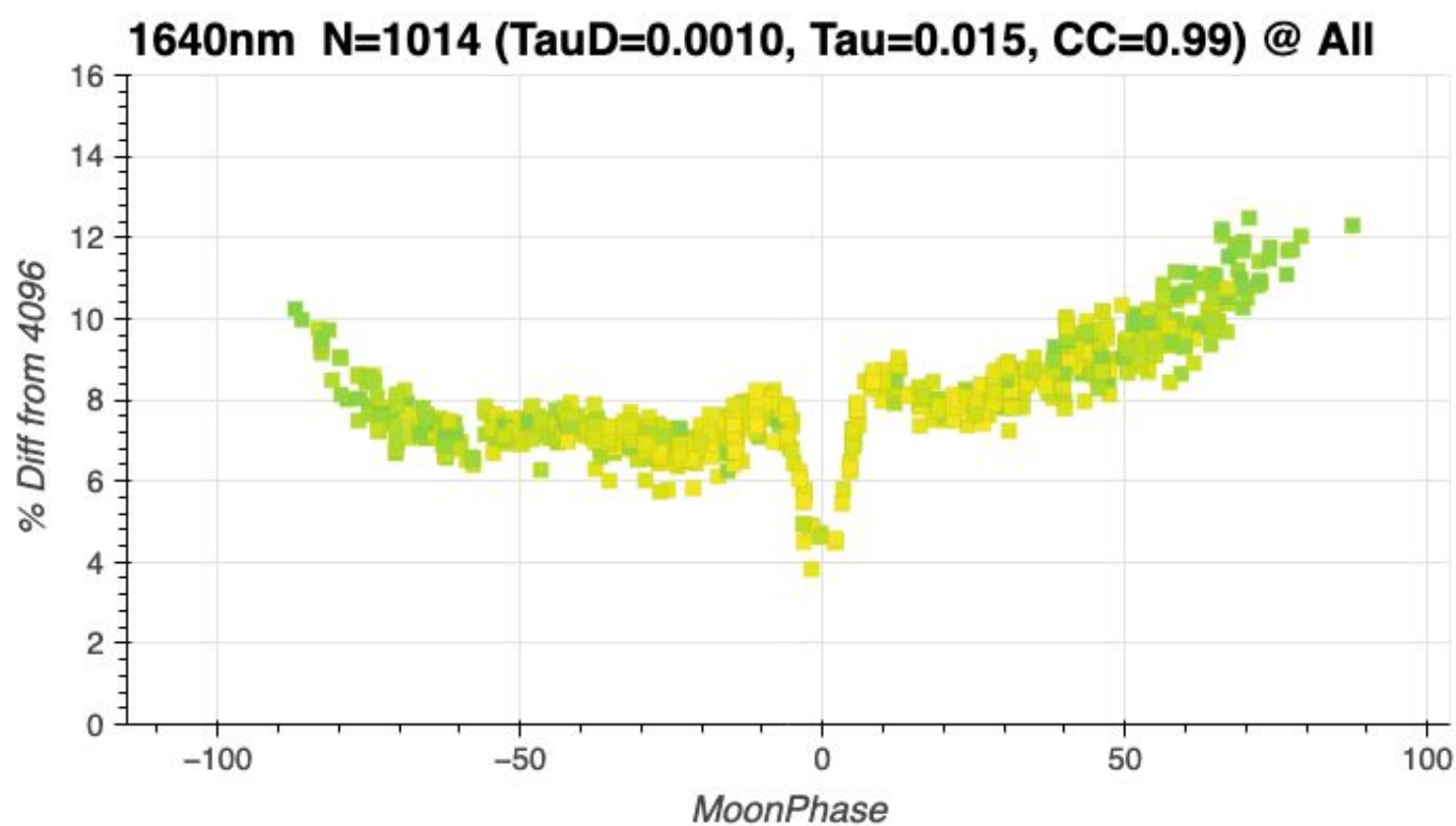
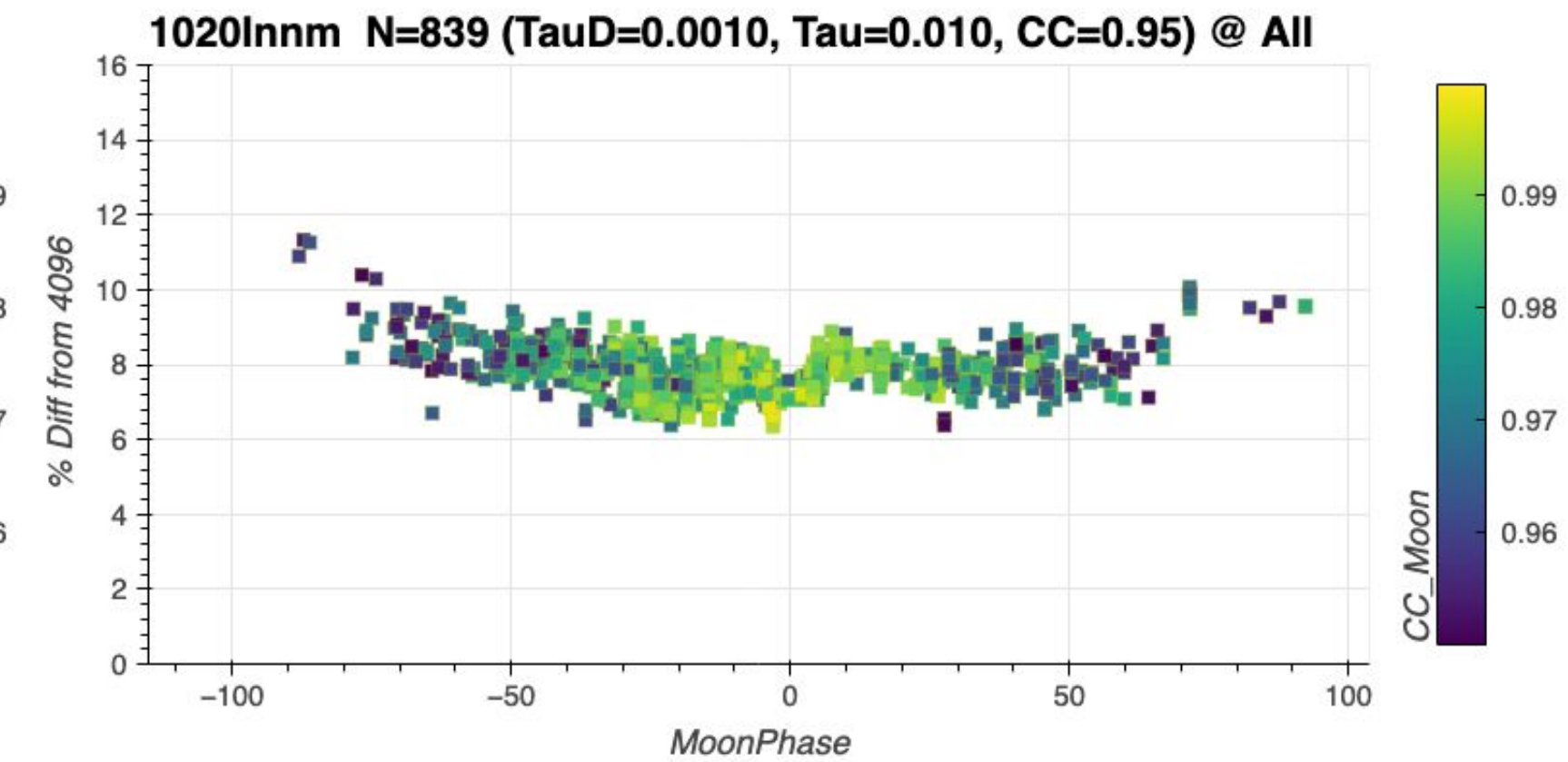
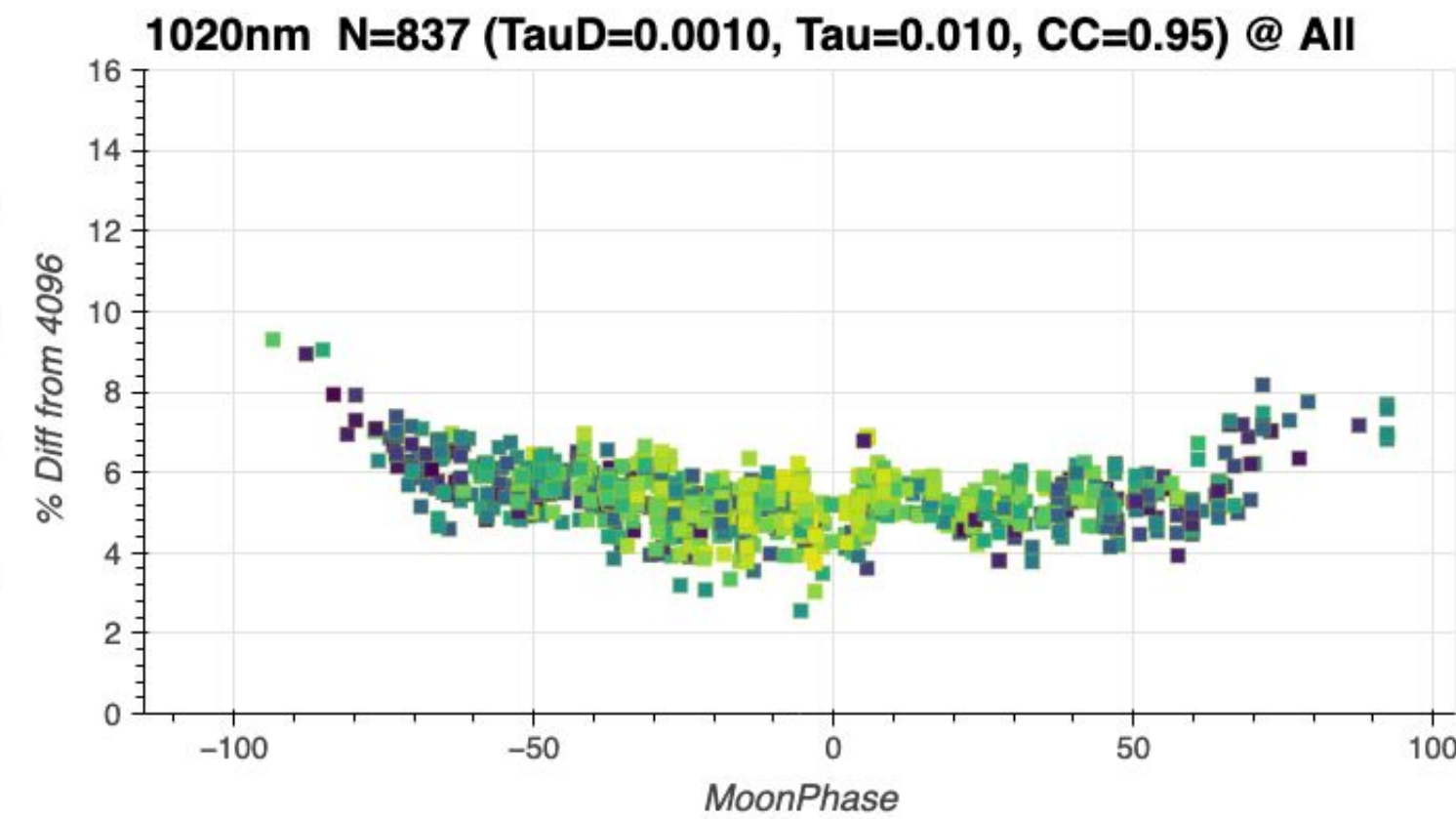
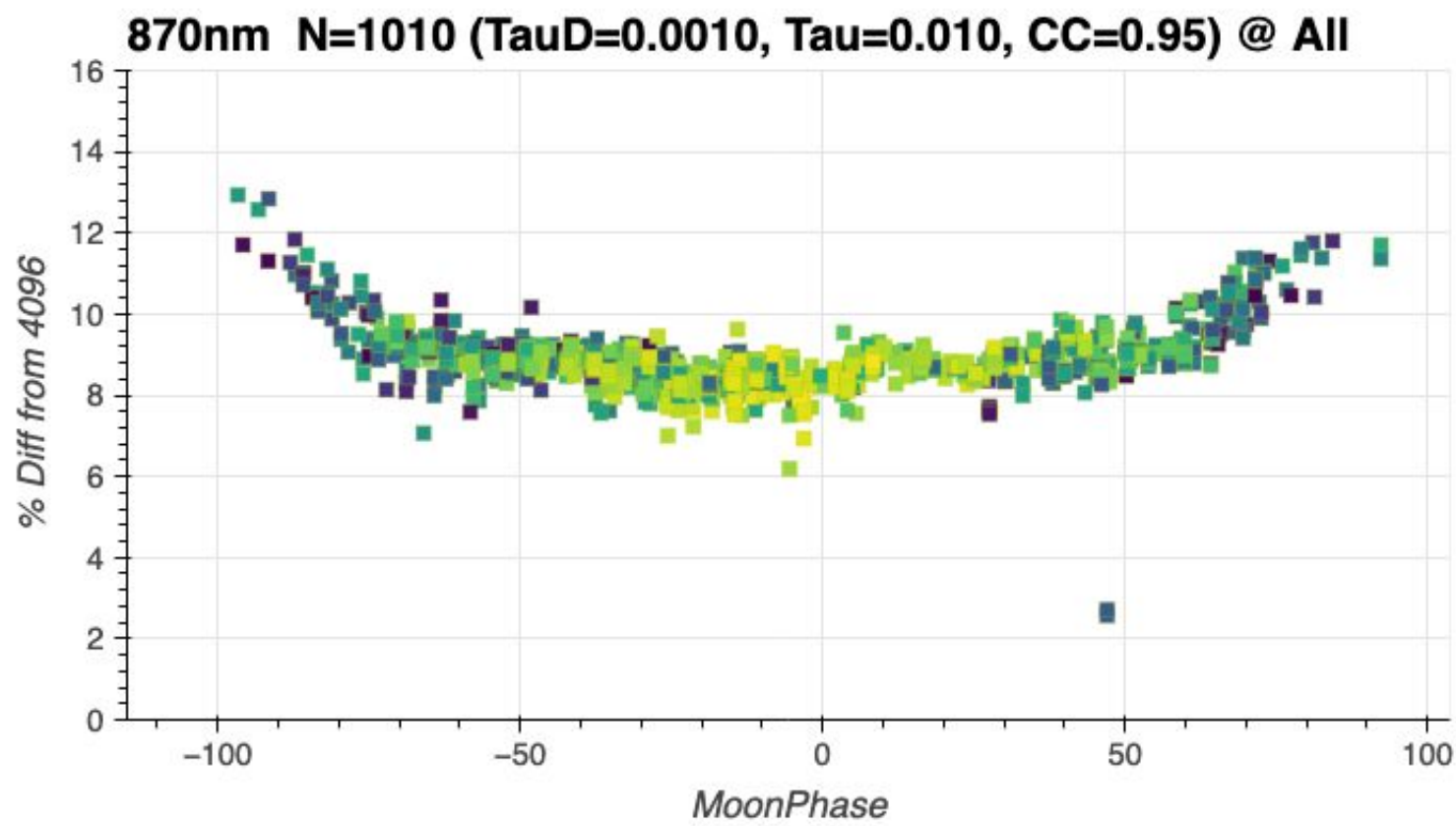
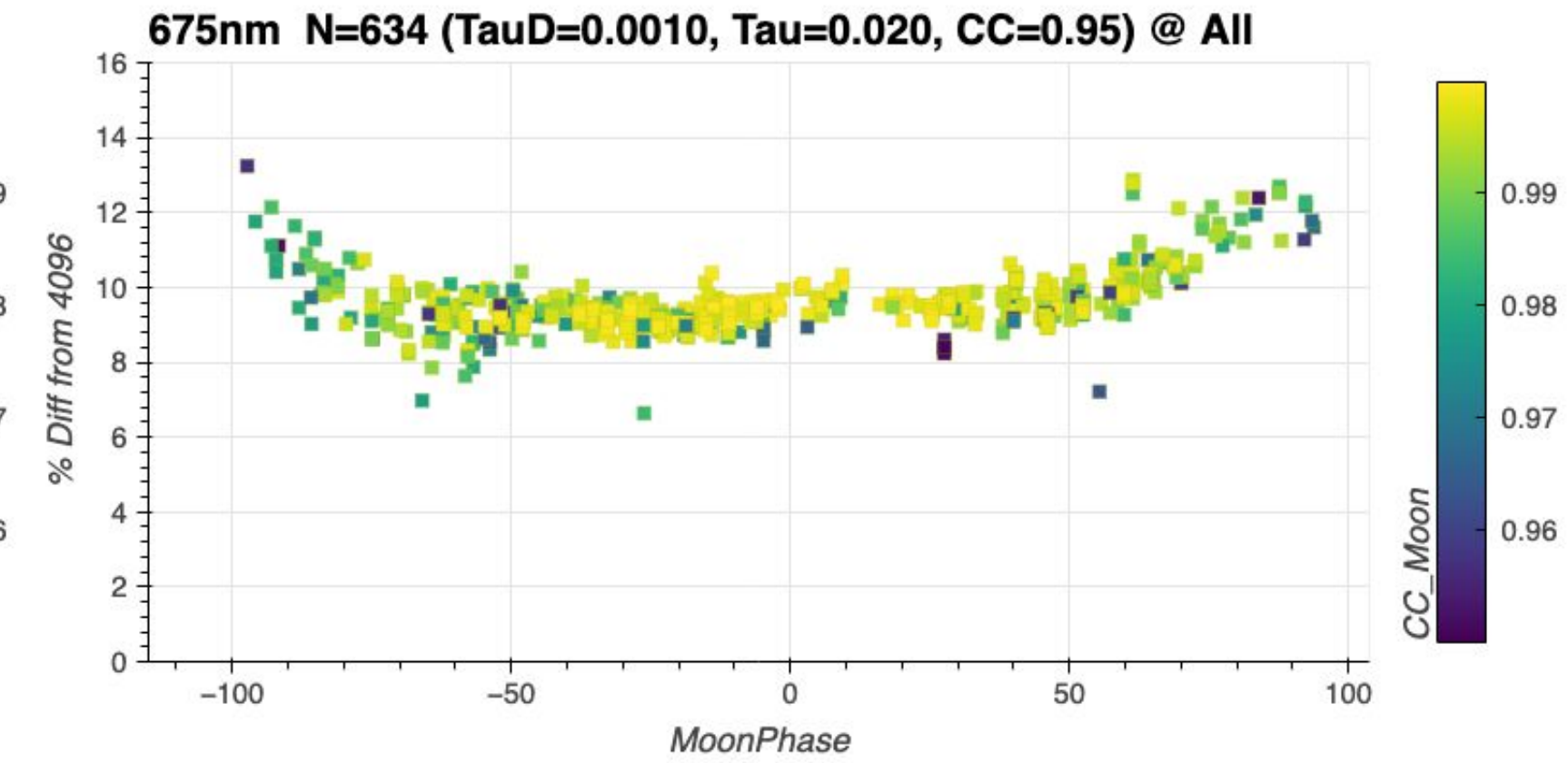
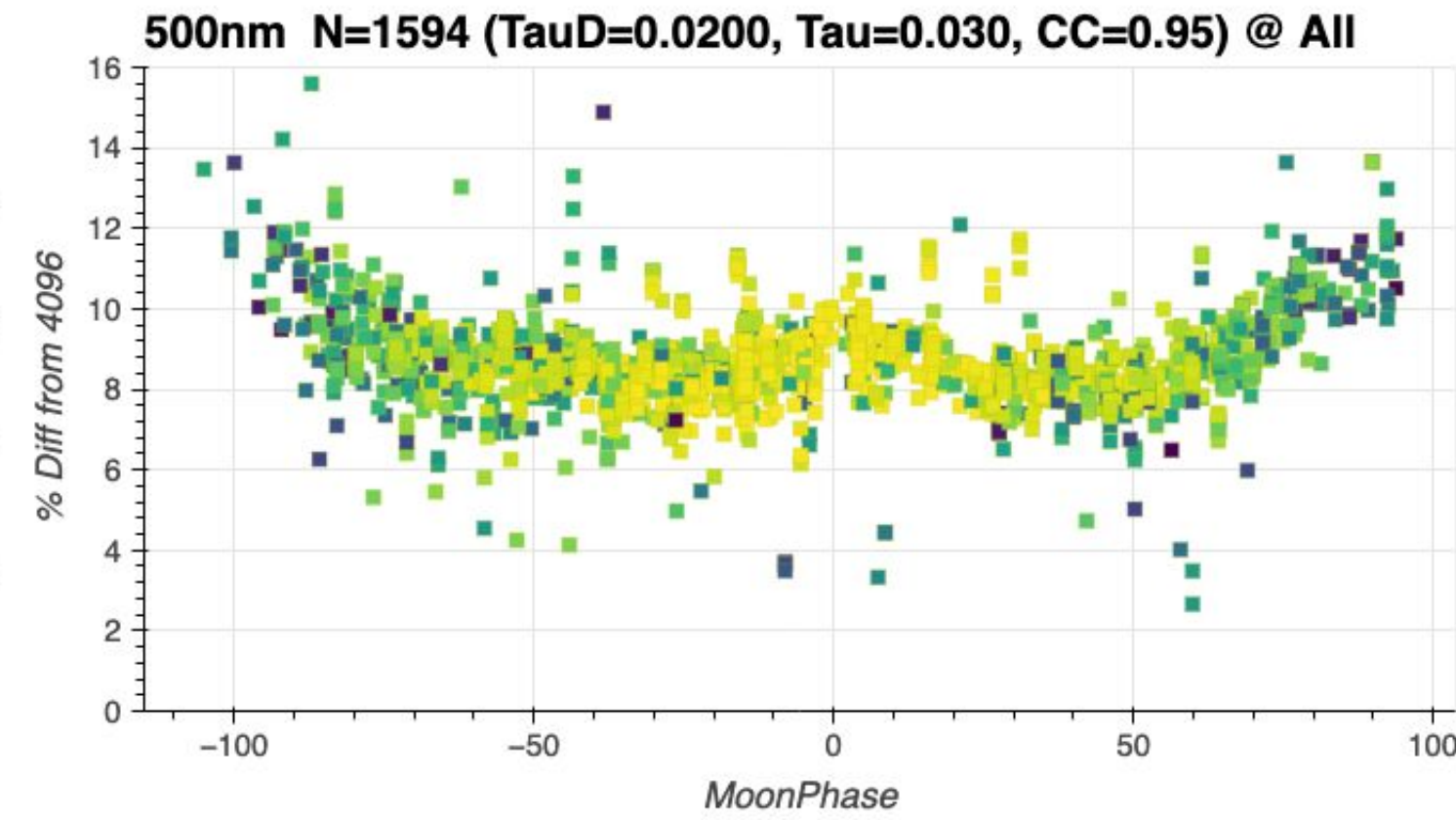
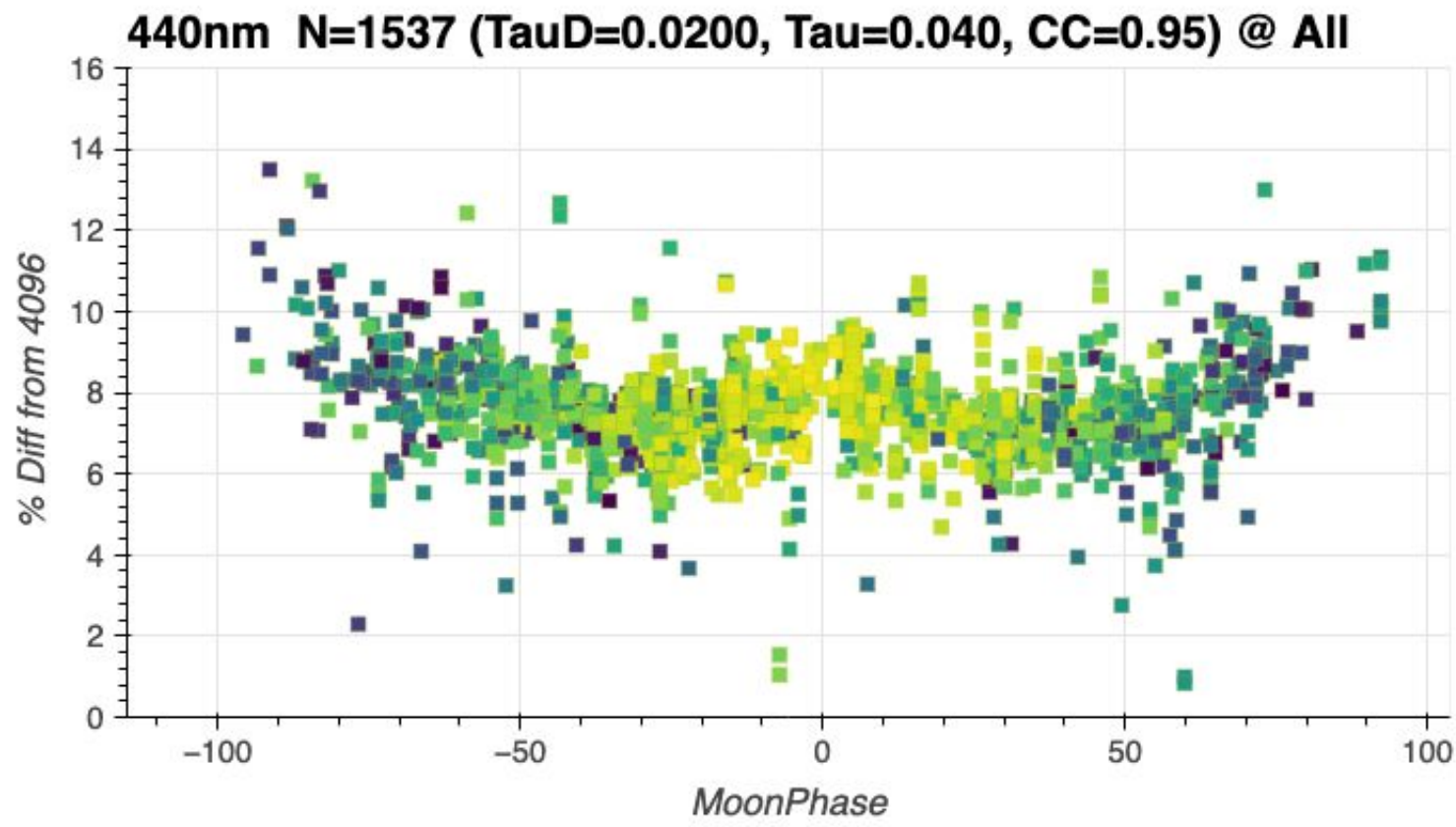
1020nm





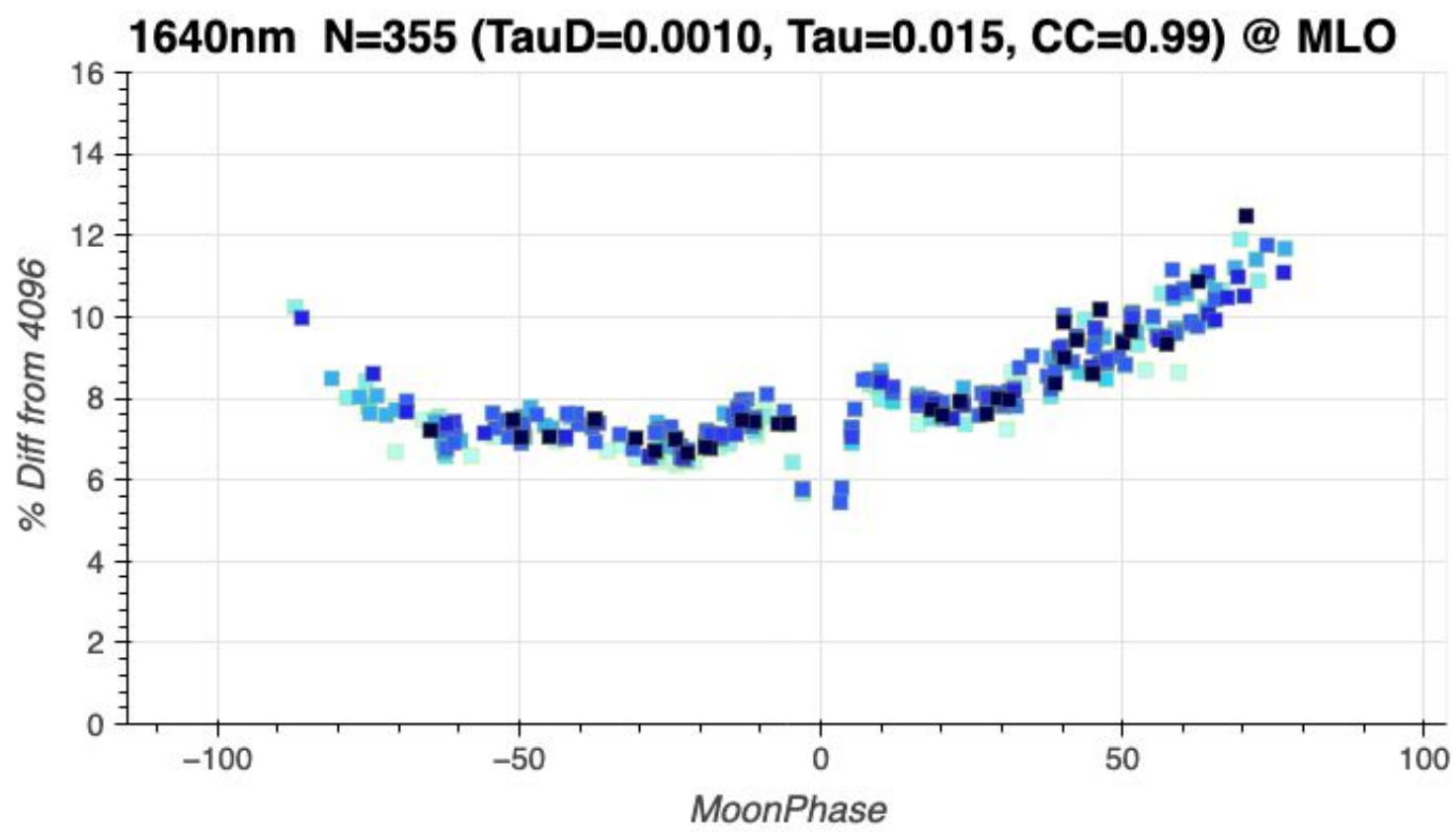
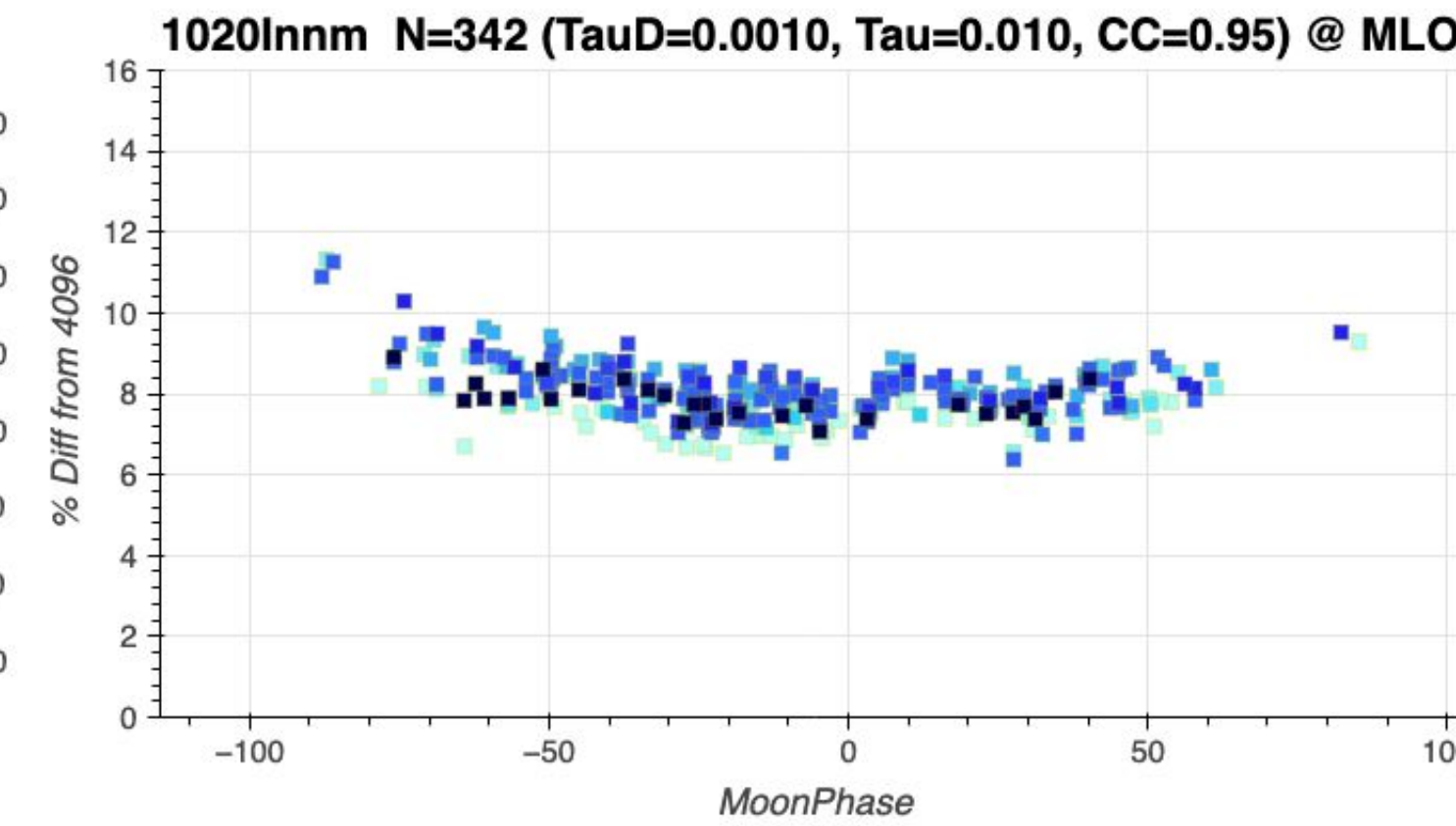
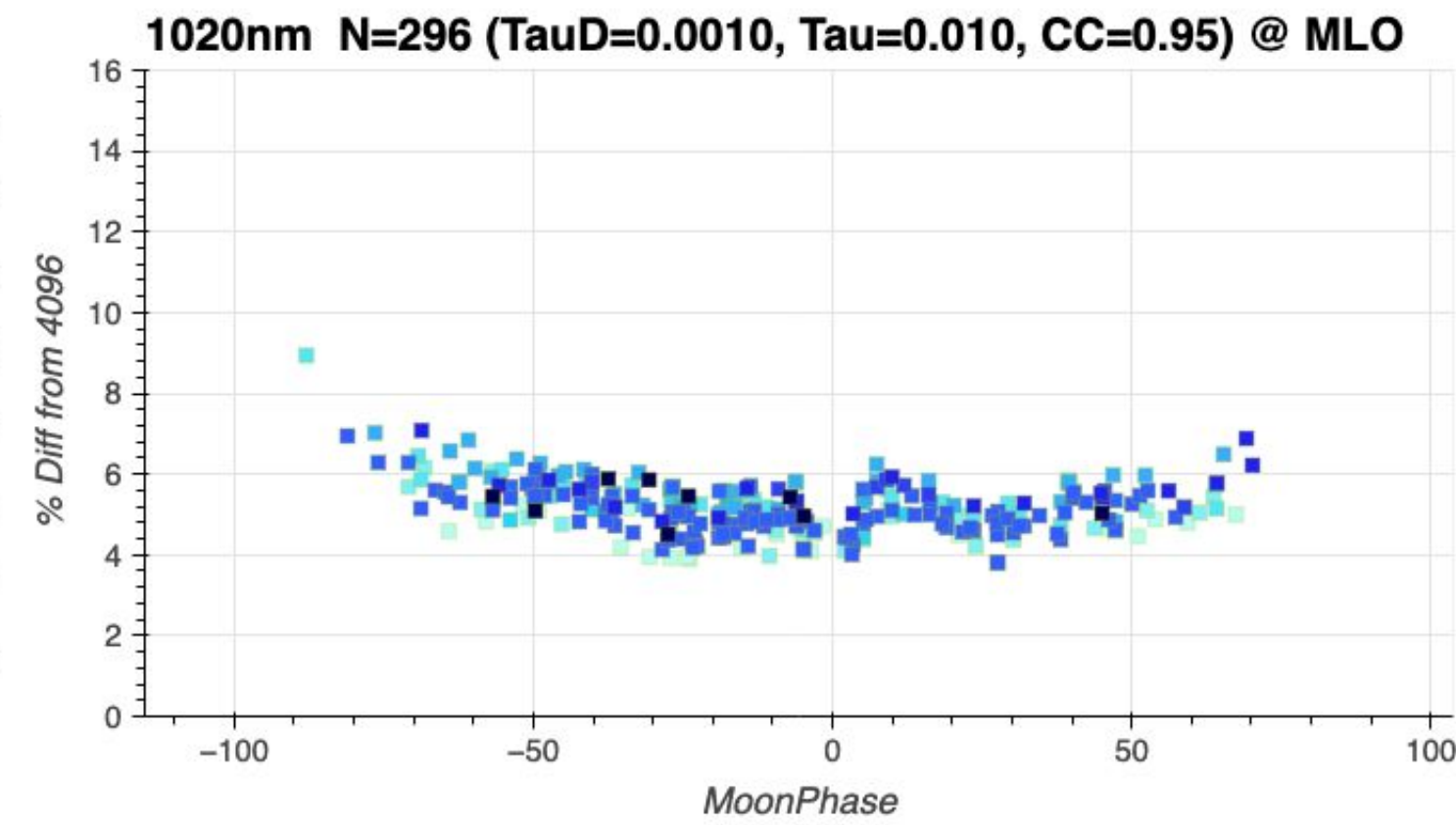
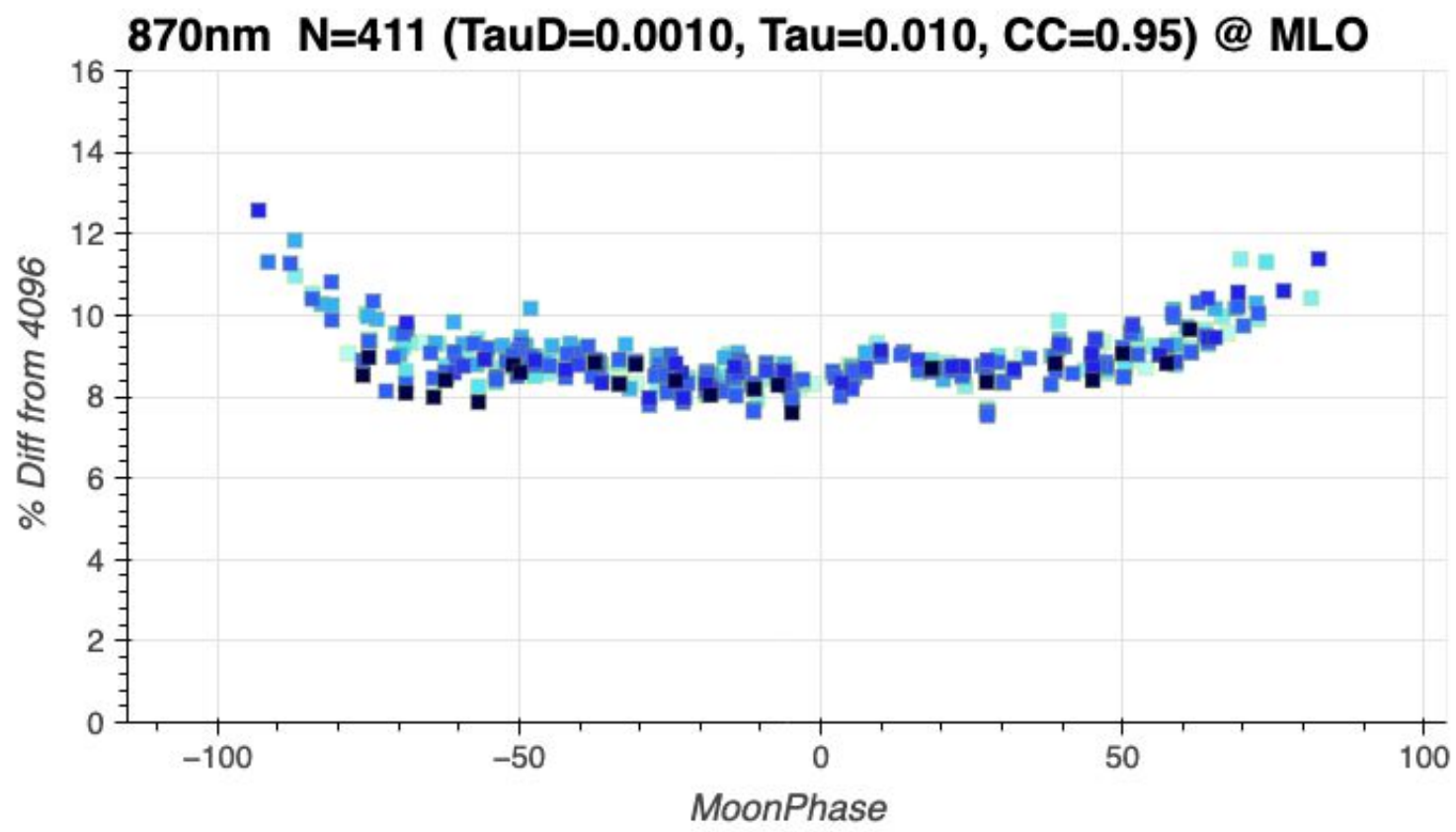
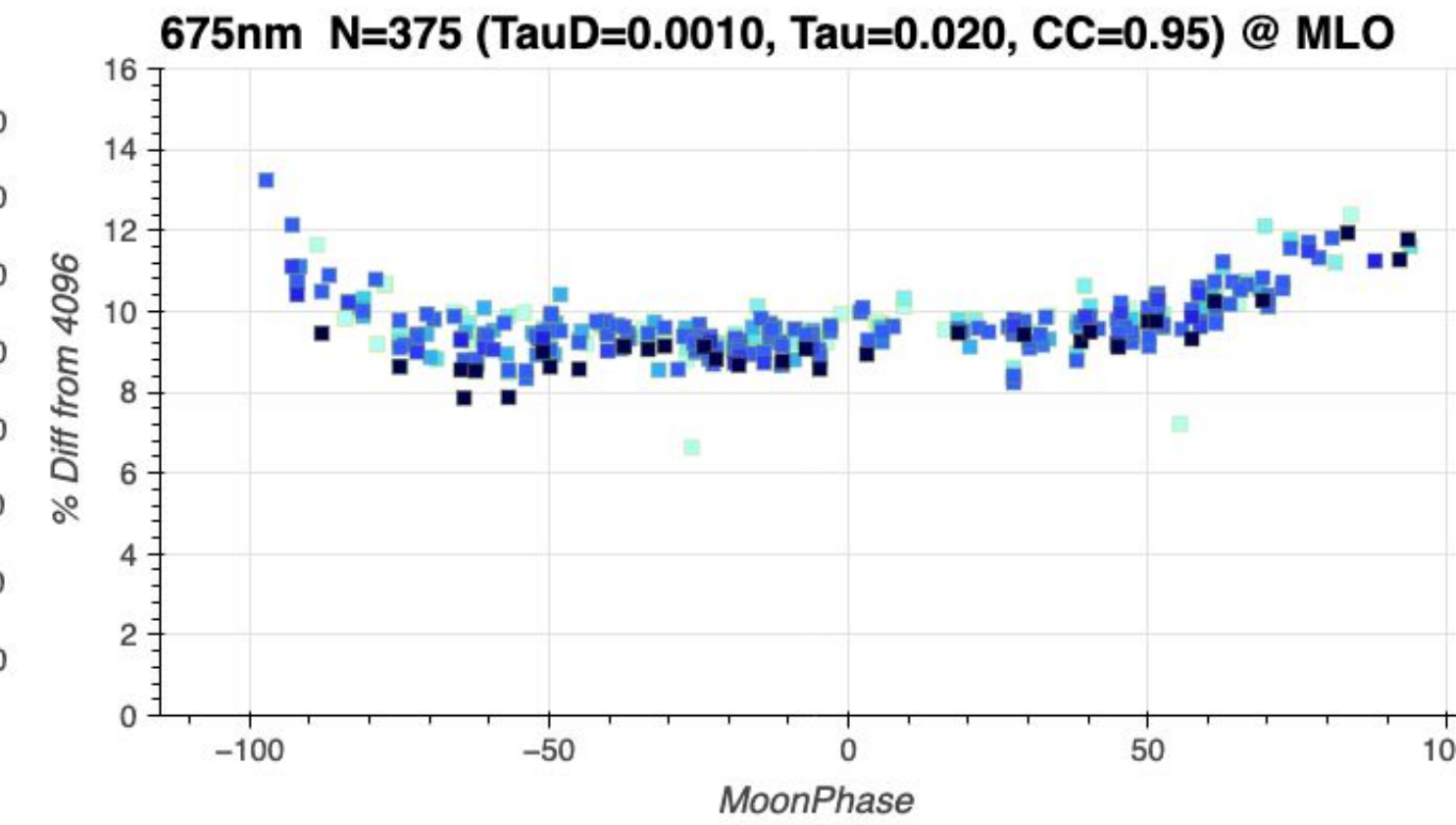
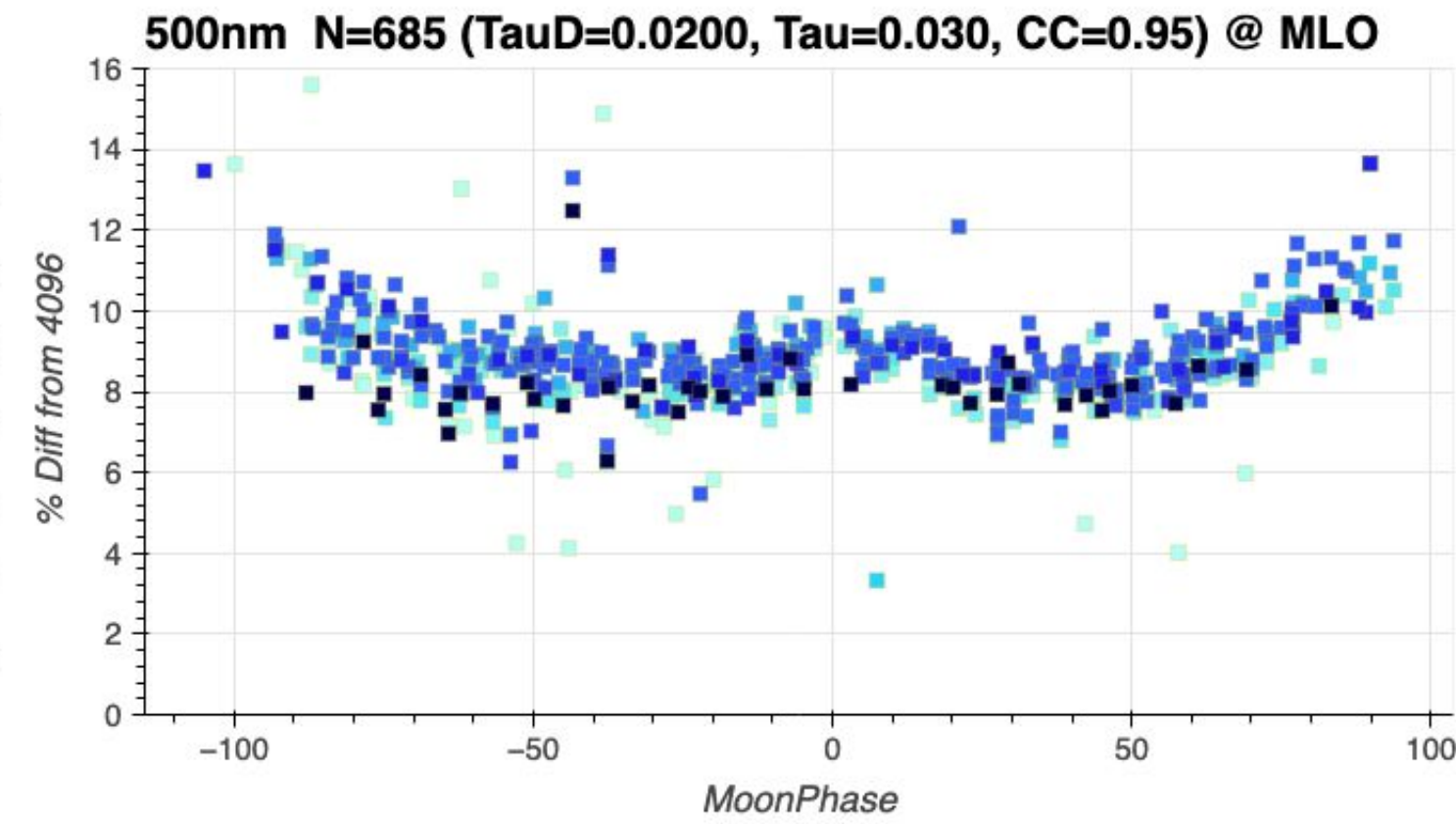
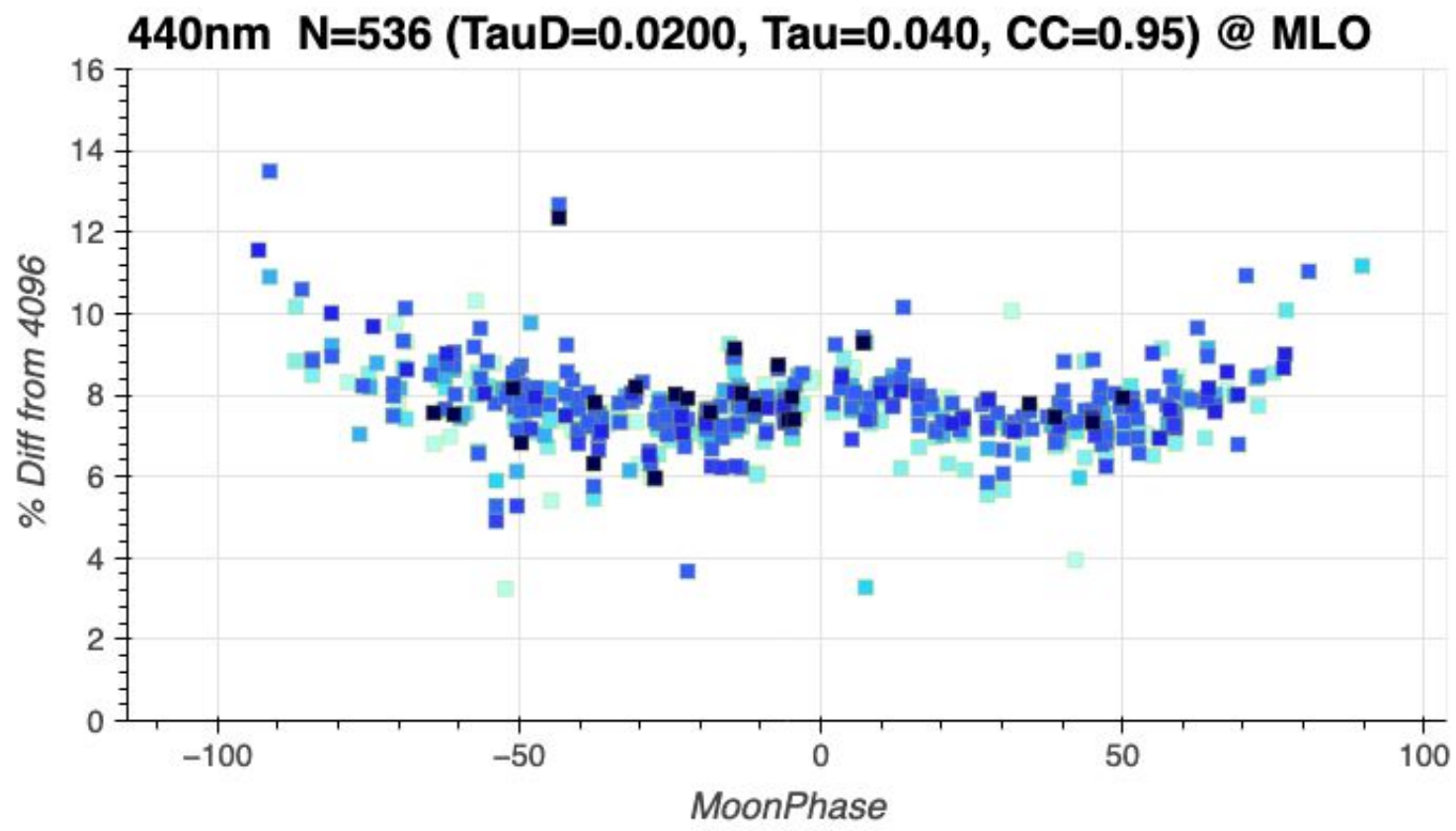


# Sensitivity to minimum CC



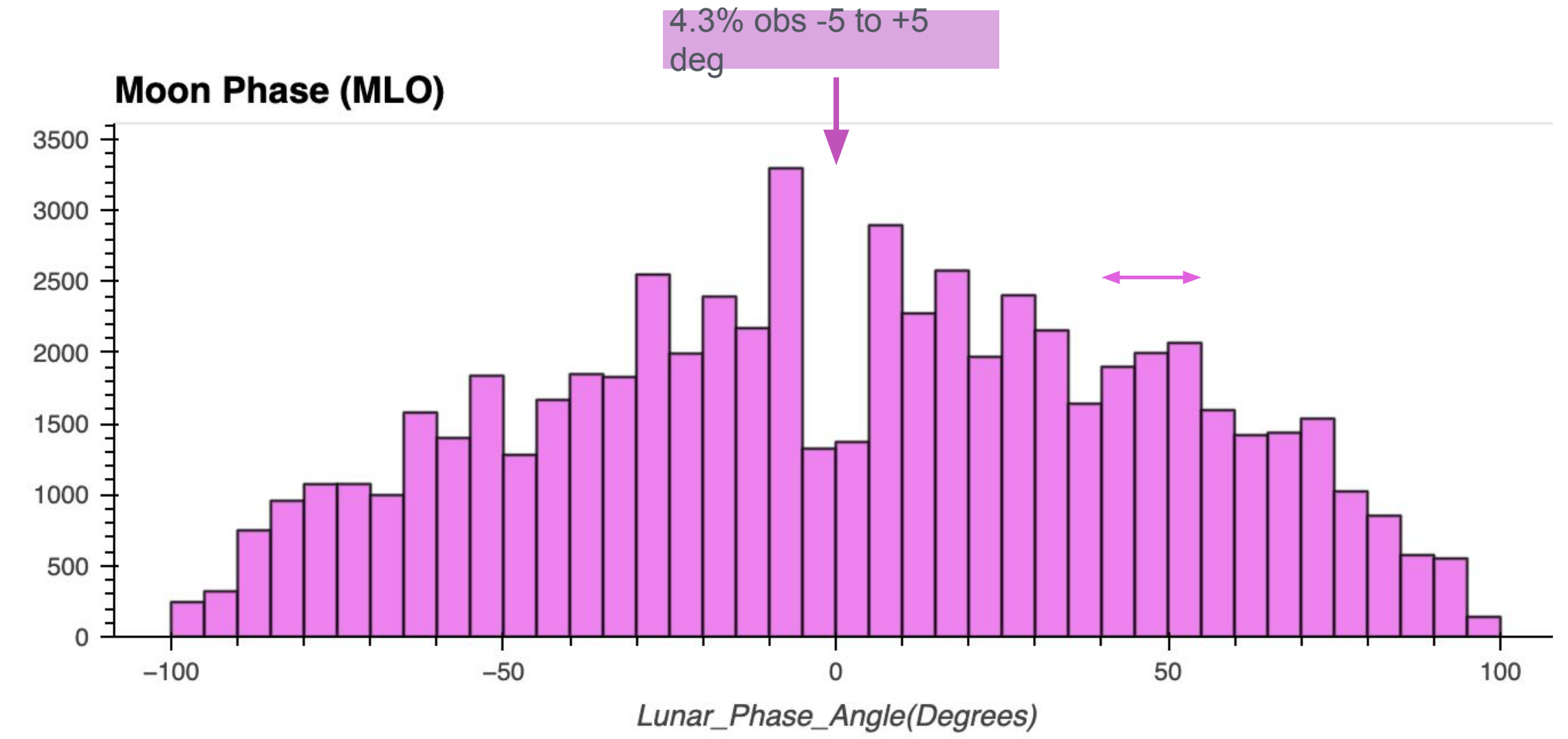
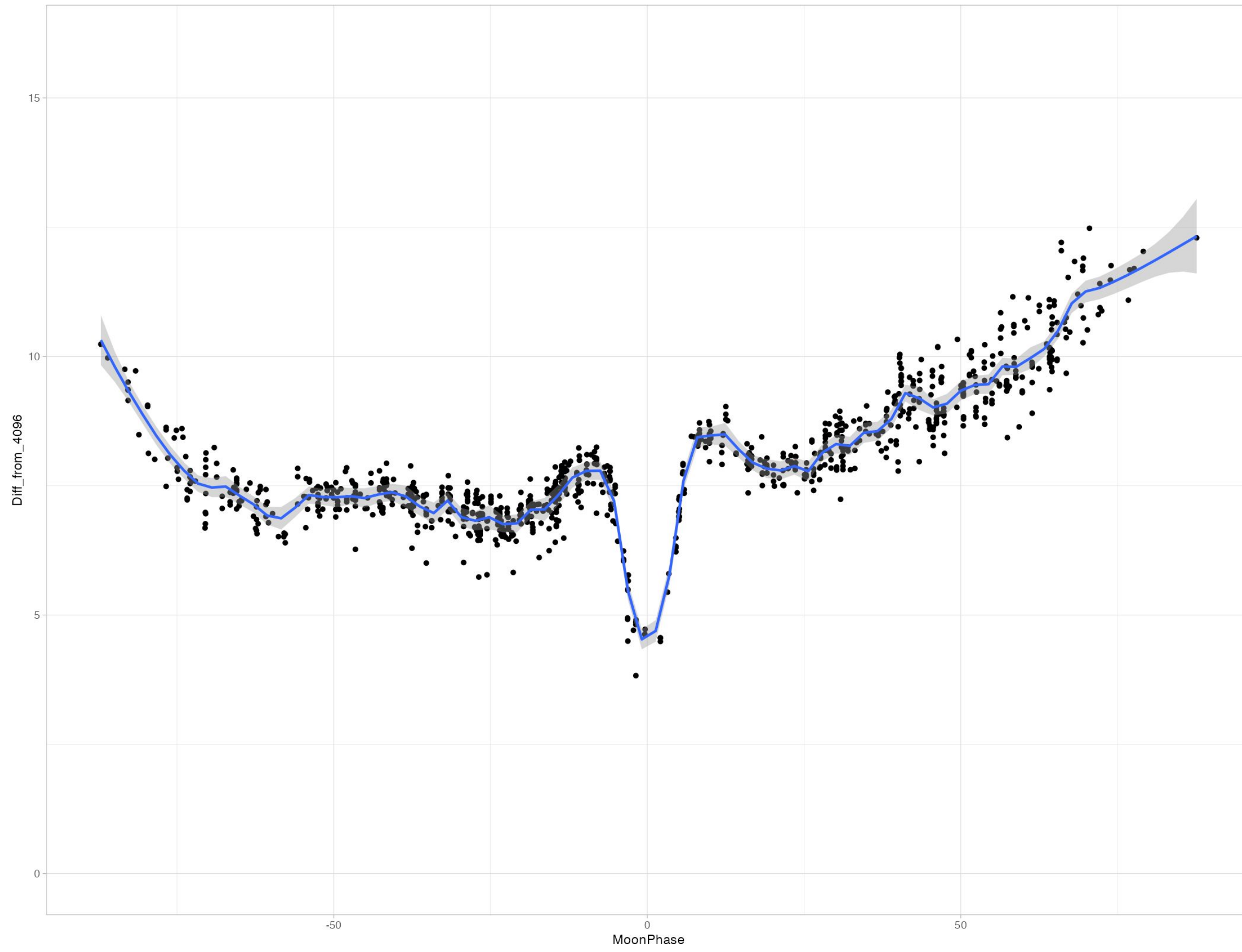
The main goal was to capture more of the large (absolute) MPA  
The loss of these observations was mostly due to the CC limit

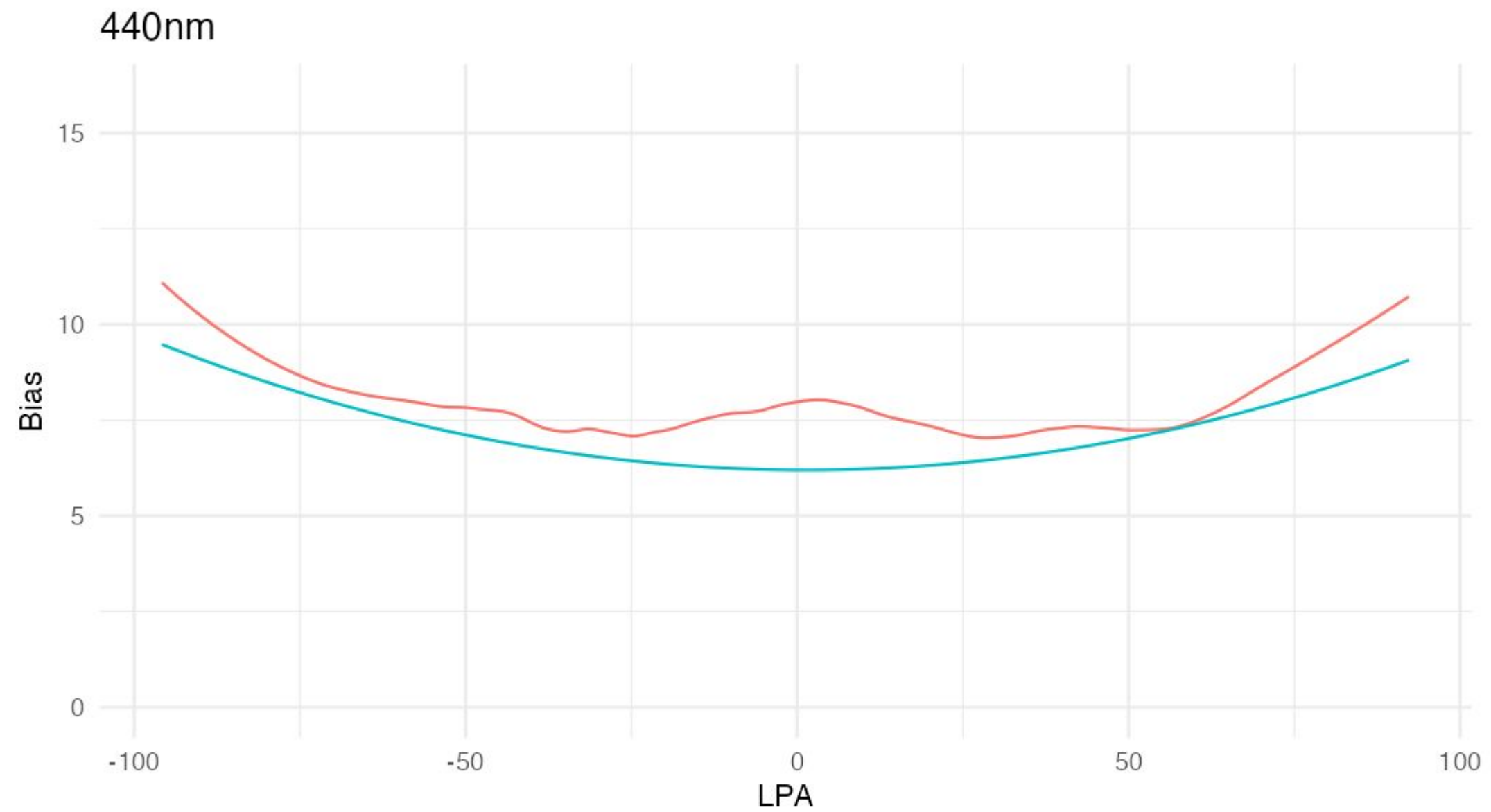
MLO only



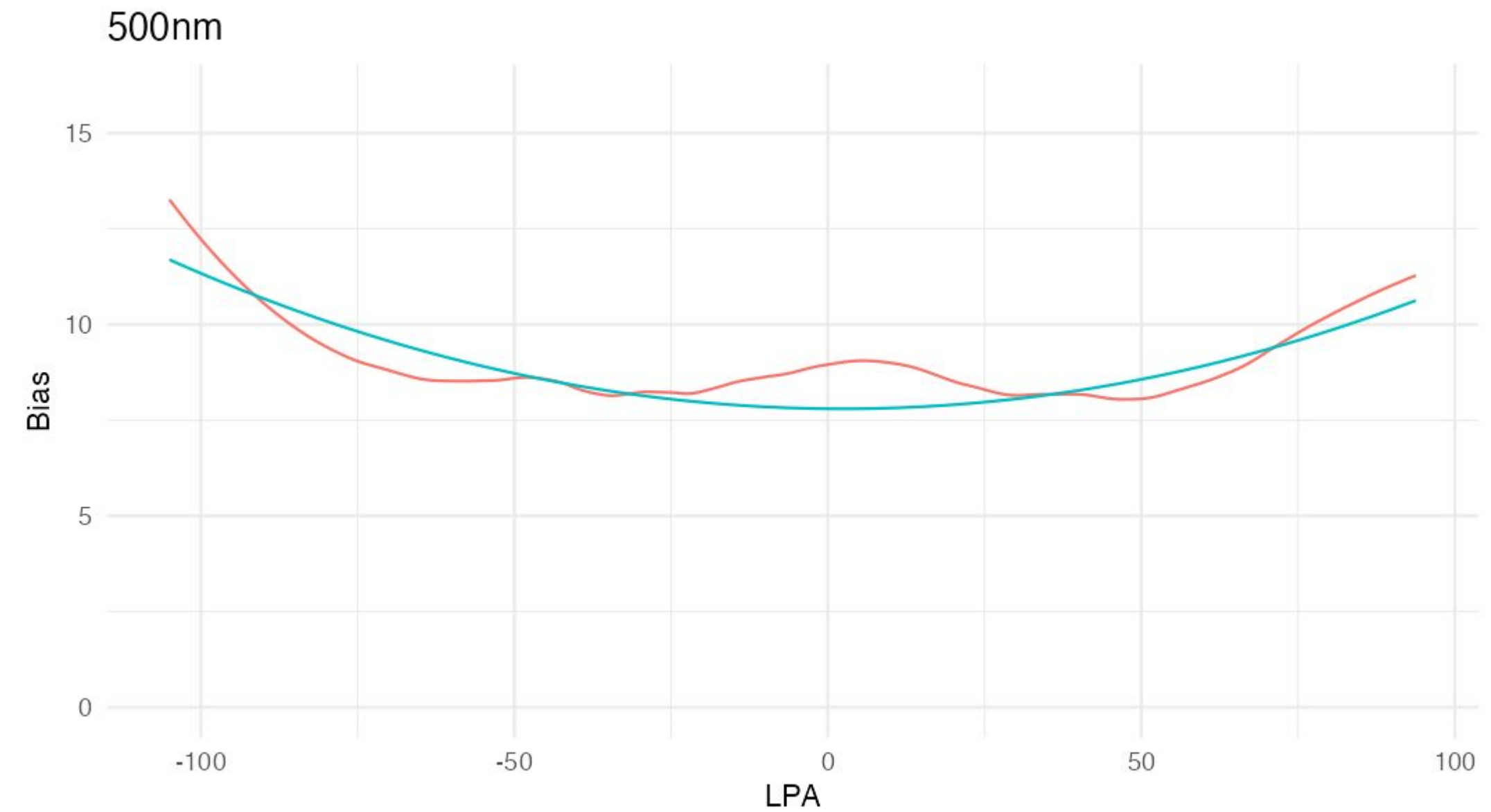
MLO remains a much less variable dataset than Izaña

1640nm, span= 0.05

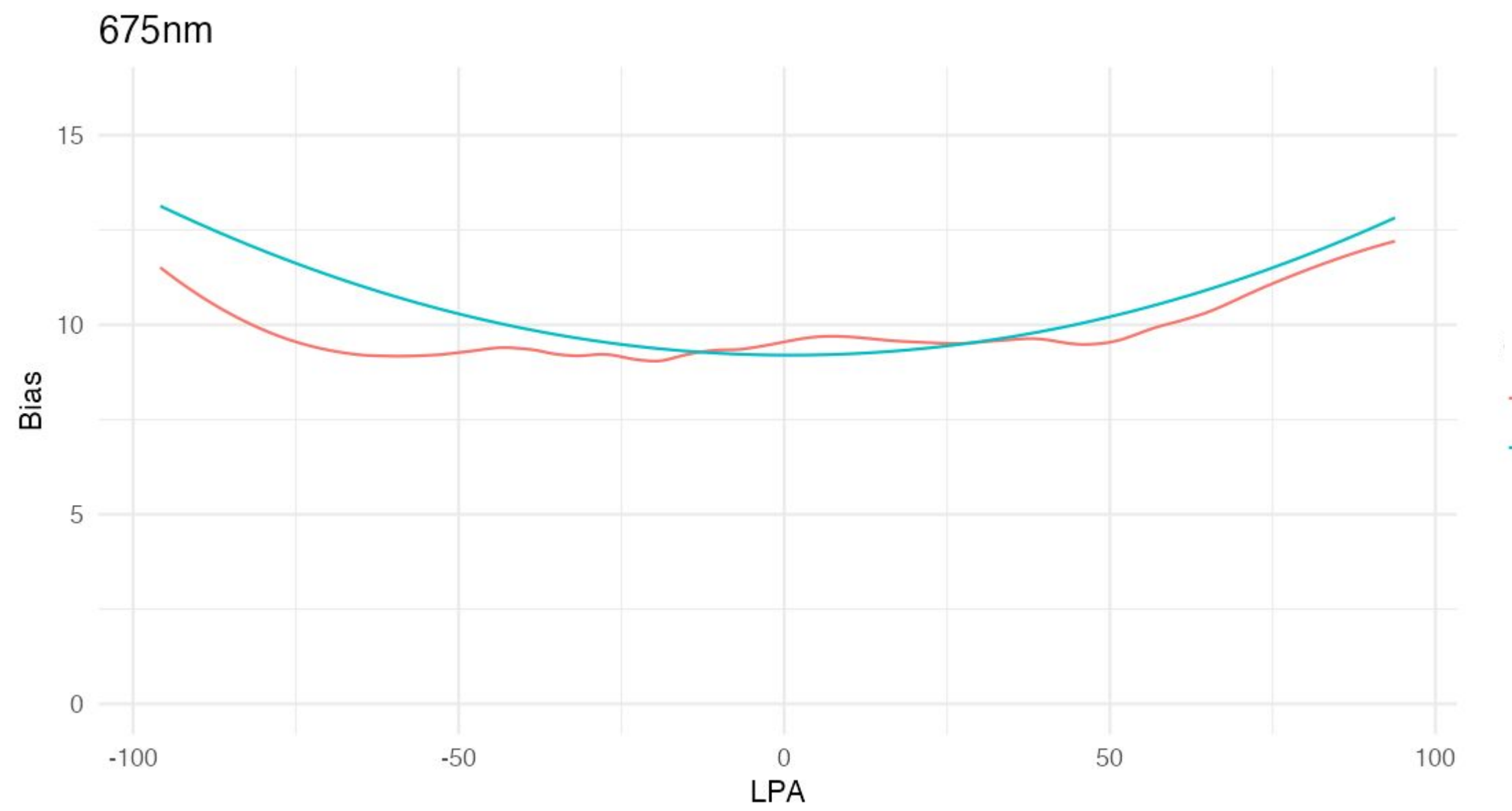




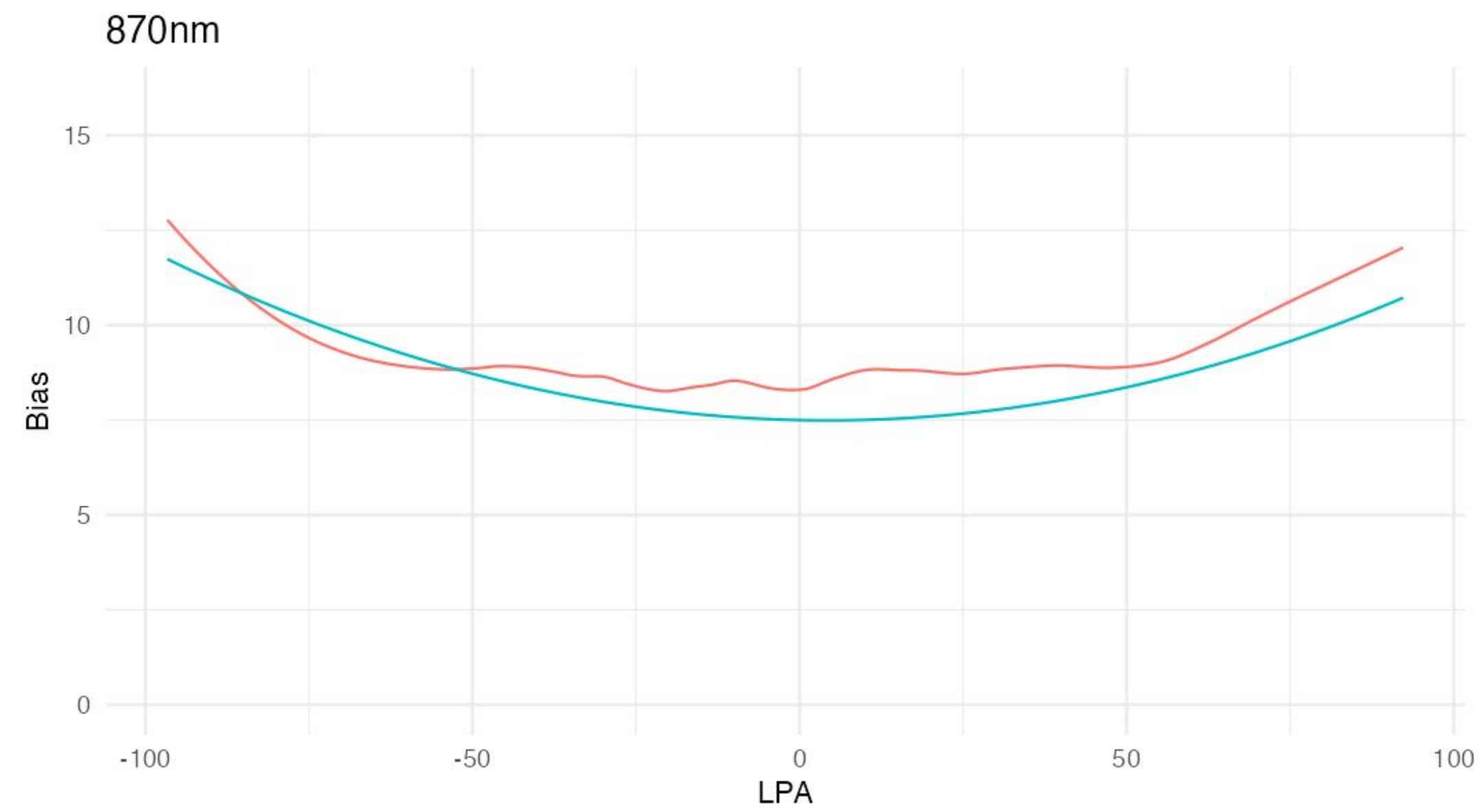
**Fit**  
— LUT  
— RCF



**Fit**  
— LUT  
— RCF



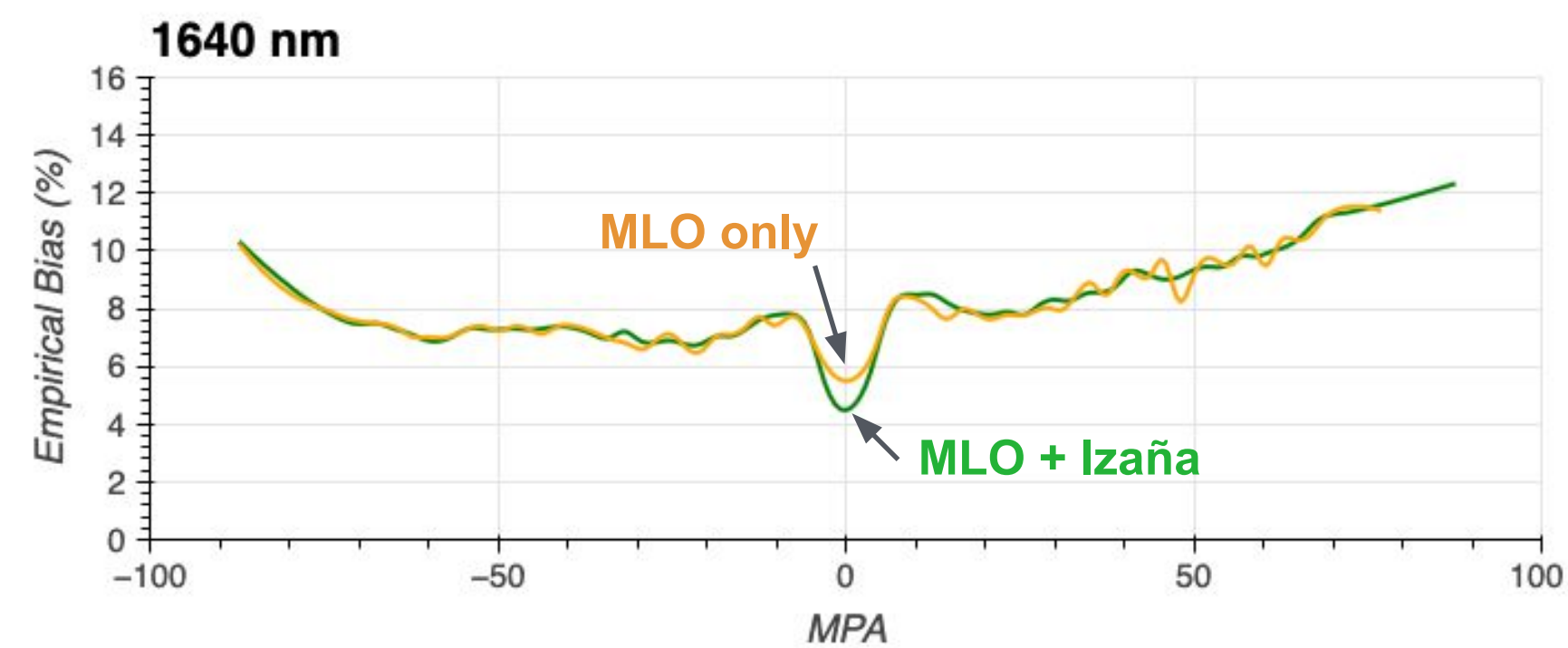
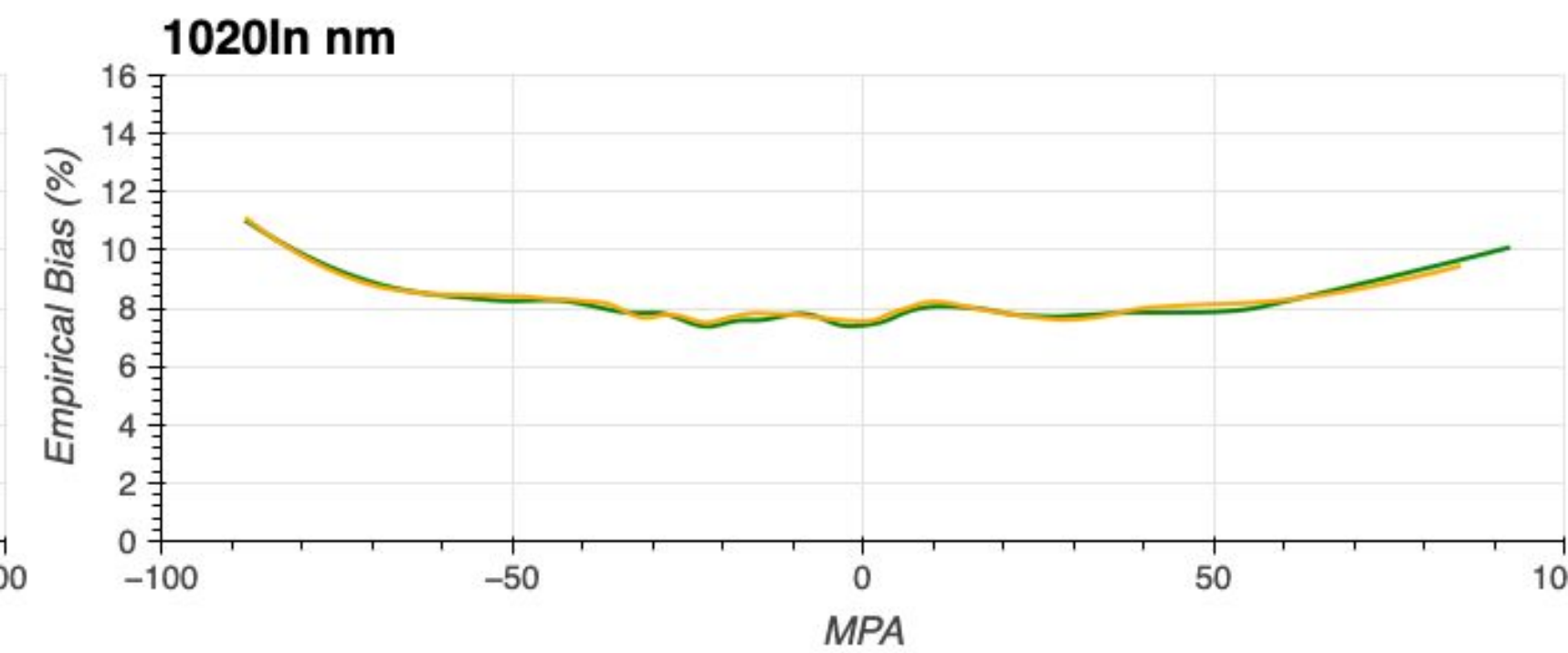
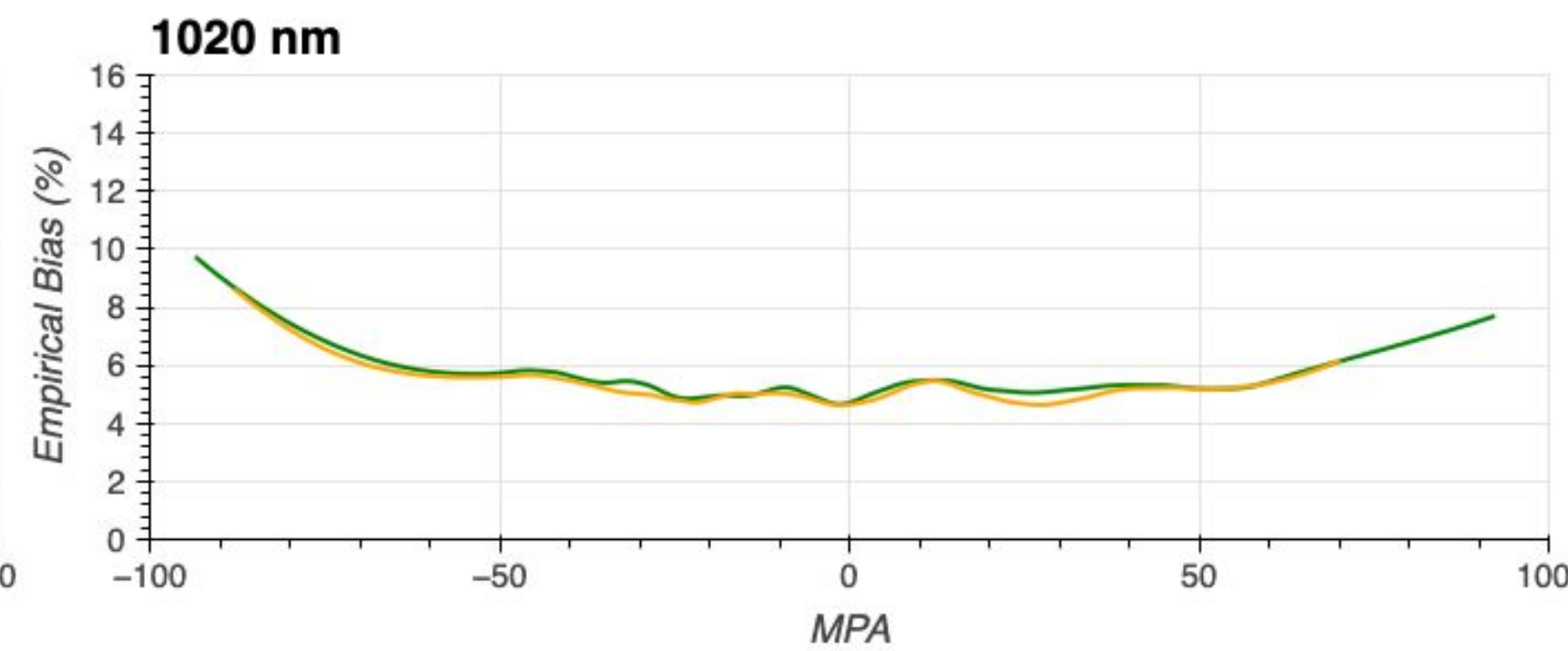
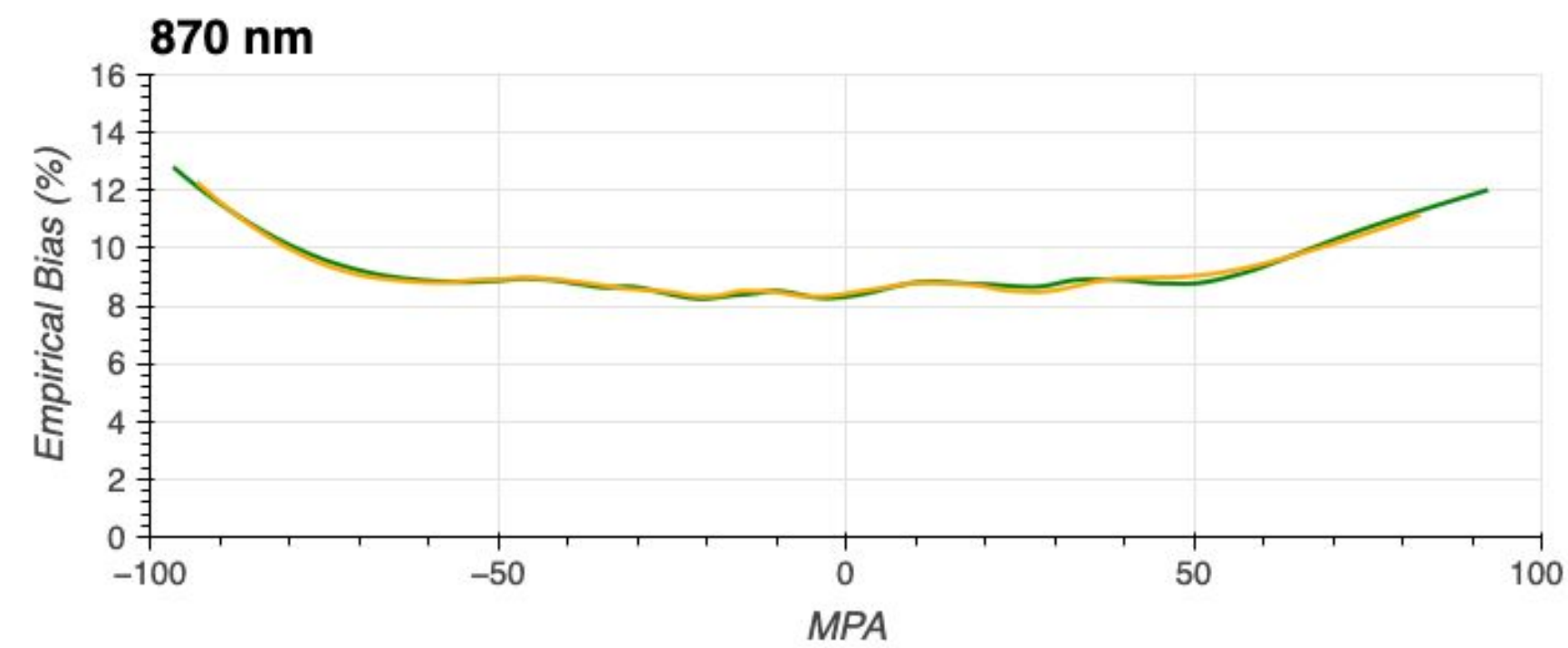
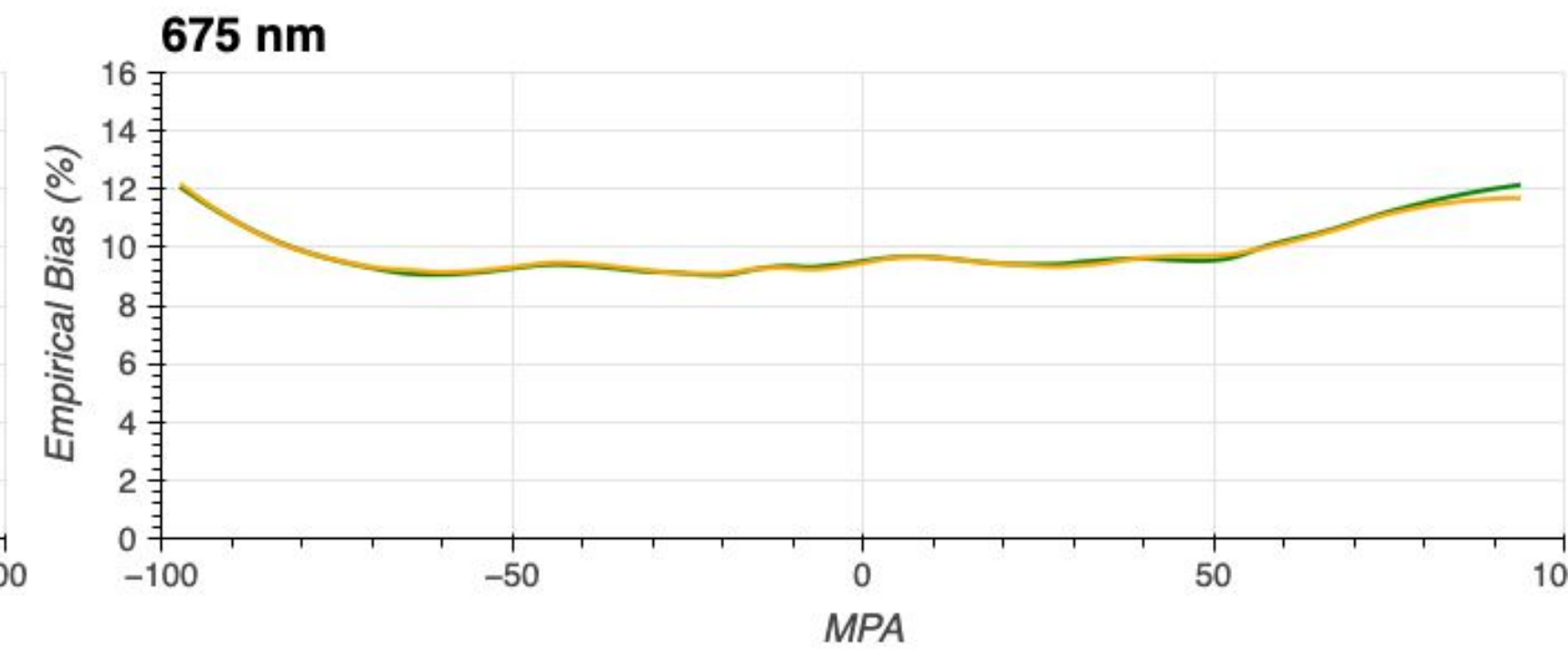
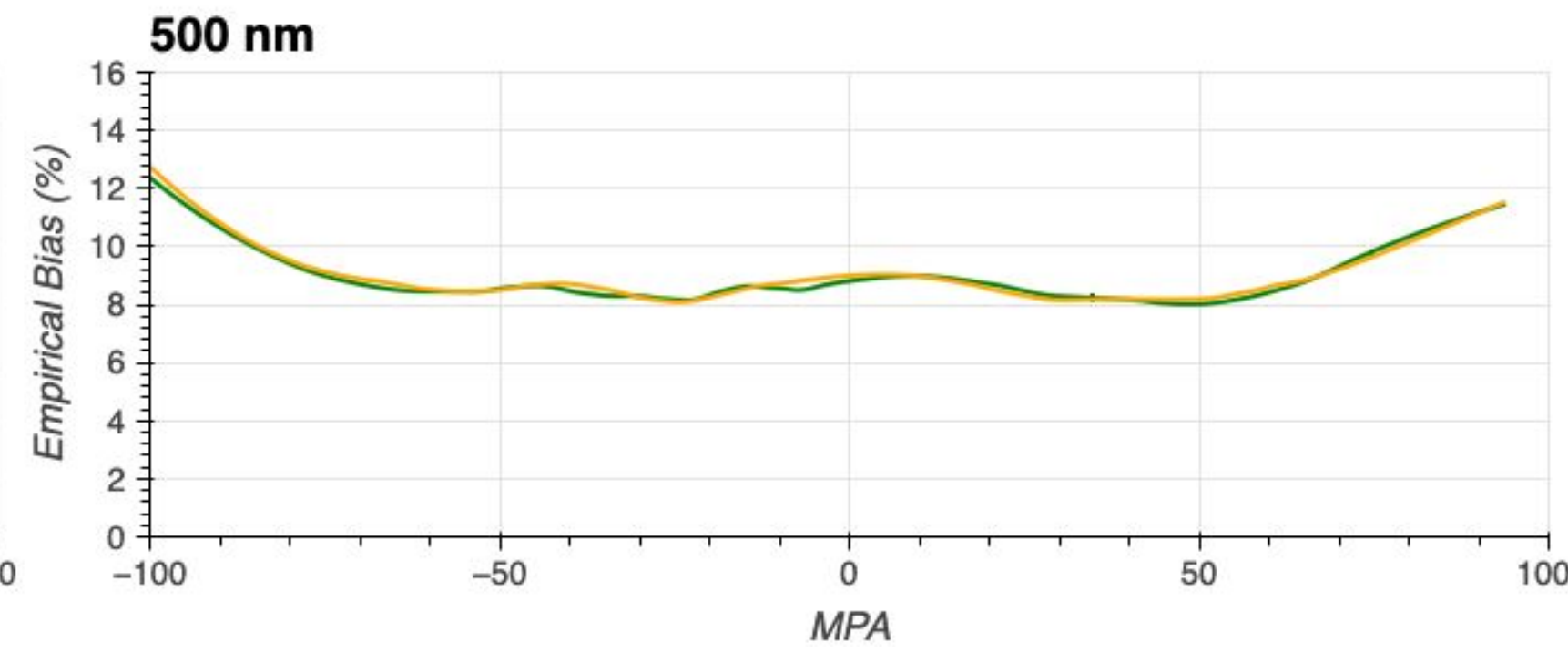
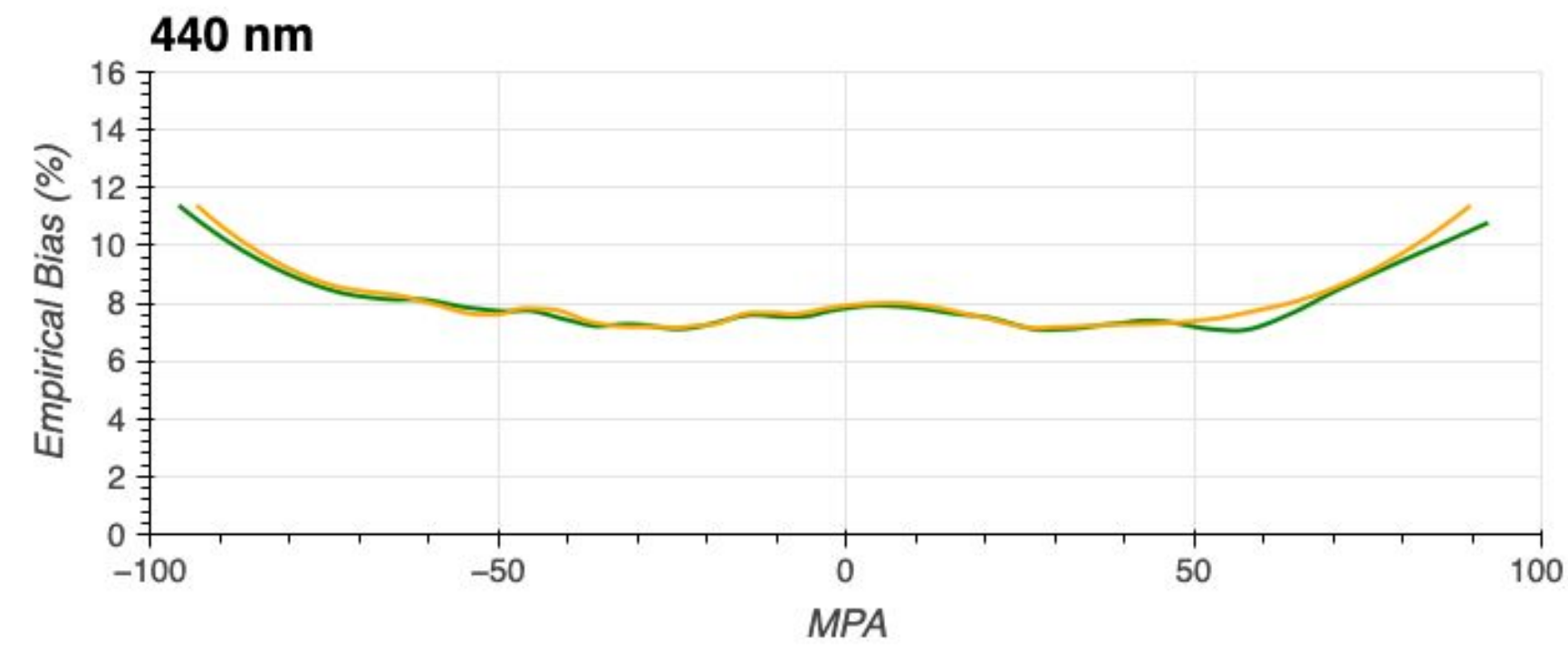
**Fit**  
— LUT  
— RCF



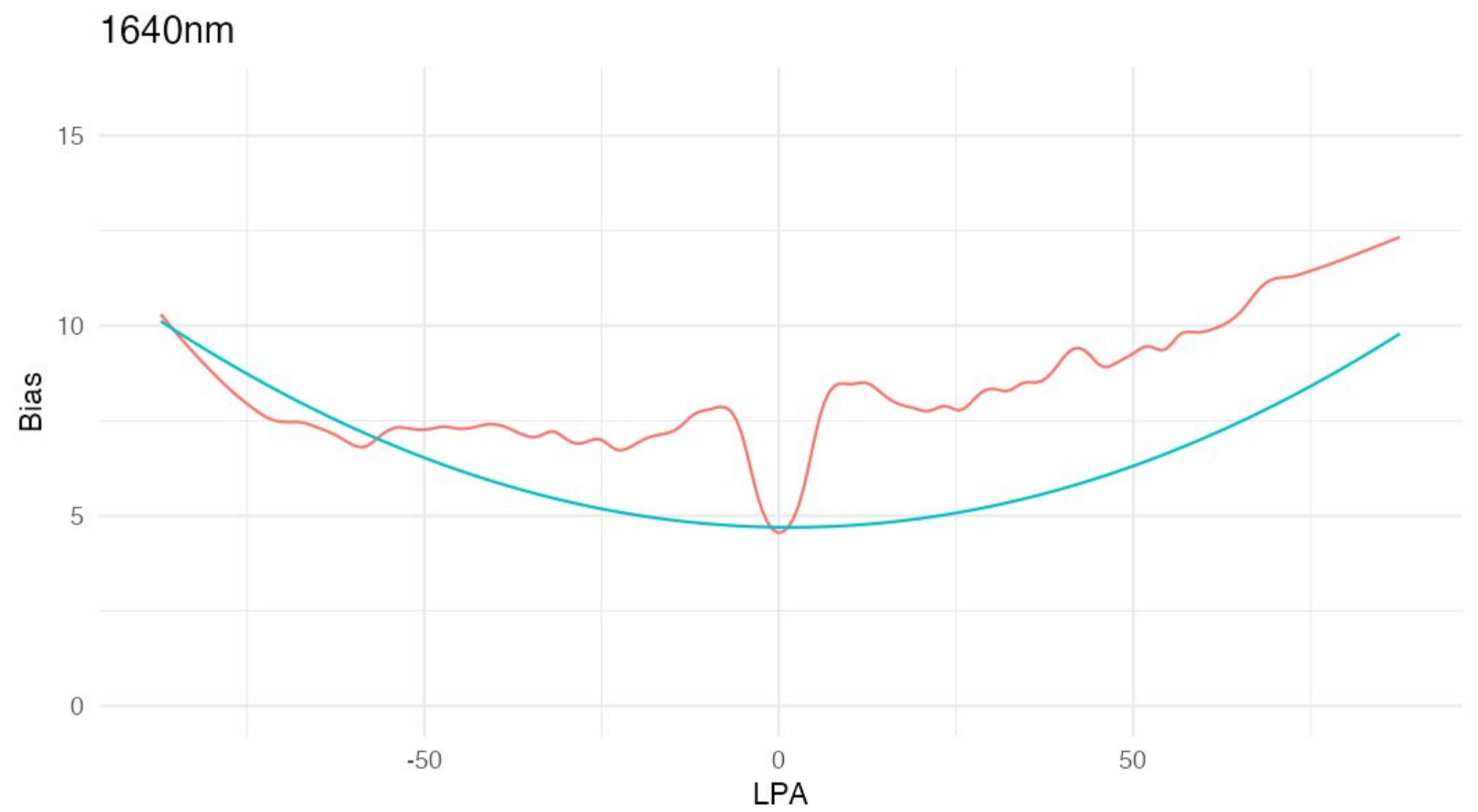
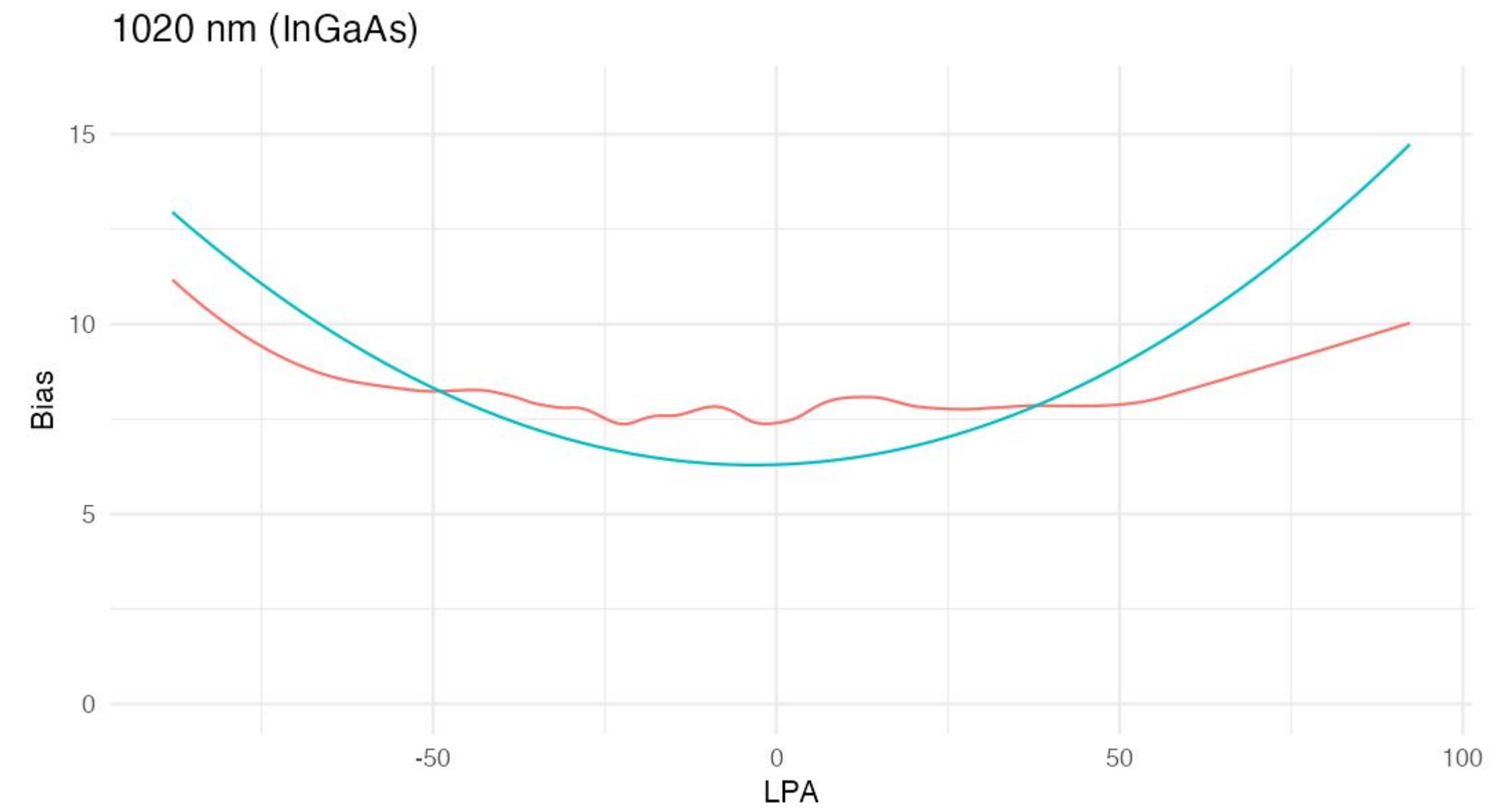
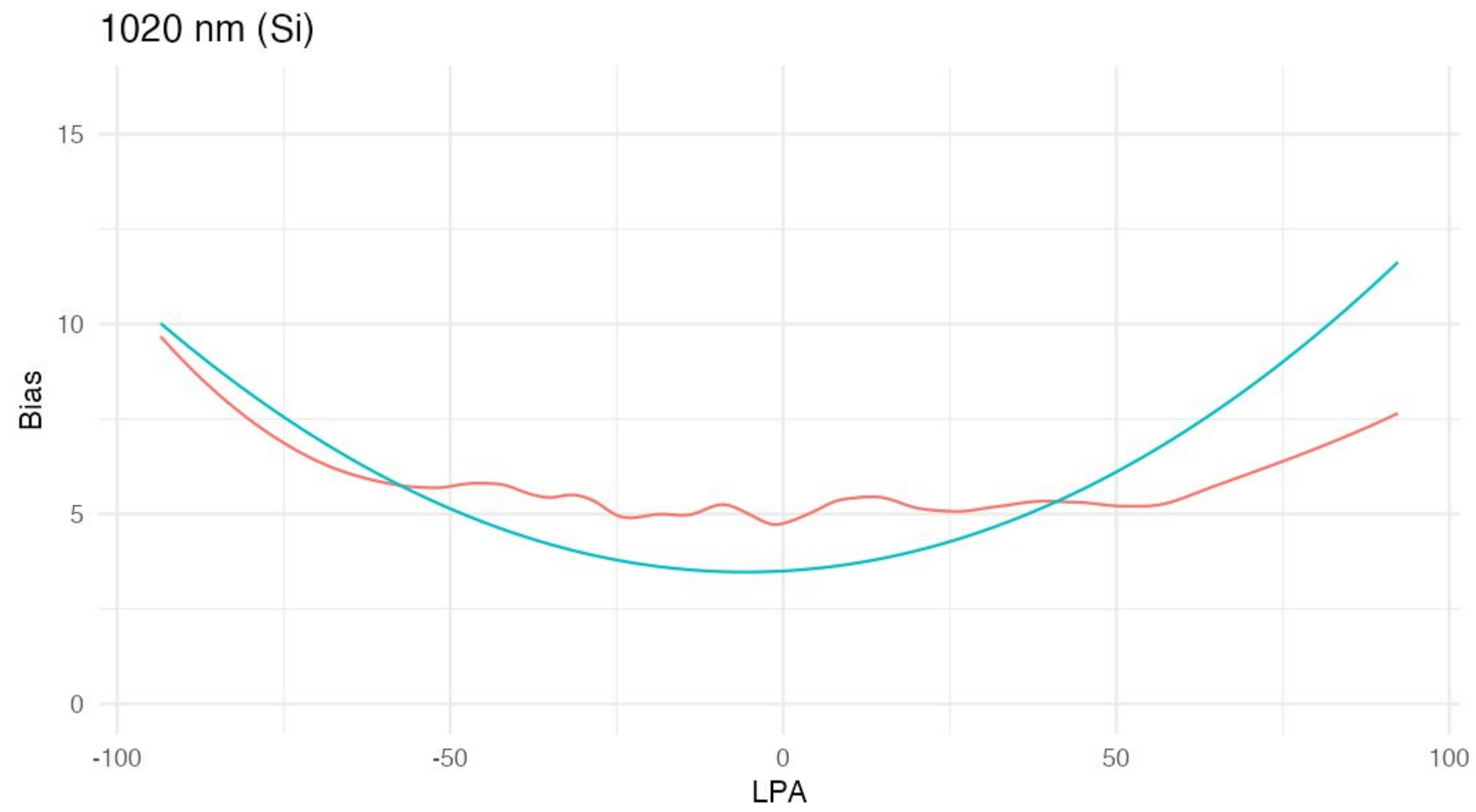
**Fit**  
— LUT  
— RCF

# Comparing LUT and RCF Empirical Bias Correction

LOESS Fit is quite similar for Combined Sites and MLO Only  
Likewise, the form of the trend is very similar except for 1640nm



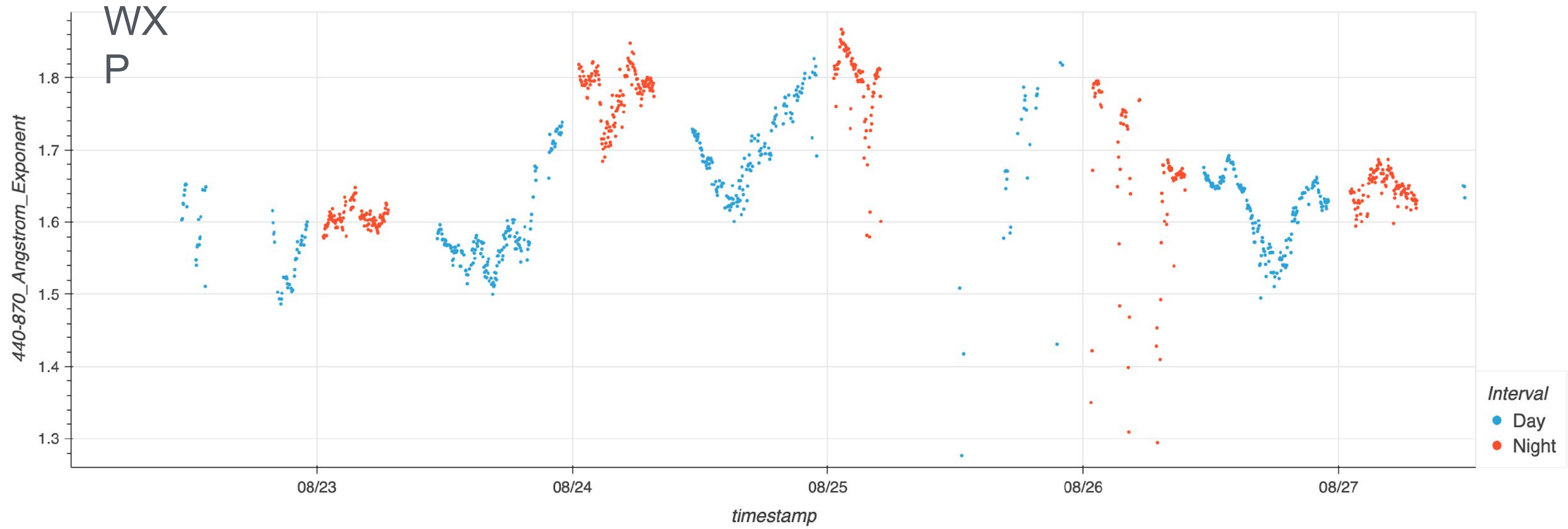
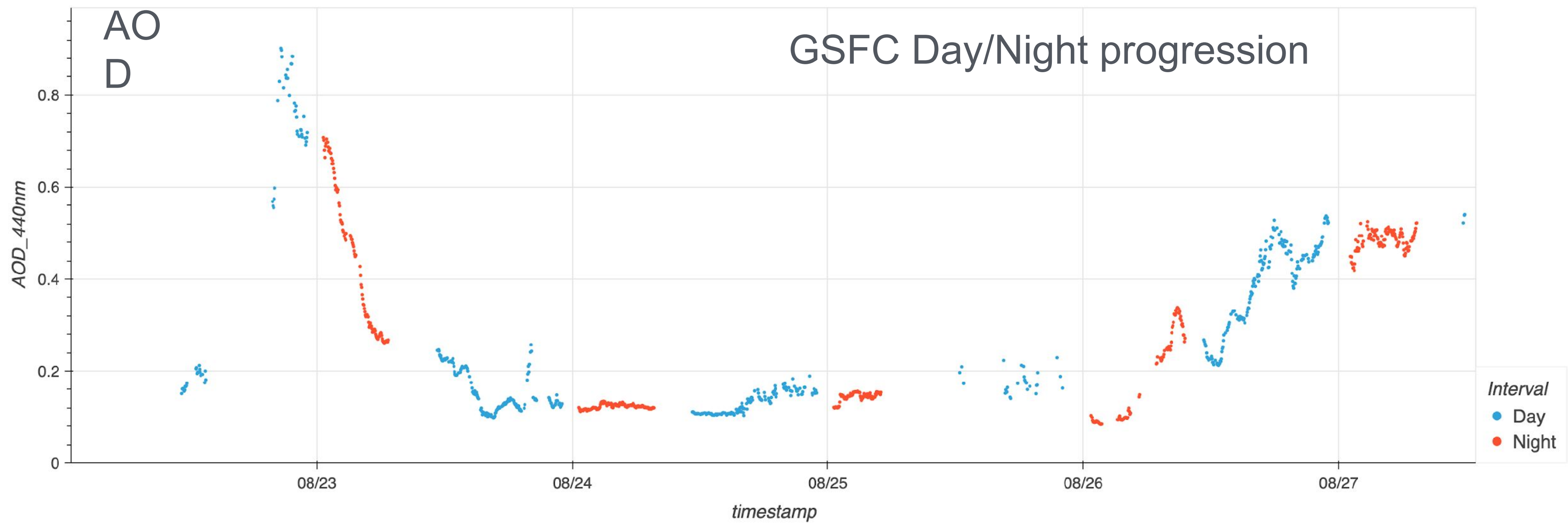
1640nm slightly different near full moon dip  
Combined sites data is more resistant to overfitting



Only 1640nm is notably different (RCF transects dip minimum)  
 1020/1020In show the same ~2.5% upward translation for LUT and RCF

## Comparing LUT and RCF Empirical Bias Correction

# **Assessment of Lunar Product**





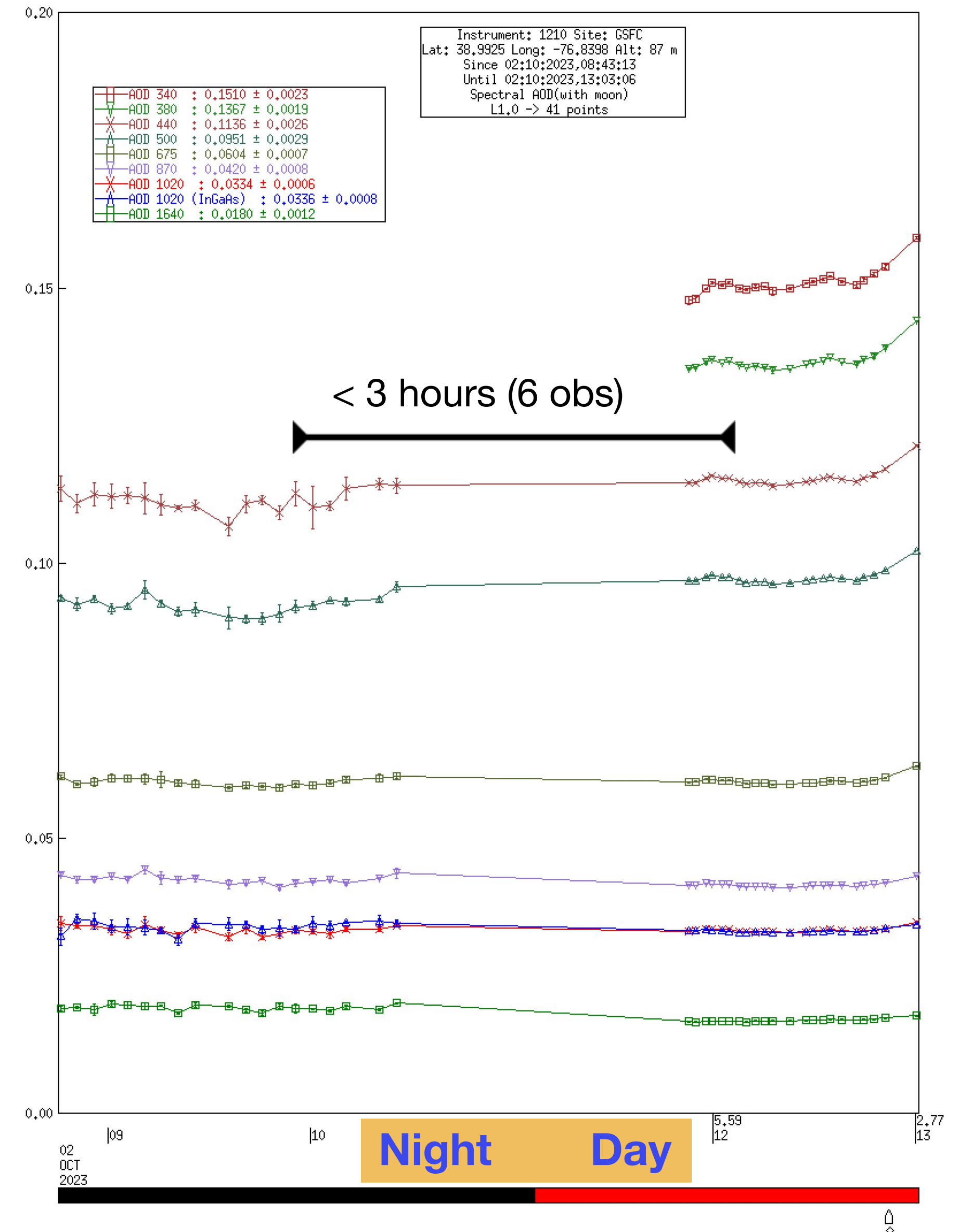
## Evaluation of full lunar database based on AOD continuity

Select all sites with Lunar AOD

N= 198 sites (> 5000 obs)

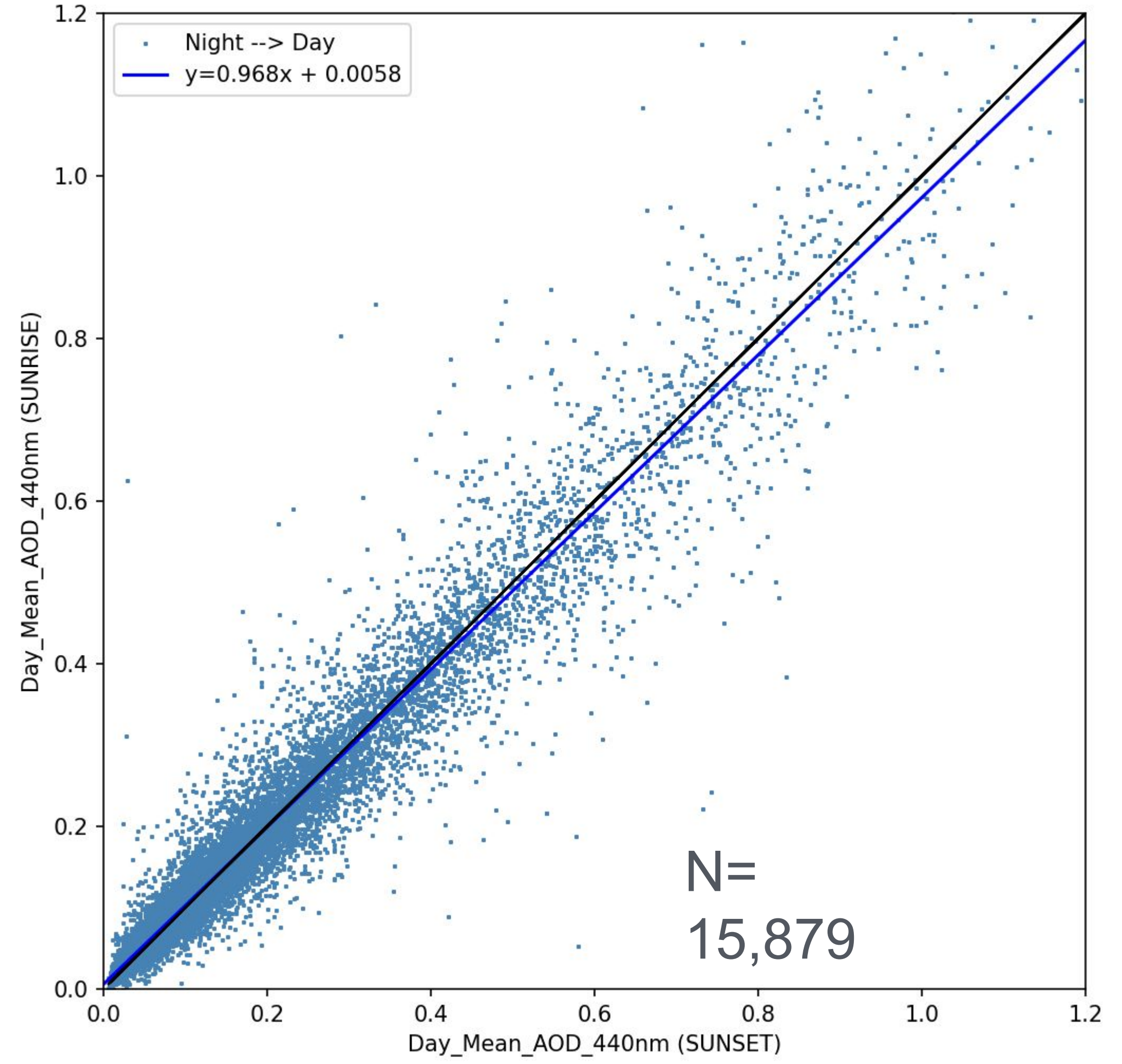
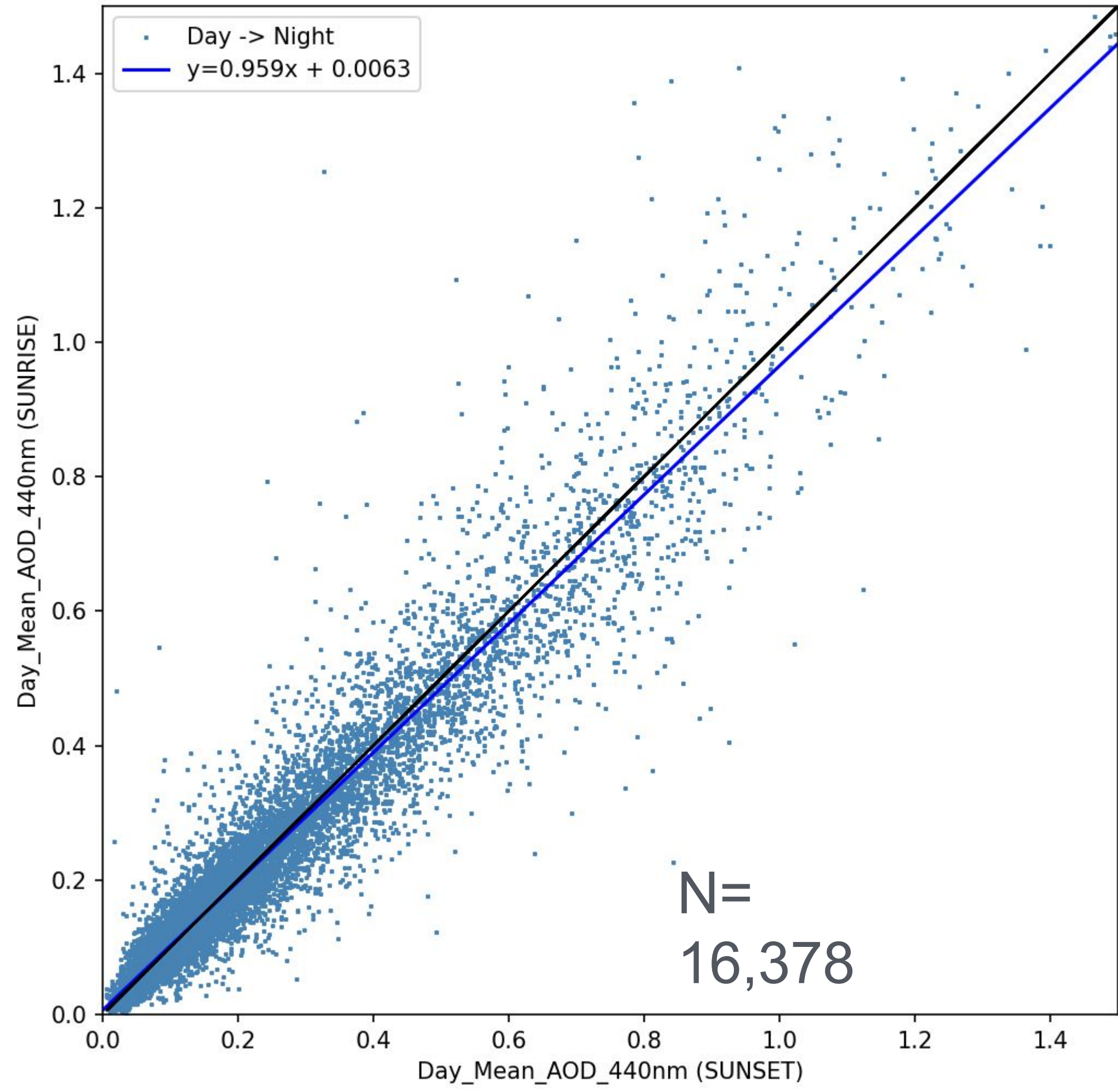
Lunar data record beginning around 2015

Average AOD from pre-/post-transition with total time span < 180 min



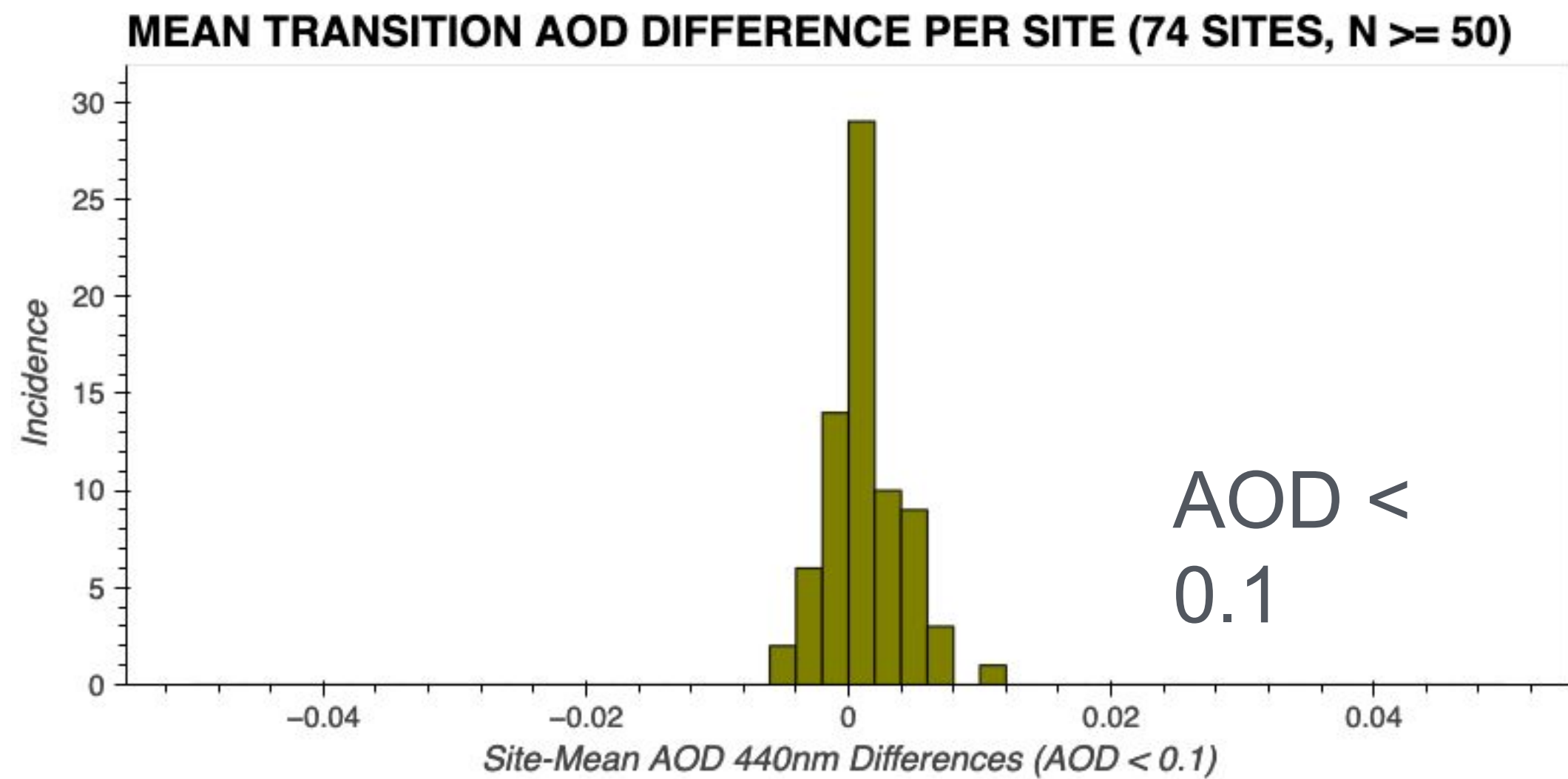
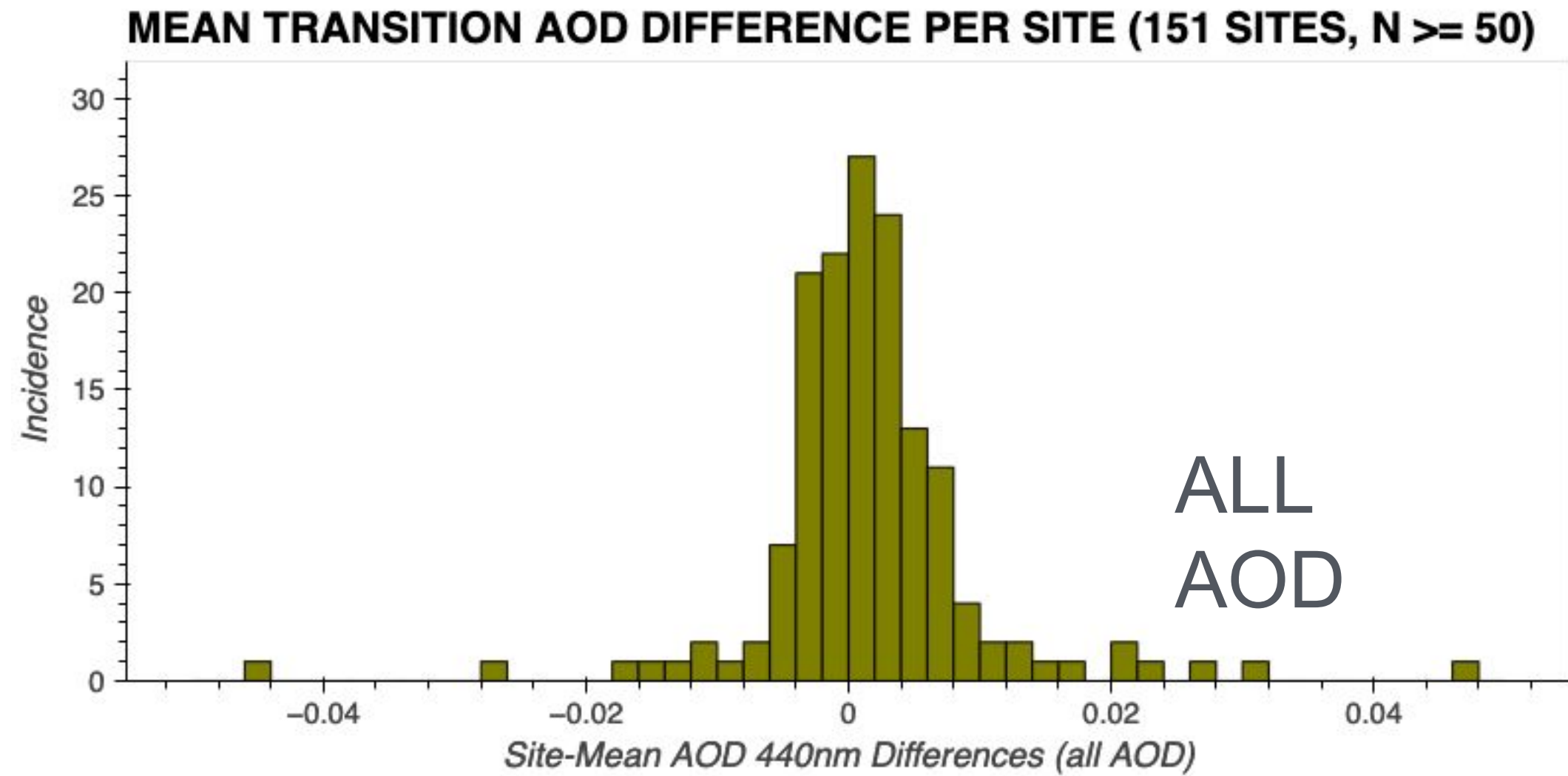
# All AOD Transition Interval Comparisons

AOD Differences at Sunrise/Sunset

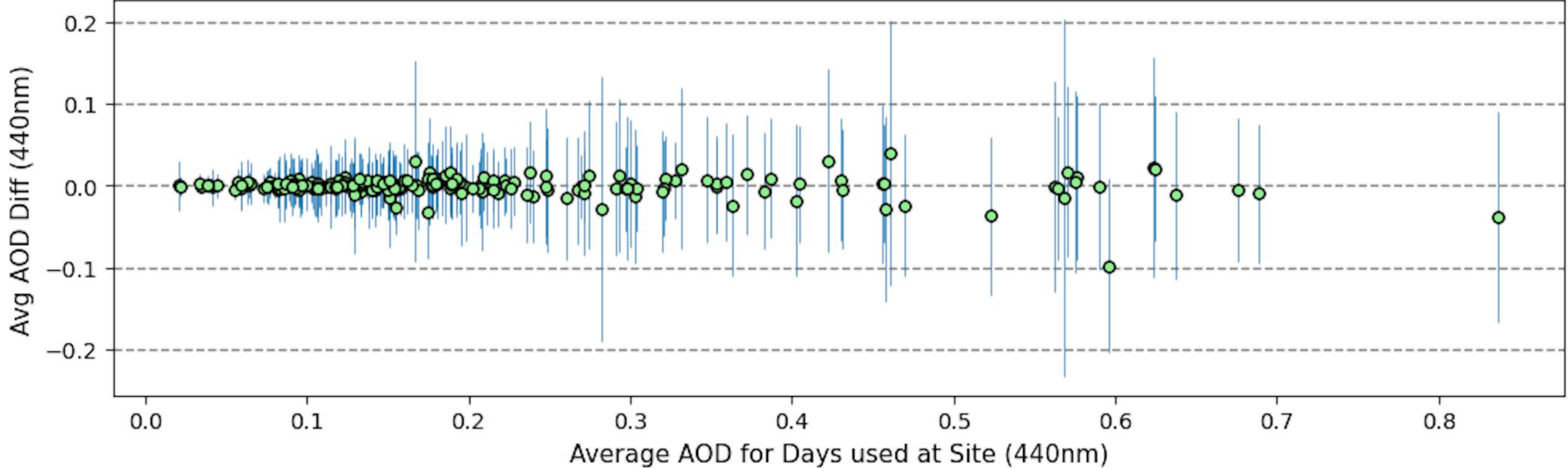


All site aggregated stats for transition interval AOD differences by channel

Channel	mean	std
440nm	0.0015	0.0154
500nm	0.0014	0.0142
675nm	0.0008	0.0133
870nm	0.0004	0.0122
1020nm	0.0003	0.0123
1640nm	-0.00003	0.0103

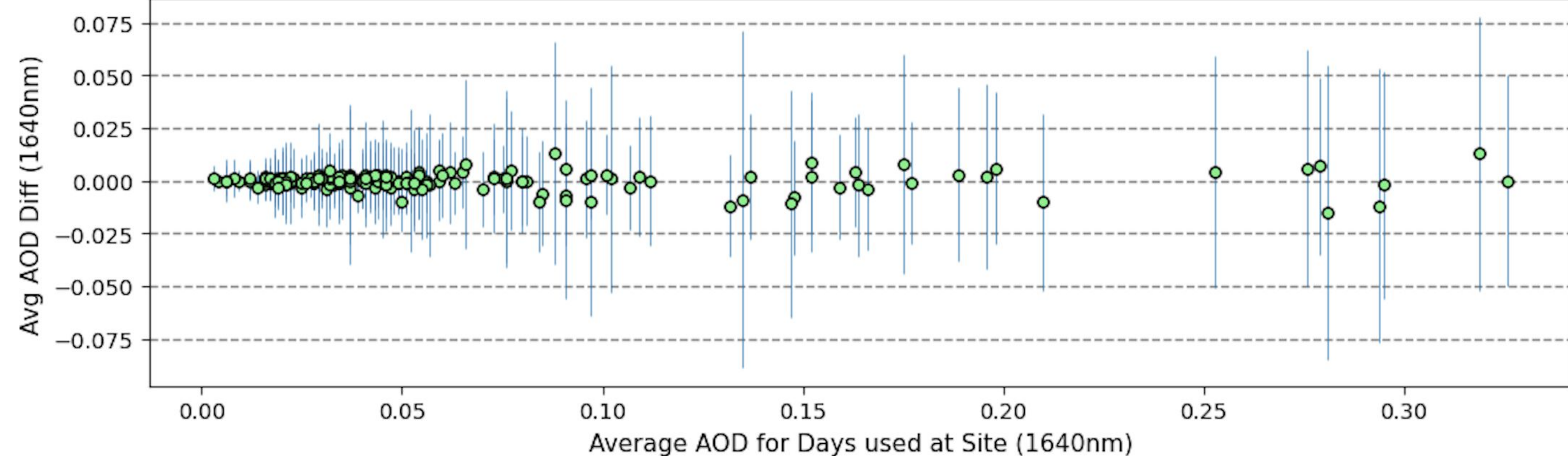


AOD DIFFERENCE IN TRANSITION INTERVAL (< 180 MINS) FOR ALL SITES (> 5K MOON OBS)



440nm

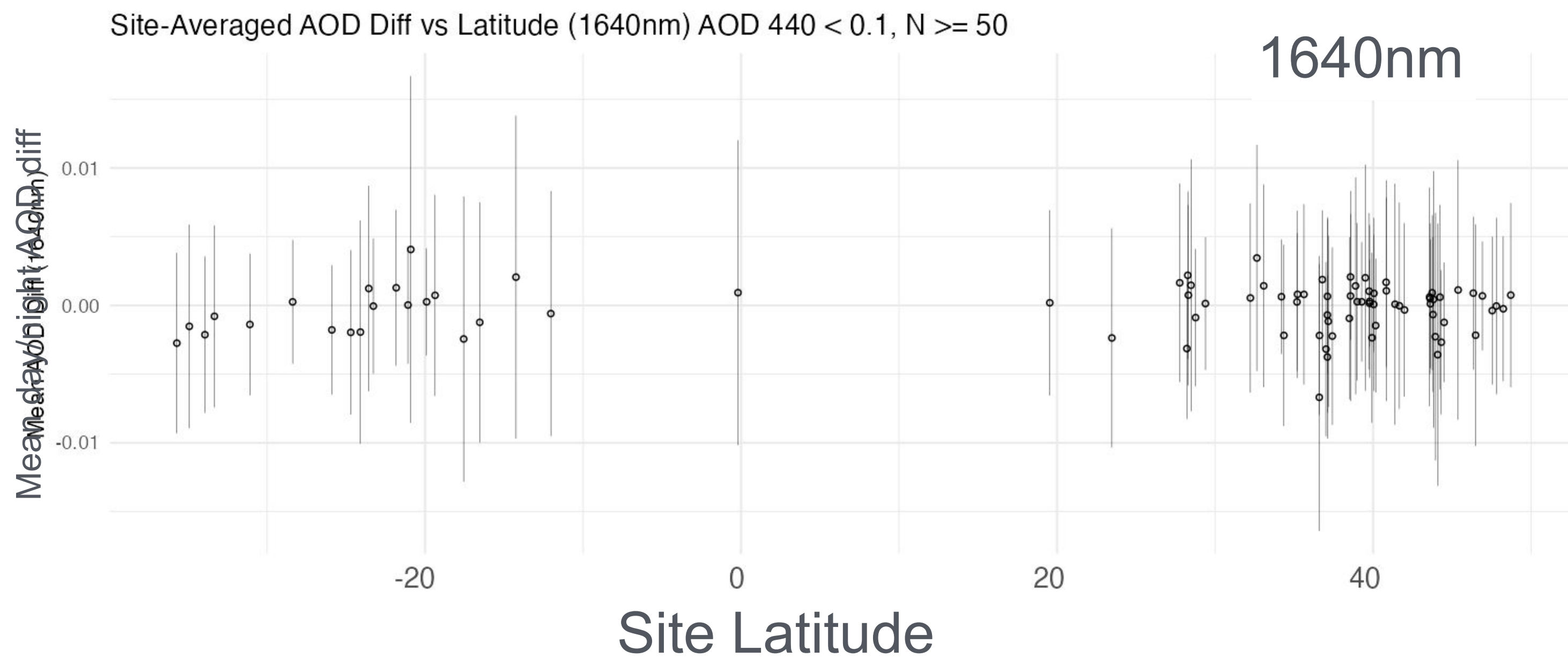
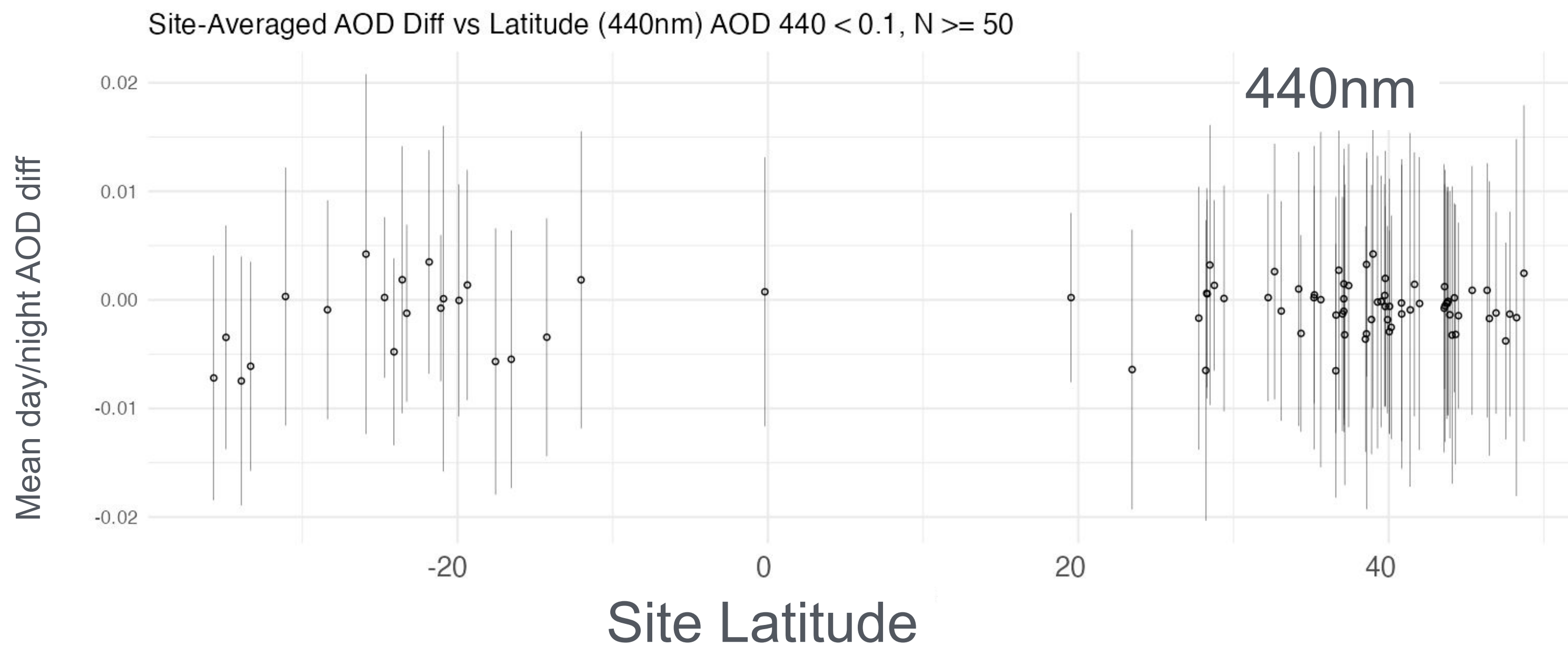
AOD DIFFERENCE IN TRANSITION INTERVAL (< 180 MINS) FOR ALL SITES (> 5K MOON OBS)



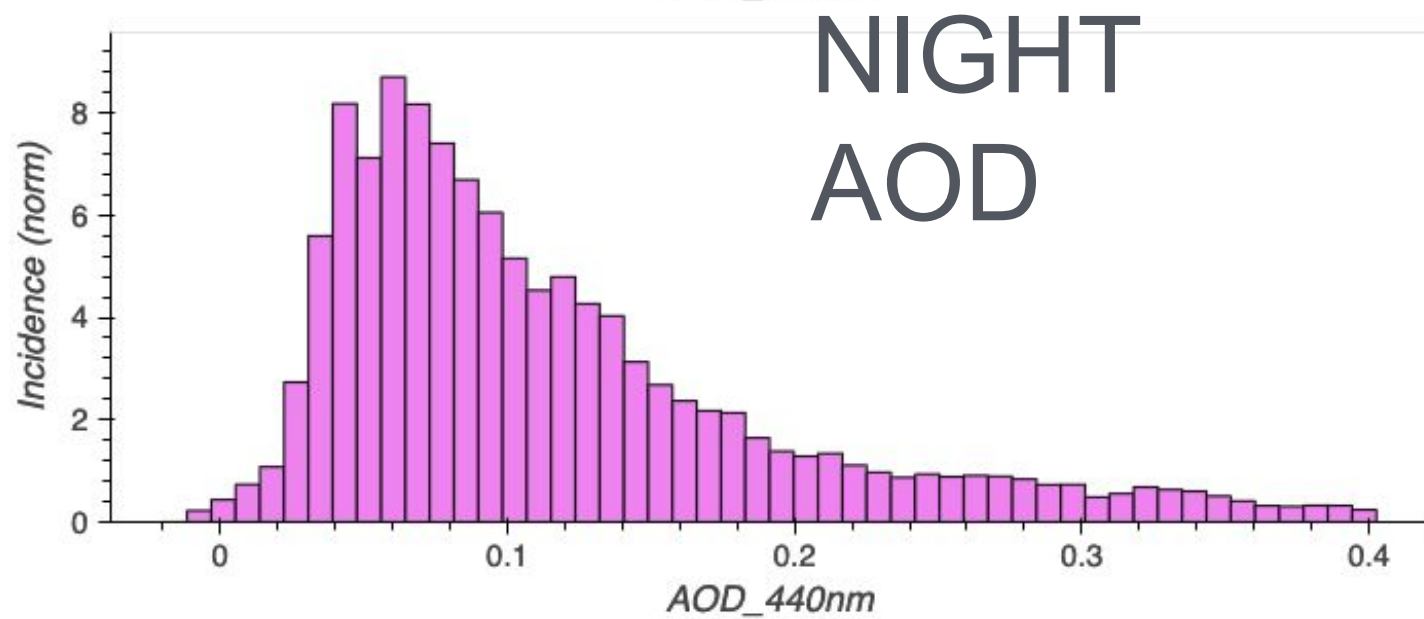
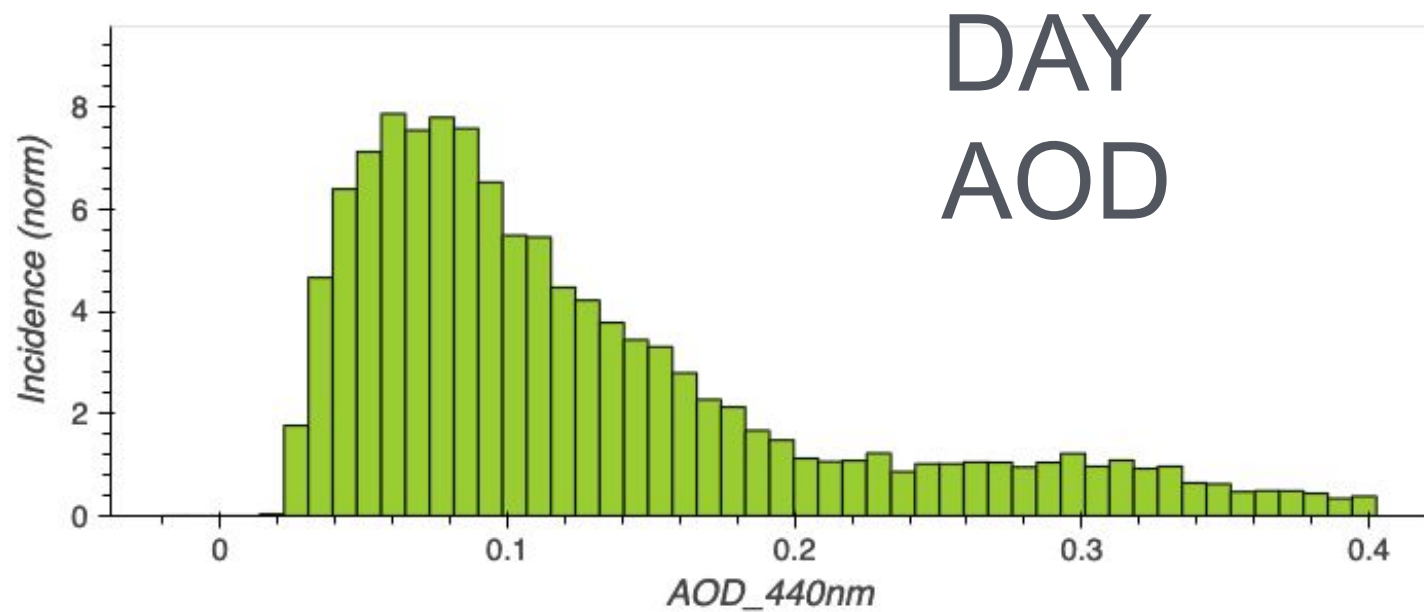
1640nm

# Check potential latitudinal dependence since bias defined at MLO/Izaña

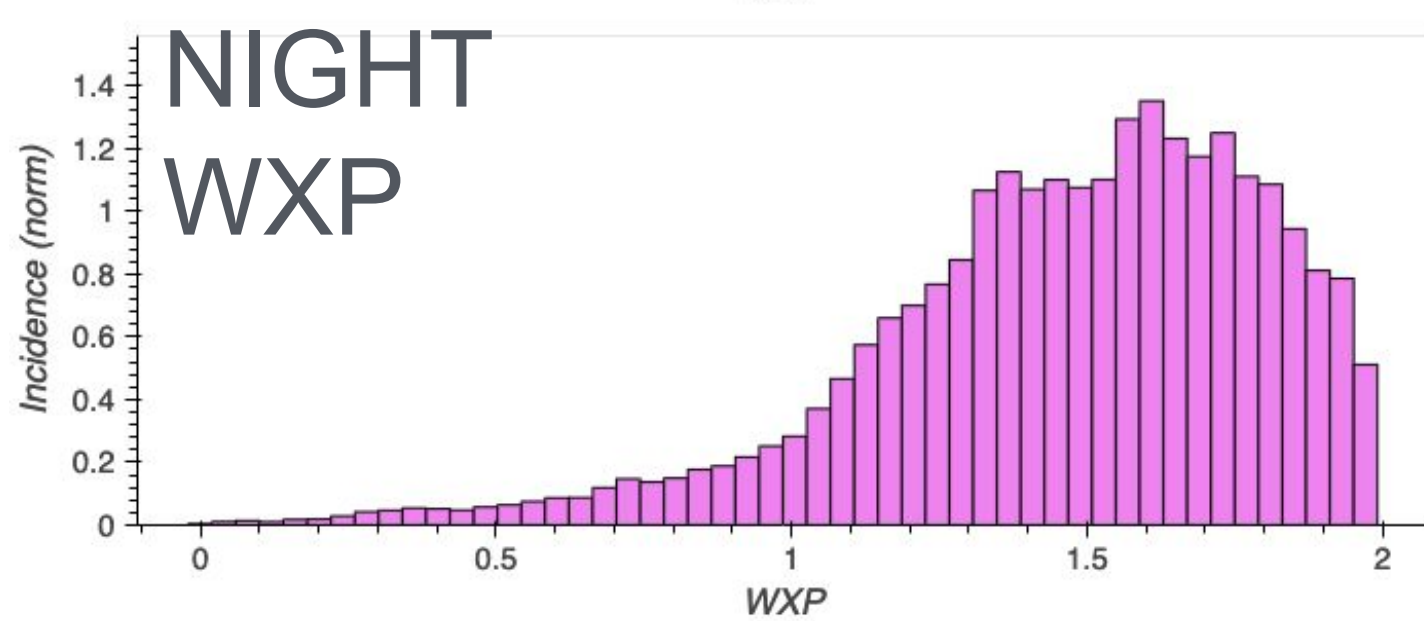
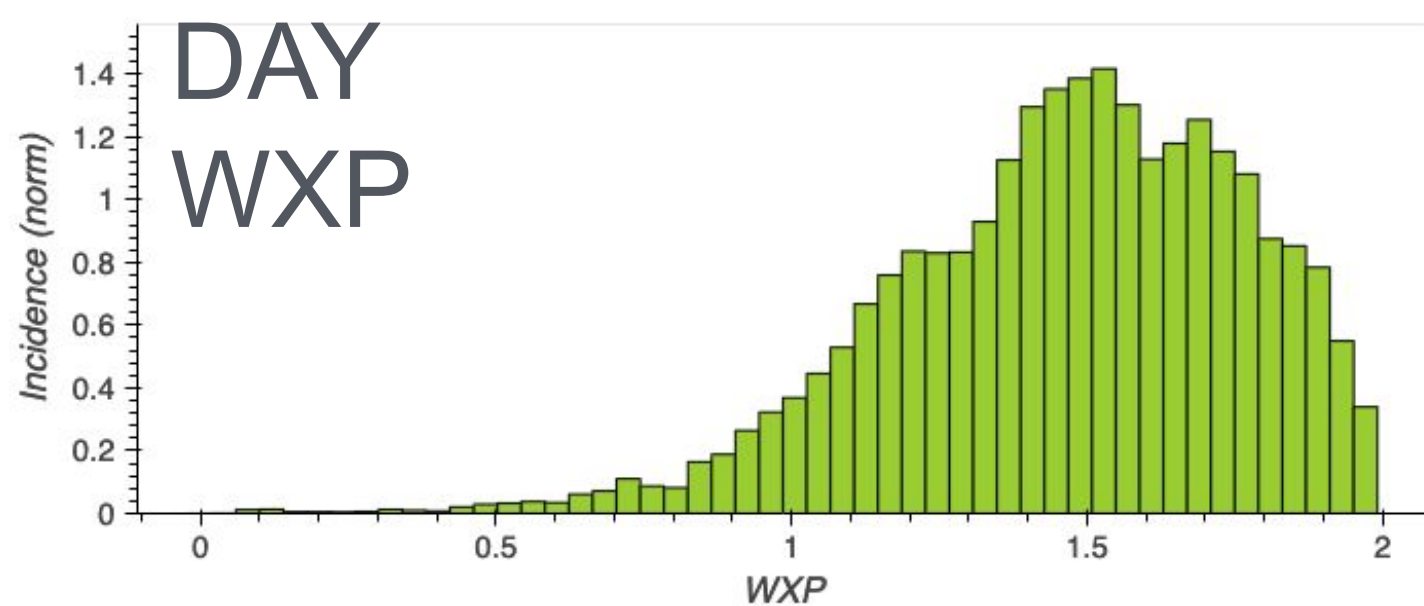
AOD Differences at Sunrise/Sunset



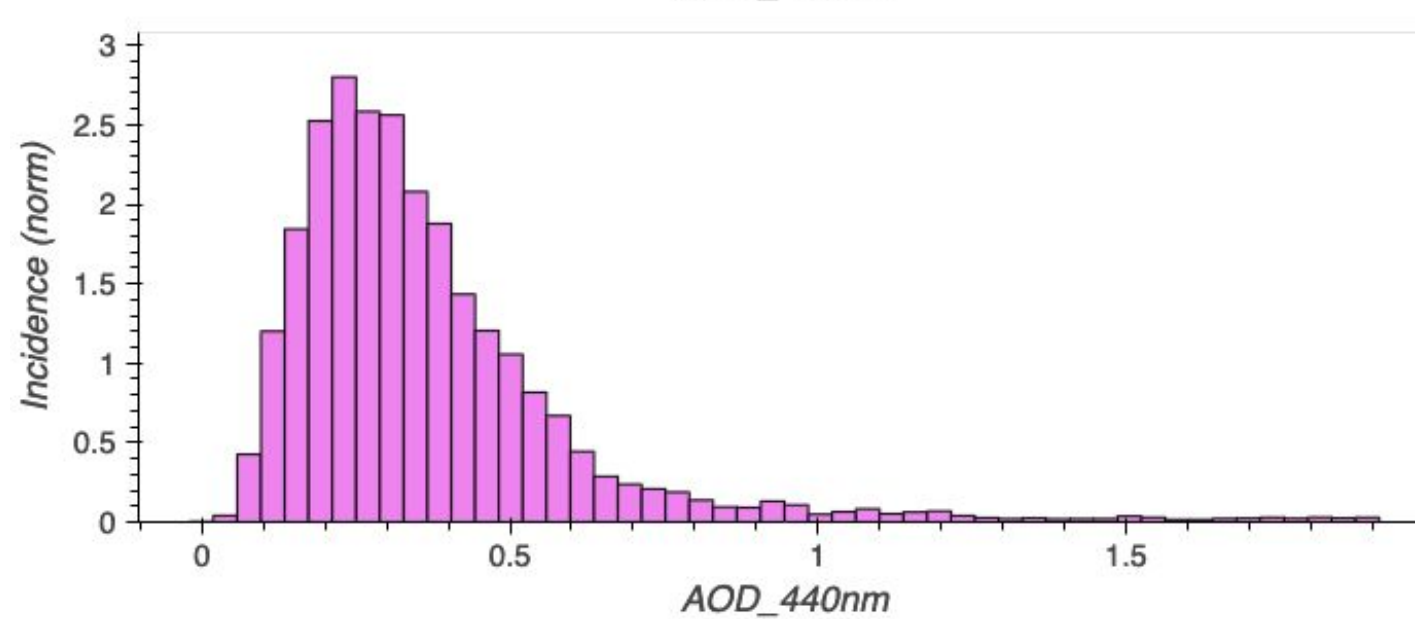
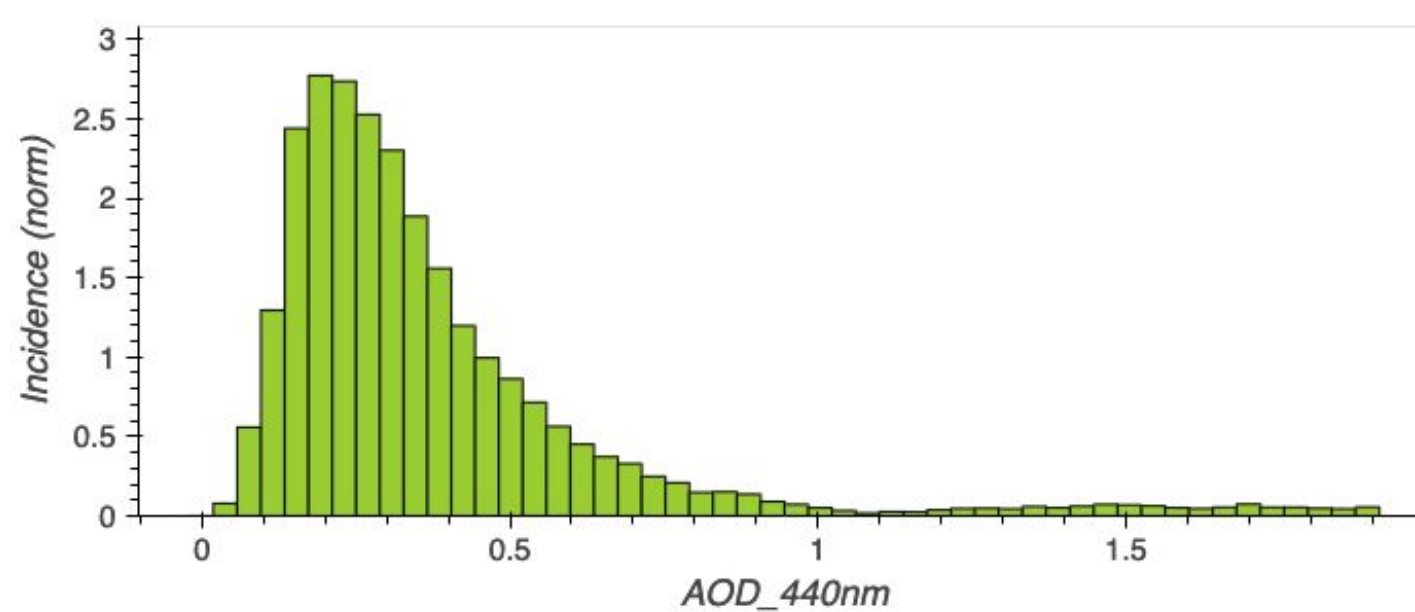
GSFC [Common Days with N > 10 Obs]



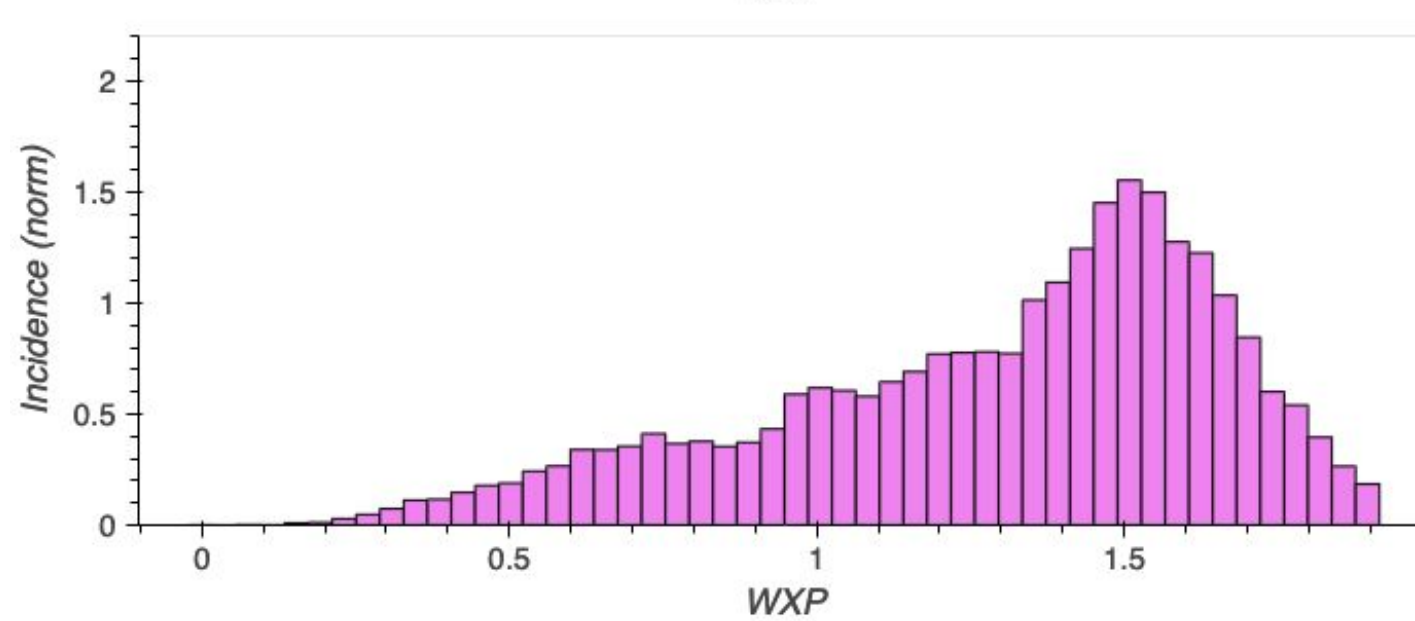
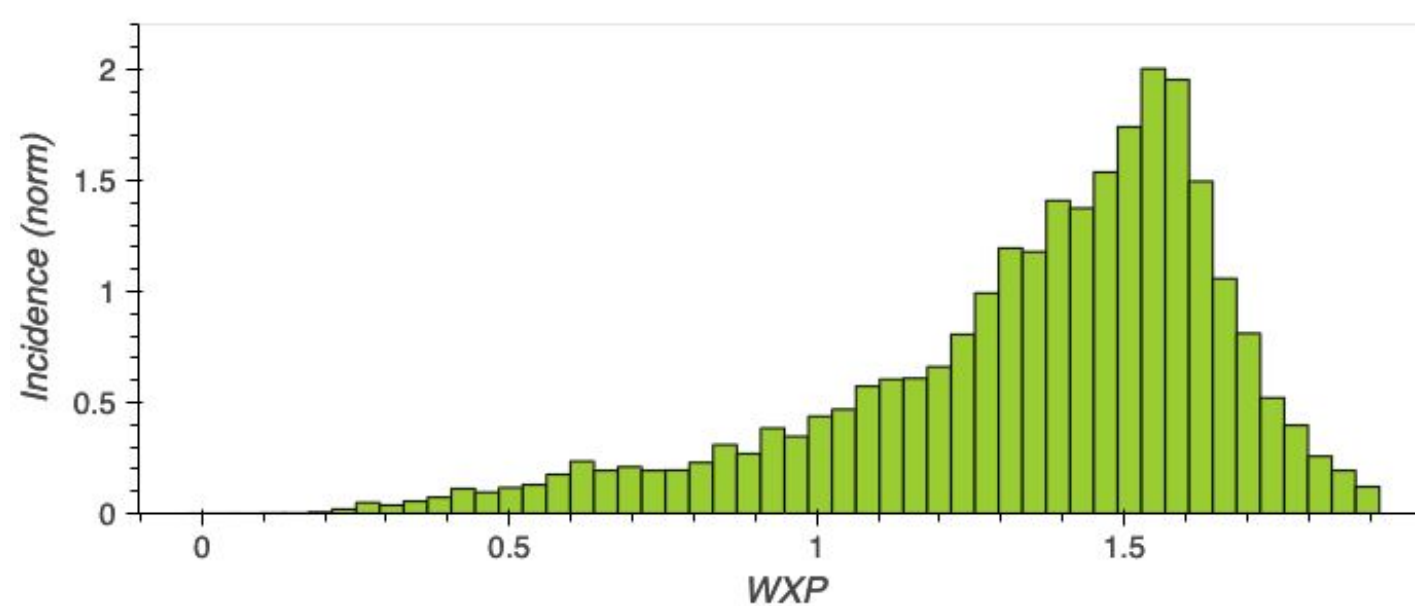
GSFC [Common Days with N > 10 Obs]



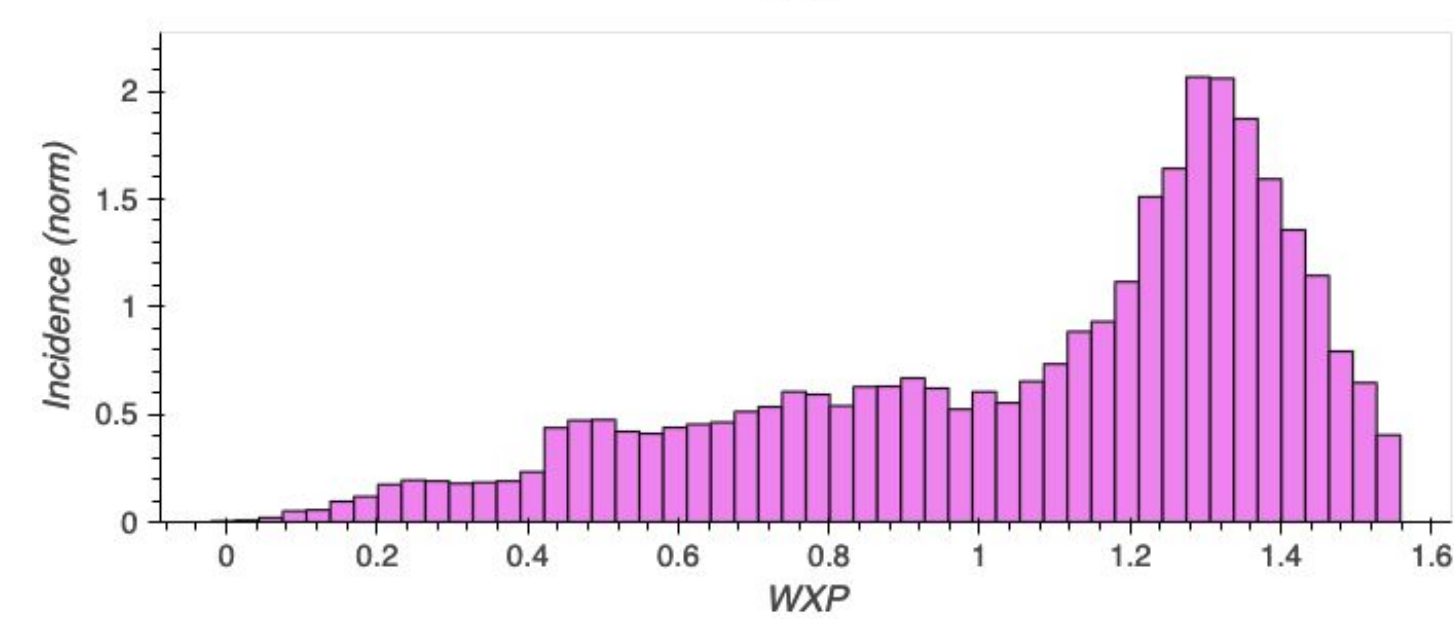
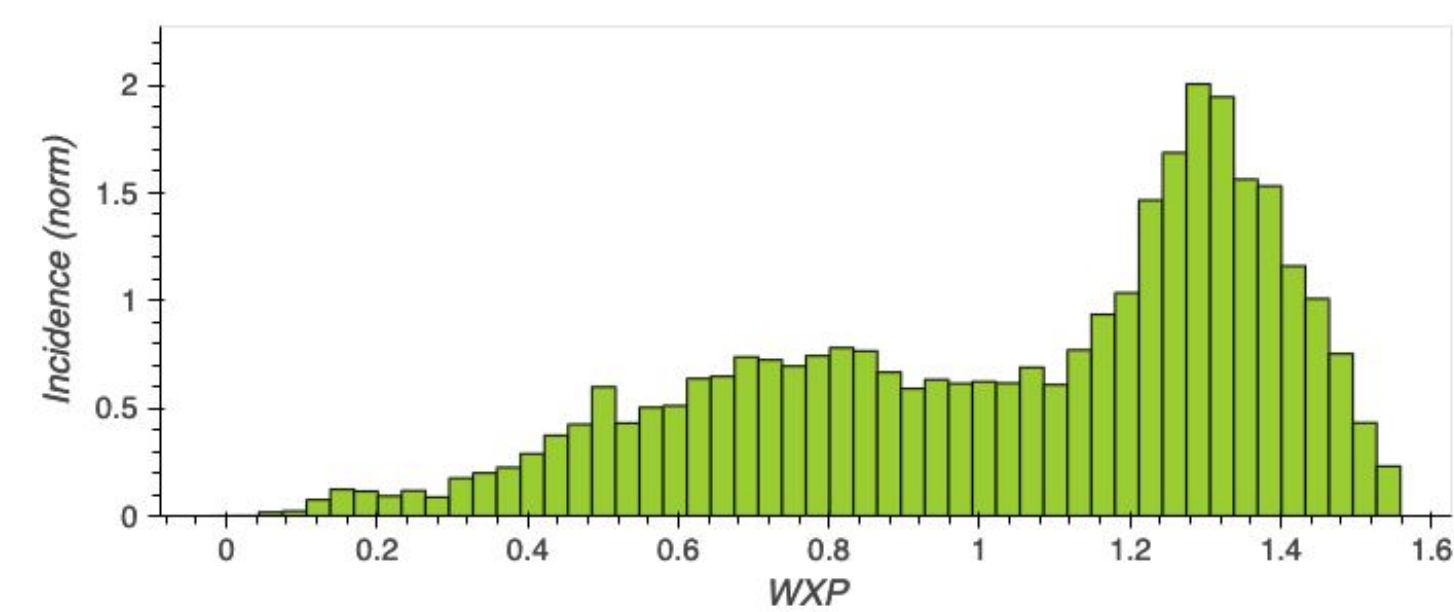
Singapore [Common Days with N > 10 Obs]



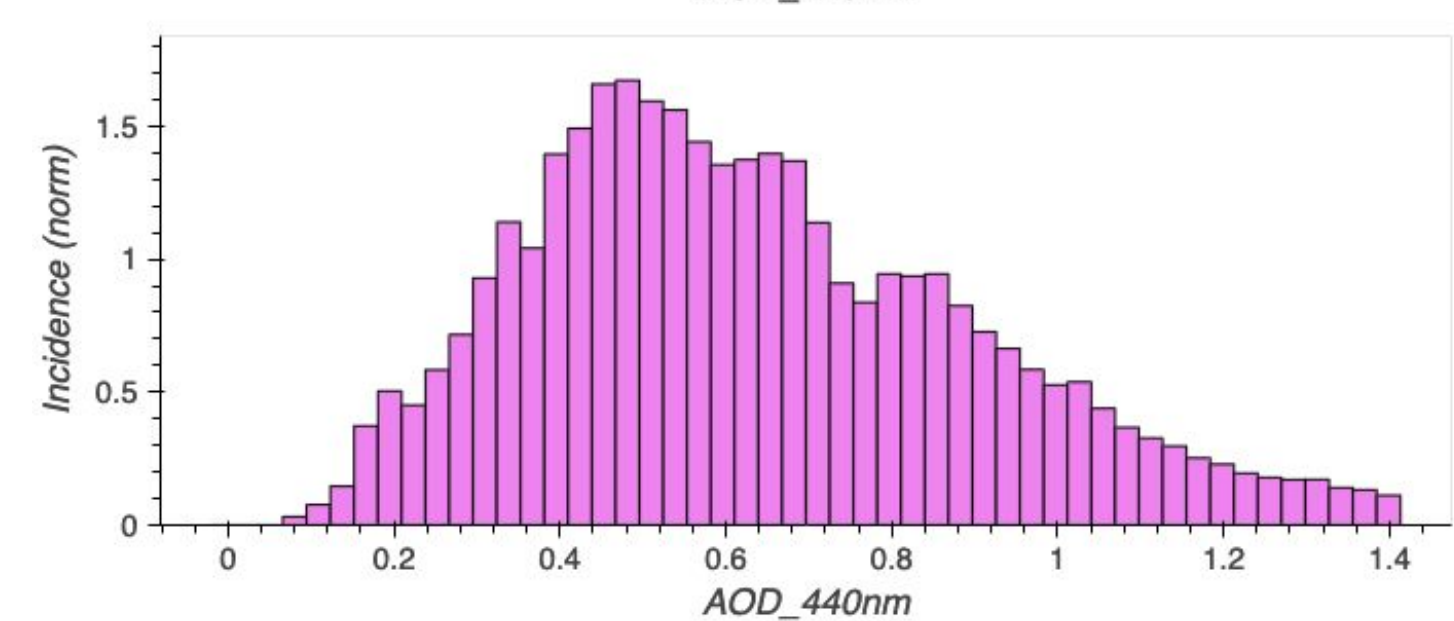
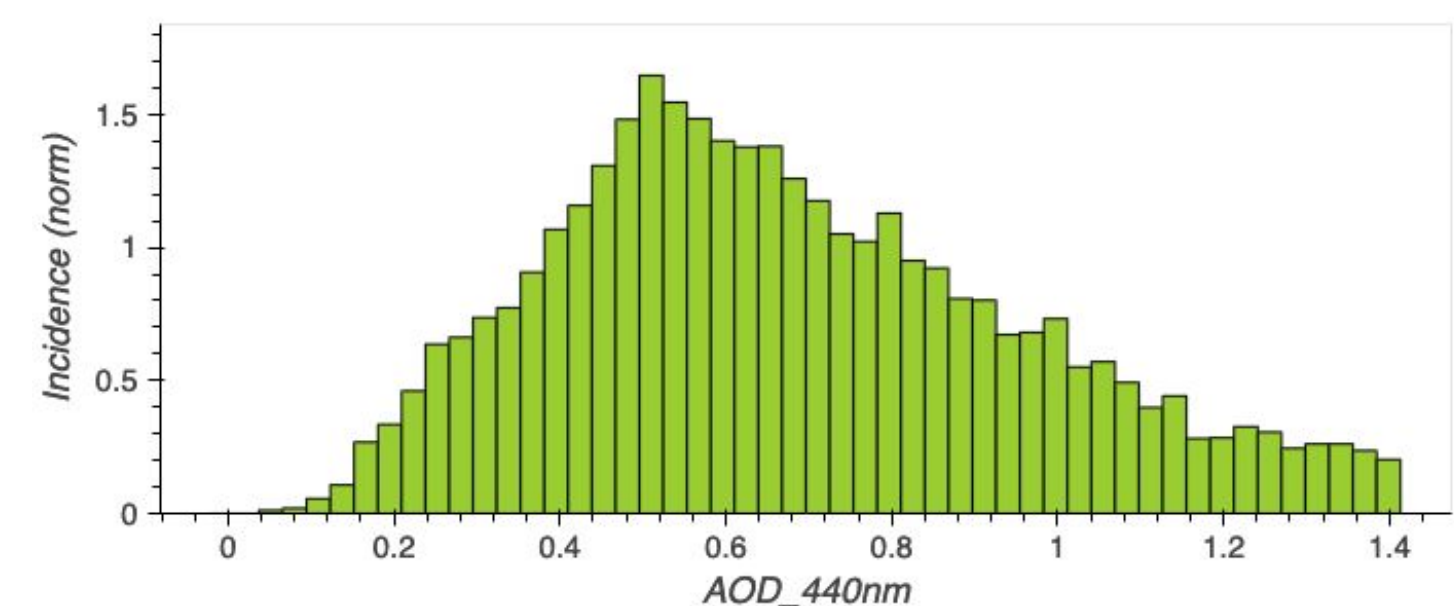
Singapore [Common Days with N > 10 Obs]



Kanpur [Common Days with N > 10 Obs]



Kanpur [Common Days with N > 10 Obs]



Thank you

AERONET Exchange  
2024

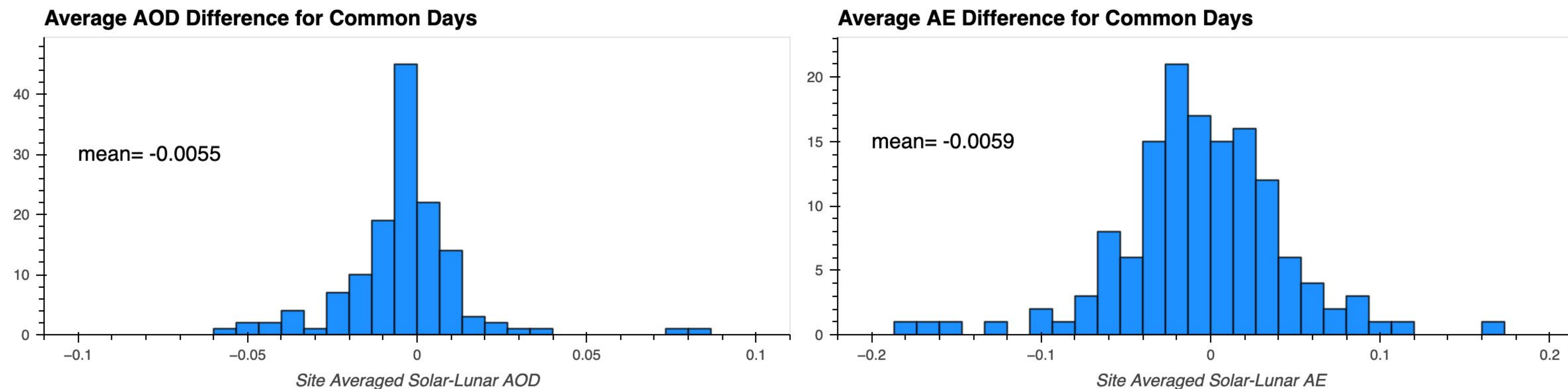
# PAIRED DAY AOD/AE DIFFERENCES PER SITE

AOD/AE averages computed for each day and only retained

for days with both solar and lunar averages ( $N \geq 3$ )

For each day, the difference in average AOD and AE was determined

Then, for each site the overall average of all daily differences was found



111 Sites. Average number of day pairs per site= 459